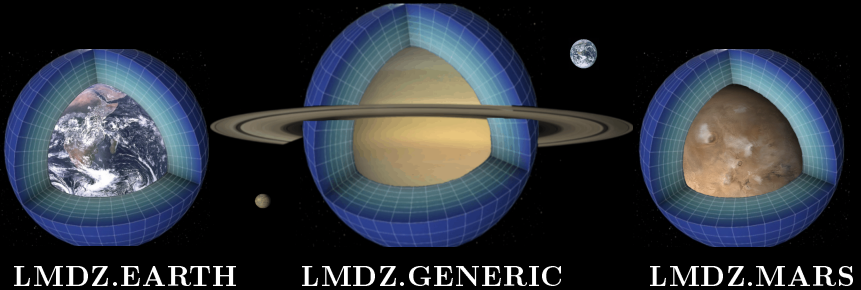


Energy spectra and dissipation

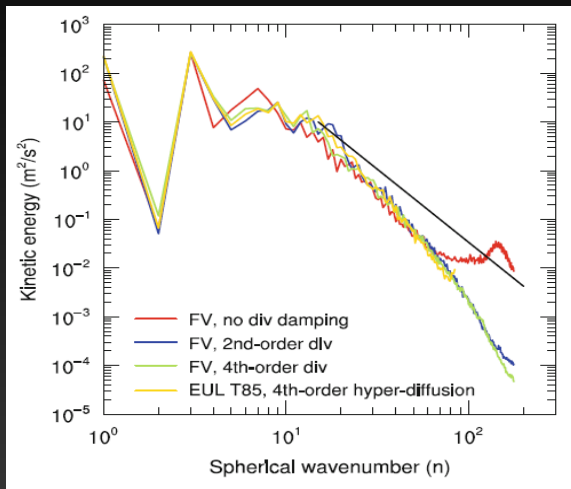


Mikel Indurain, Equipe Planéto
LMD

Dissipation : introduction

Energy transfer from large scales to small scales.

Problem : energy accumulation if dissipative structures are not resolved



Jablonowski,
2011

Dissipation : introduction

Energy transfer from large scales to small scales.

Problem : energy accumulation if dissipative structures are not resolved

⇒ dissipation operator : $F_\psi = (-1)^{q+1} K_\psi \nabla^{2q} \psi$

K_ψ : diffusion coefficient

$2q$: diffusion order

$$\frac{\partial \psi}{\partial t} = \text{Dyn}(\psi) + \text{Phy}(\psi) + F_\psi$$

Energy spectra

To study dissipation effect, we compute the spectra of specific kinetic energy :

$$KE = \frac{1}{2} \vec{V} \cdot \vec{V}$$

For $1 \leq n \leq N$, spatial mean of specific kinetic energy :

$$\overline{KE}_n = \xi_n^0 (\xi_n^0)^* + \delta_n^0 (\delta_n^0)^* + 2 \sum_{m=1}^n [\xi_n^m (\xi_n^m)^* + \delta_n^m (\delta_n^m)^*]$$

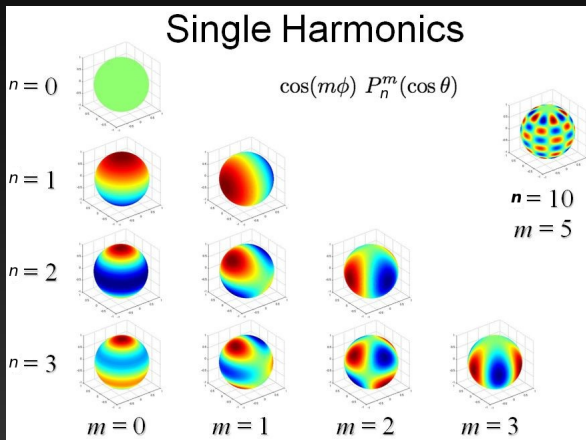
where ξ_n^m : spectral coefficient of vorticity

δ_n^m : spectral coefficient divergence

Spherical harmonics

Spherical harmonics form a basis for functions on the sphere :

$$f(\theta, \phi) = \sum_{n=0}^N \sum_{m=-n}^n c_n^m Y_n^m(\theta, \phi)$$



Truncation at :

$$N = \max(N_{lat}, \frac{N_{long}}{2})$$

Dissipation in LMDZ.EARTH

$$F_\psi = (-1)^{q+1} K_\psi \nabla^{2q} \psi$$

Dissipation operator is applied to :

horizontal velocity

potential temperature

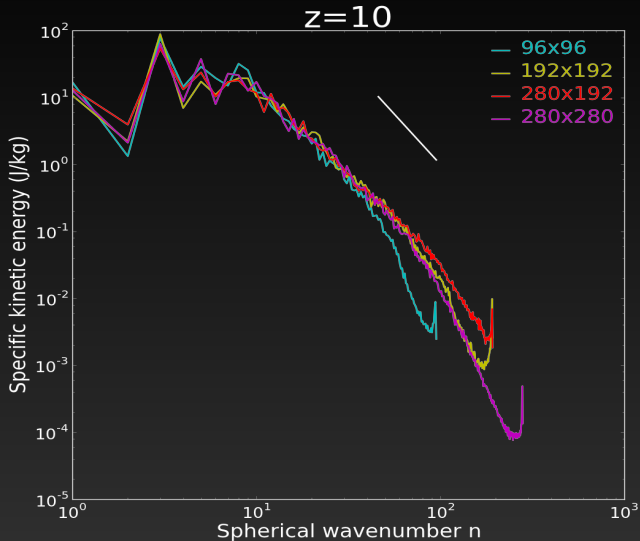
Diffusion coefficient $K_\psi \propto \frac{1}{\text{tetagdiv}}$ or $\propto \frac{1}{\text{tetagrot}}$ or $\propto \frac{1}{\text{tetatemp}}$

Vertical variation of K_ψ with `vert_prof_dissip = True`

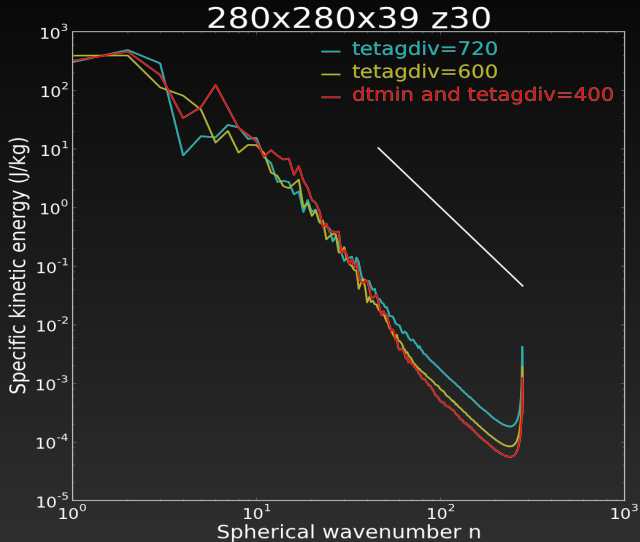
Dissipation coefficients increase with altitude

Diffusion order $2q$: `nitergdiv (=1)` or `nitergrot (=2)` or `niterh (=2)`

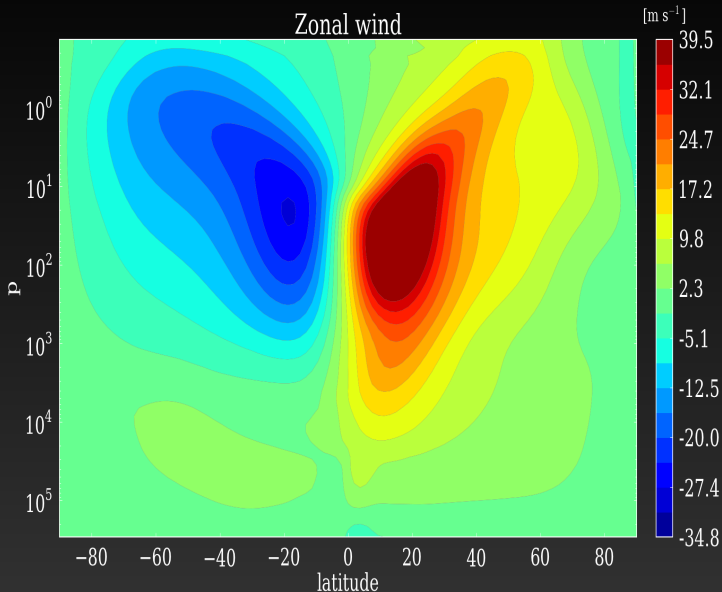
Energy spectra : Earth



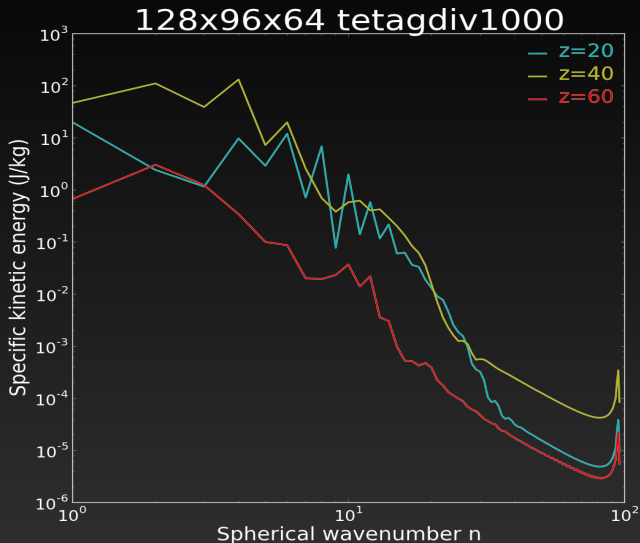
Energy spectra : Earth



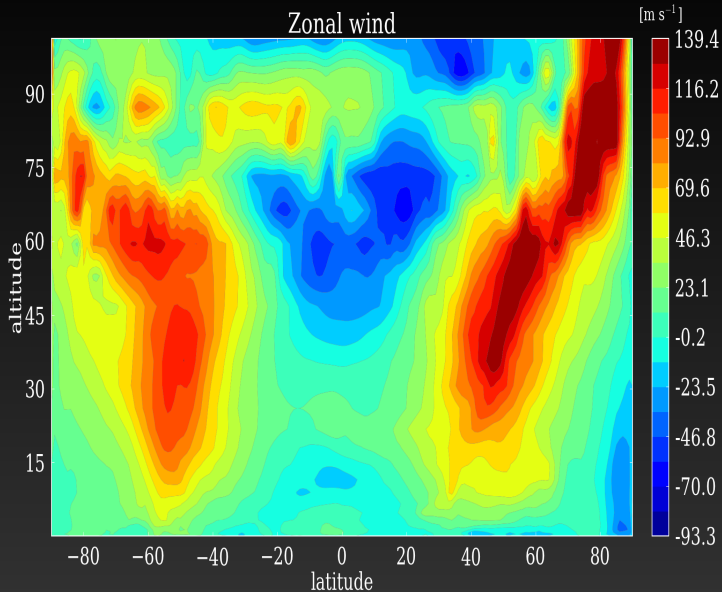
Energy spectra : Saturn's stratosphere



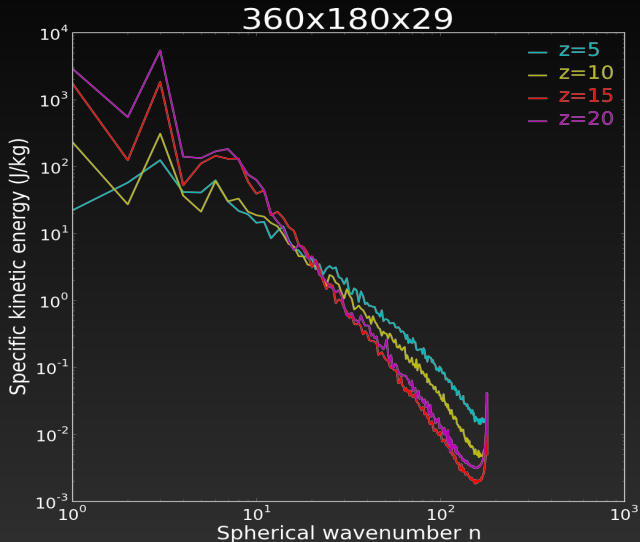
Energy spectra : Saturn's stratosphere



Energy spectra : Mars

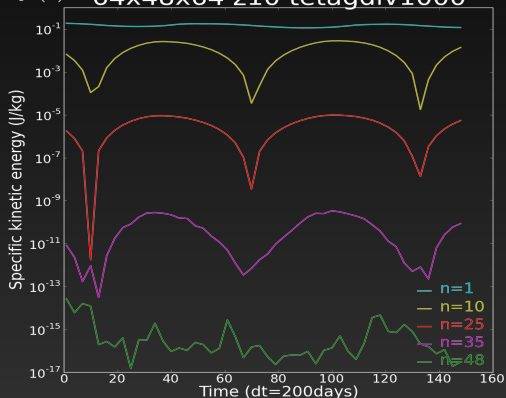


Energy spectra : Mars



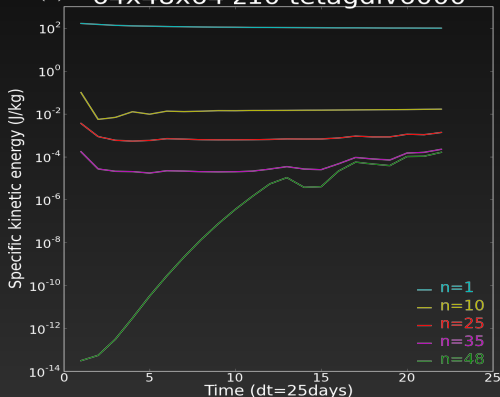
Questions

- ▶ Number of truncation for harmonic basis
- ▶ Higher dissipation order (now : 1 or 2)
- ▶ Dissipation coefficient as a function of resolution
- ▶ $\overline{KE}_n = f(t)$ 64x48x64 z10 tetagdiv1000



Questions

- ▶ Number of truncation for harmonic basis
- ▶ Higher dissipation order (now : 1 or 2)
- ▶ Dissipation coefficient as a function of resolution
- ▶ $\overline{KE}_n = f(t)$ 64x48x64 z10 tetagdiv6000



Thanks for your attention!

