

# Clouds

LMDz Training – December 2019

J-B Madeleine, C. Rio and the LMDz team

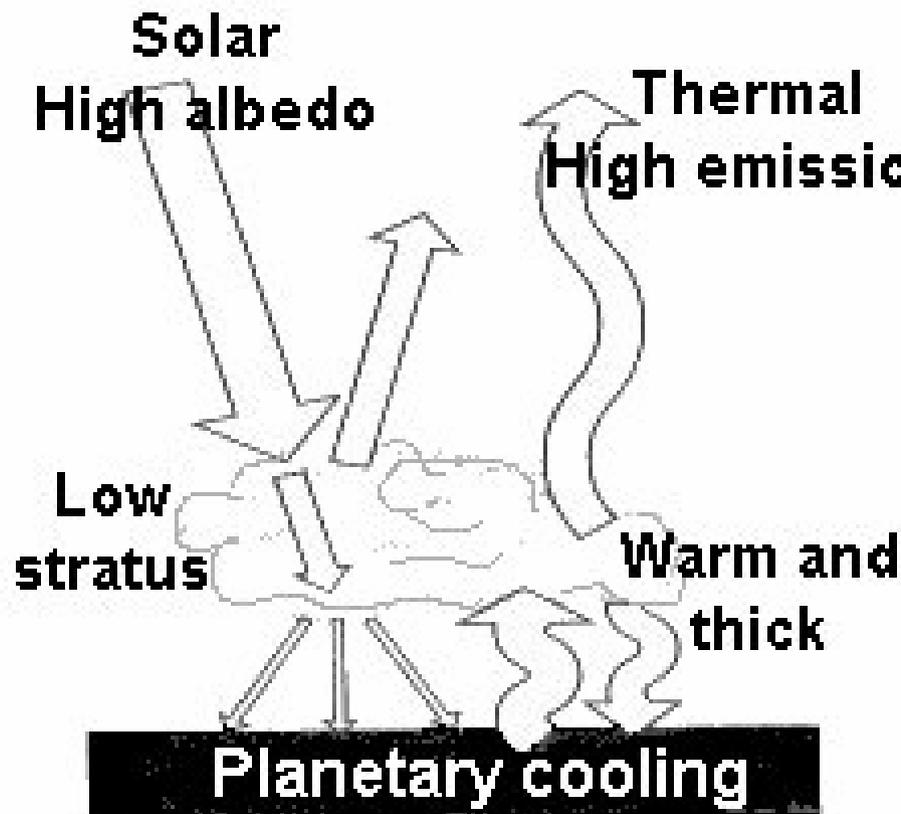
*wget <ftp://ftp.lmd.jussieu.fr/pub/jbmlmd/transfer/config.def>*



Picture by Oleg Artemyev taken from the ISS

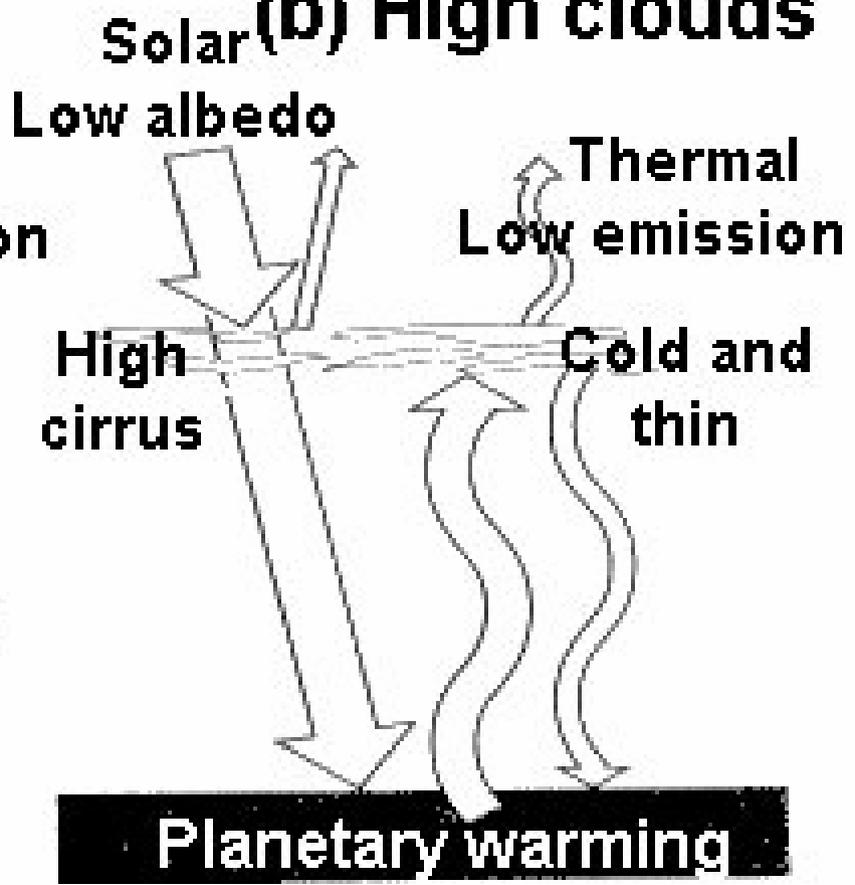
# Radiative impact of clouds

## (a) Low clouds



- Low clouds
- albedo effect (reflectivity of 40-50%)
  - weak greenhouse effect (high temp)

## (b) High clouds



- High clouds :
- weak albedo effect
  - strong greenhouse effect (cold clouds)

# Radiative forcing

## LW radiative forcing

**Positive** : clouds reduce the LW outgoing radiation

Annual mean :  $+29 \text{ W m}^{-2}$

## SW radiative forcing

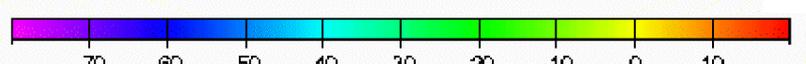
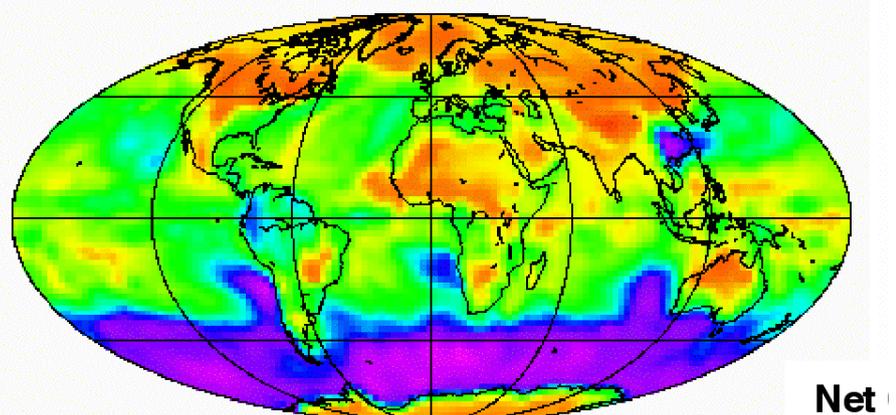
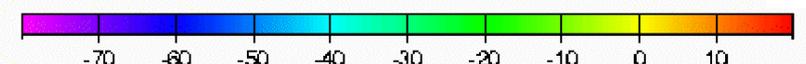
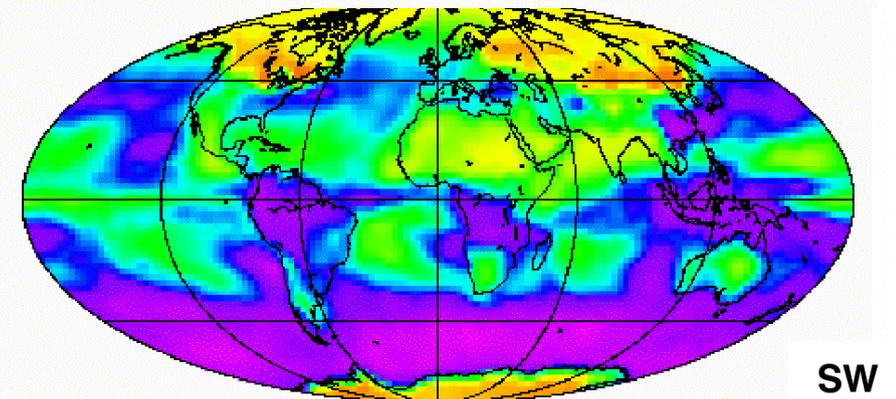
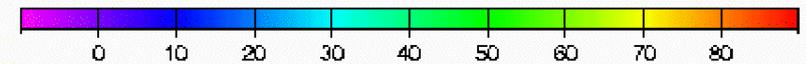
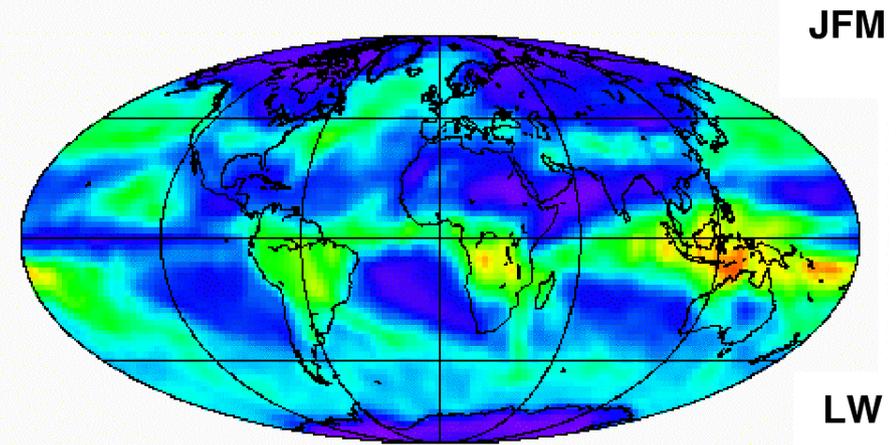
**Negative** : clouds reflect the incoming SW radiation

Annual mean :  $-47 \text{ W m}^{-2}$

Net forcing : **Cooling**

Annual mean :  $-18 \text{ W m}^{-2}$

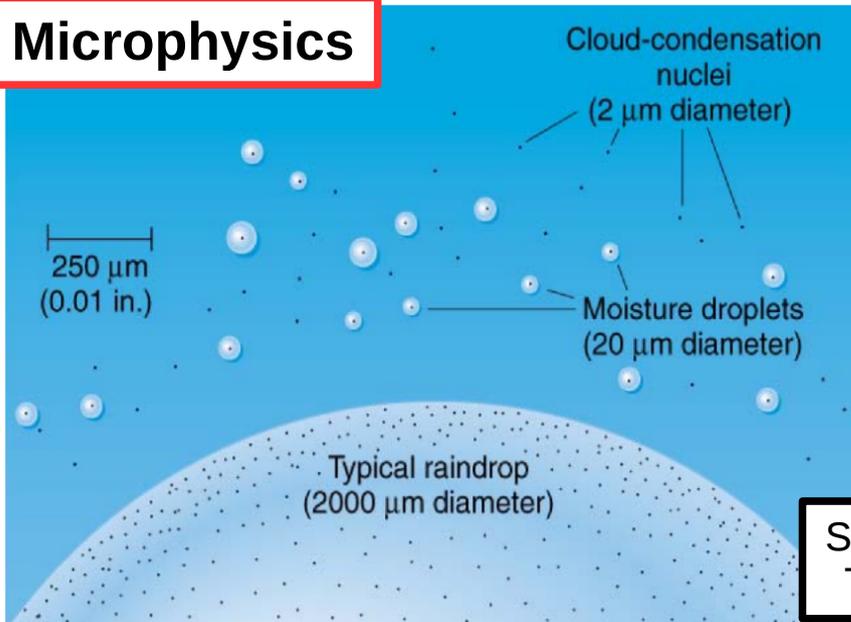
« The single largest uncertainty in determining the climate sensitivity to either natural or anthropogenic changes are clouds and their effects on radiation » IPCC report



Net

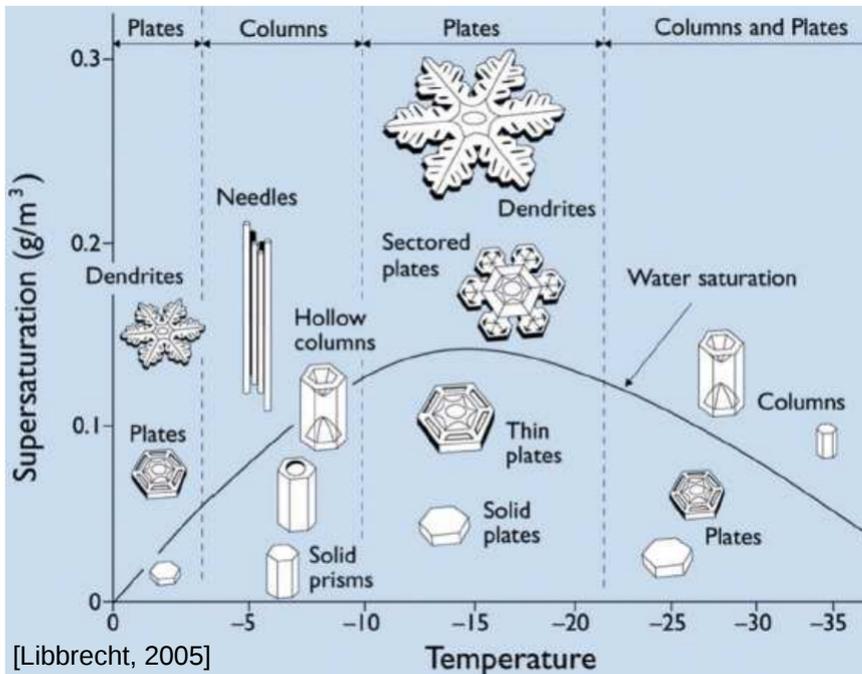
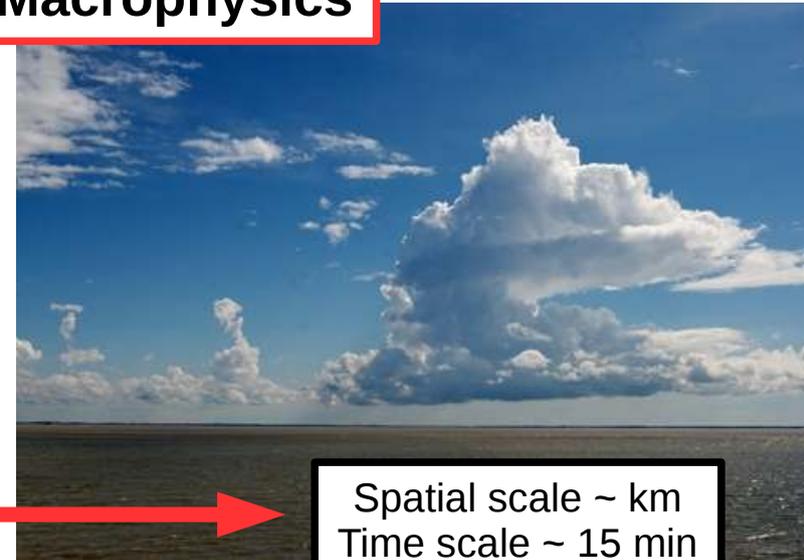
# Modeling clouds : a challenge

## Microphysics



Spatial scale  $\sim \mu\text{m}$   
Time scale  $\sim 1 \text{ s}$

## Macrophysics



# Fundamental process

- Clausius-Clapeyron equation :

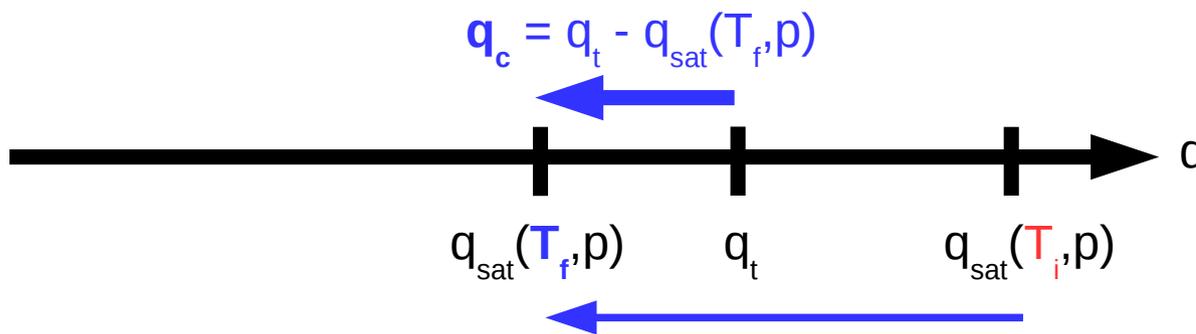
$$\frac{1}{e_{\text{sat}}} \frac{de_{\text{sat}}}{dT} = \frac{L}{R_{\text{vap}} T^2}$$

T	0°C	20°C
$e_{\text{sat}}$	6.1 hPa	23.4 hPa
$q_{\text{sat}}$	3.7 g kg <sup>-1</sup>	14.4 g kg <sup>-1</sup>

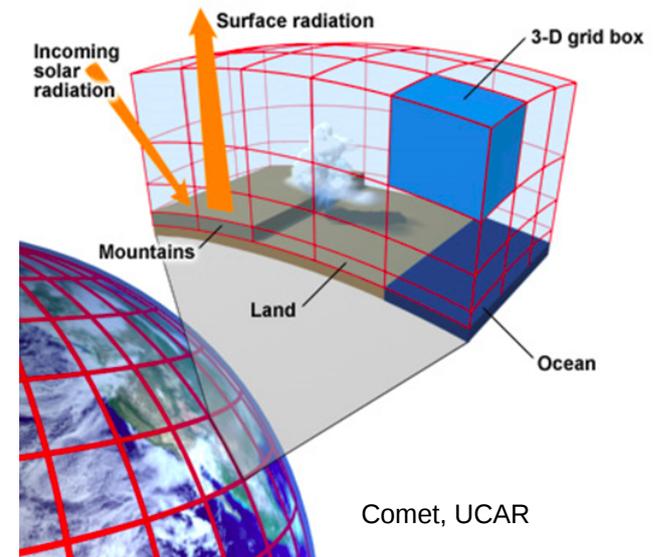
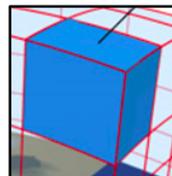
- Saturation mass mixing ratio :

$$q_{\text{sat}}(T, p) \simeq 0.622 \frac{e_{\text{sat}}(T)}{p}, \text{ where } e_{\text{sat}}(T) \text{ grows exponentially with temperature}$$

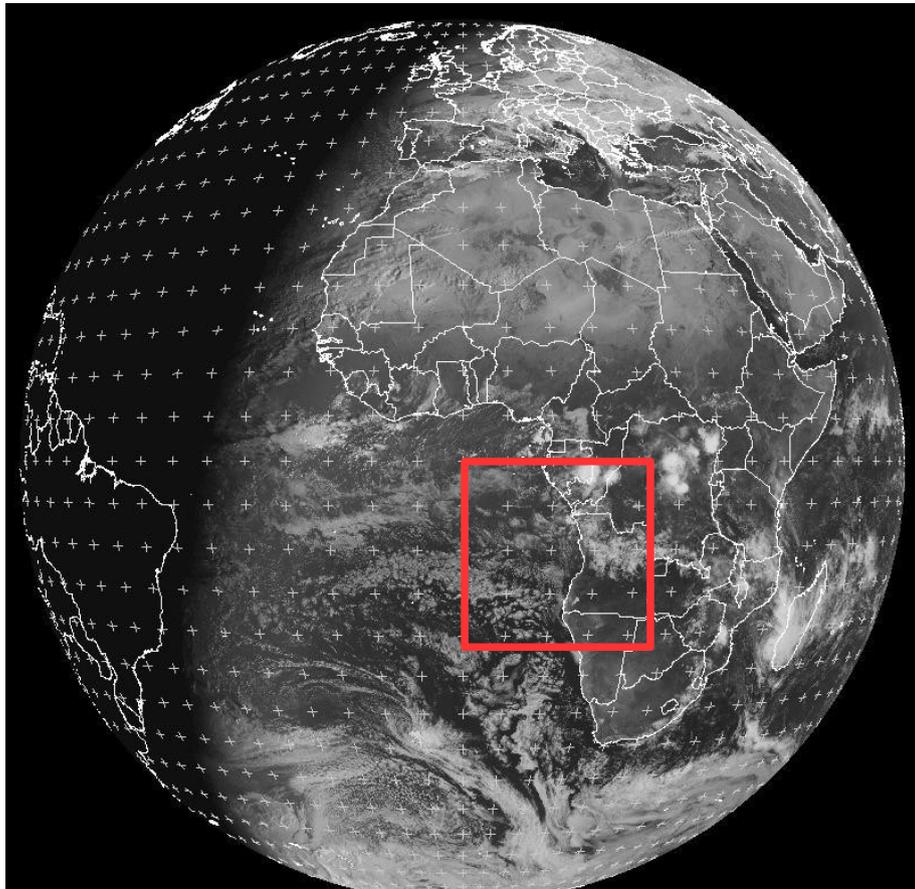
- Clouds form when an air parcel is cooled :



- But clouds do not look like that :

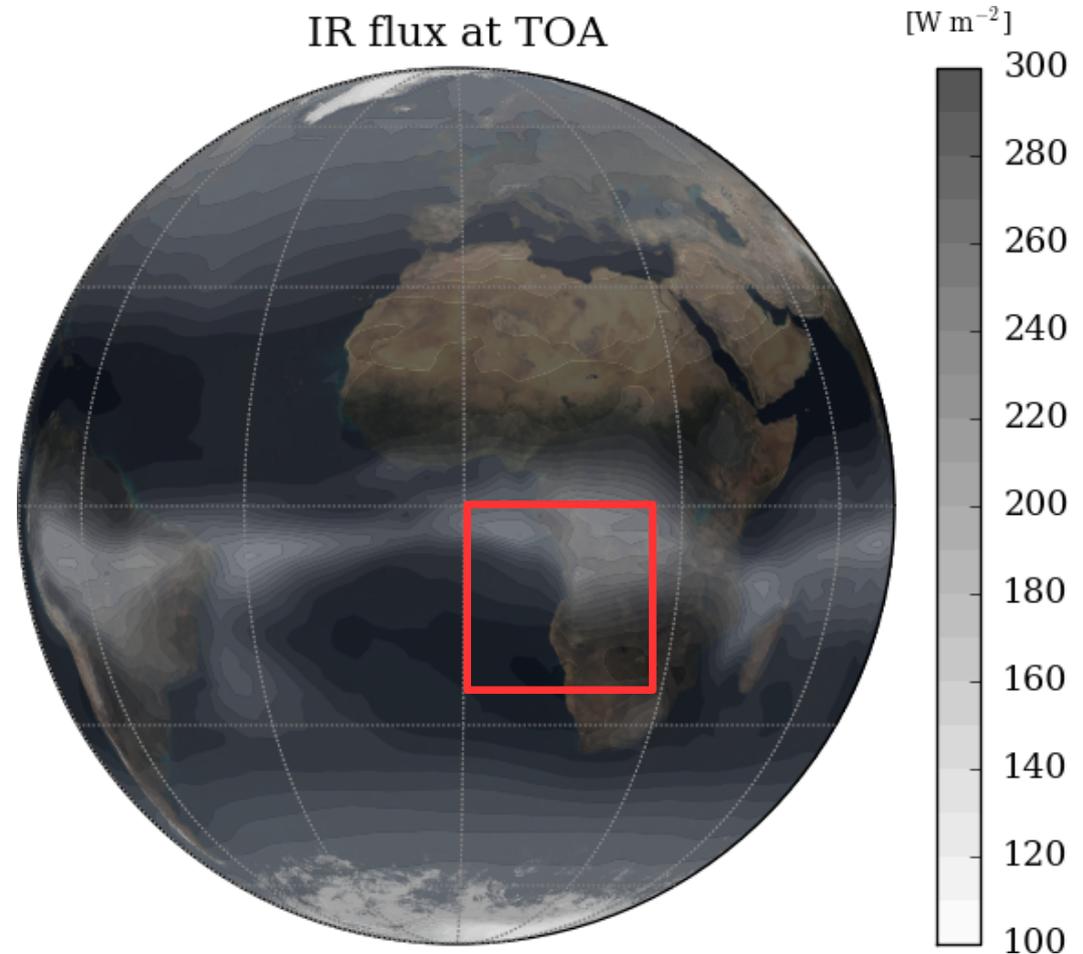


# A wide variety of processes



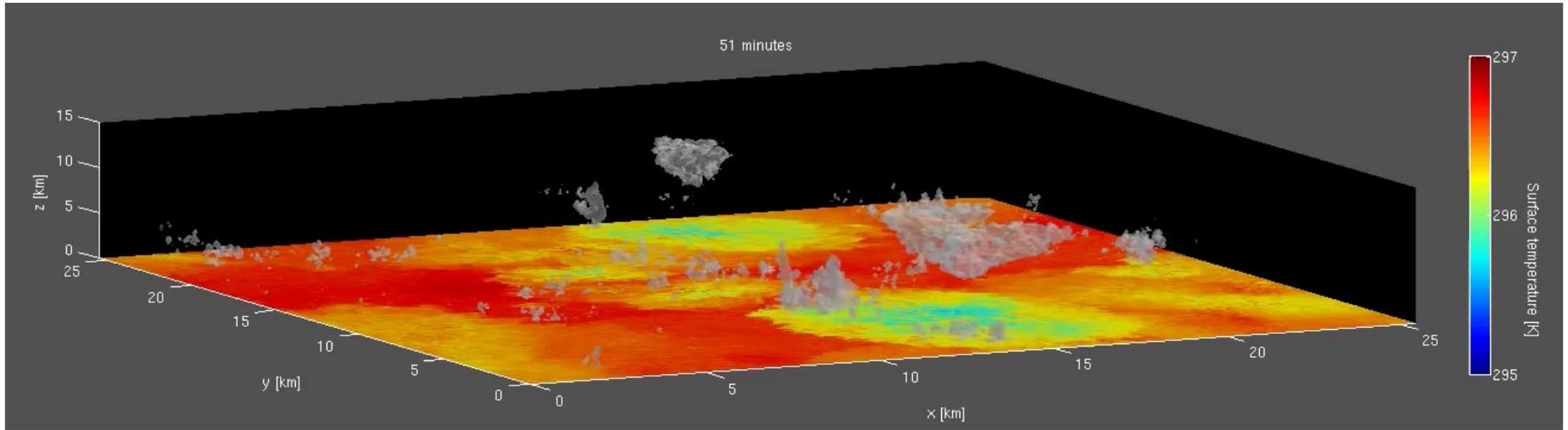
MET10 VIS006 2015-02-06 08:00 UTC

EUMETSAT



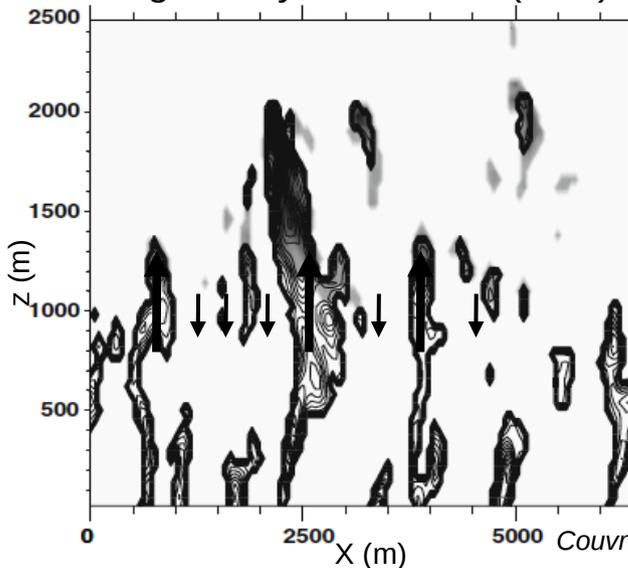
[IPSL Climate Model / 144x142 horizontal resolution / Graphisme: Planetoplot]

# Many processes in one grid cell

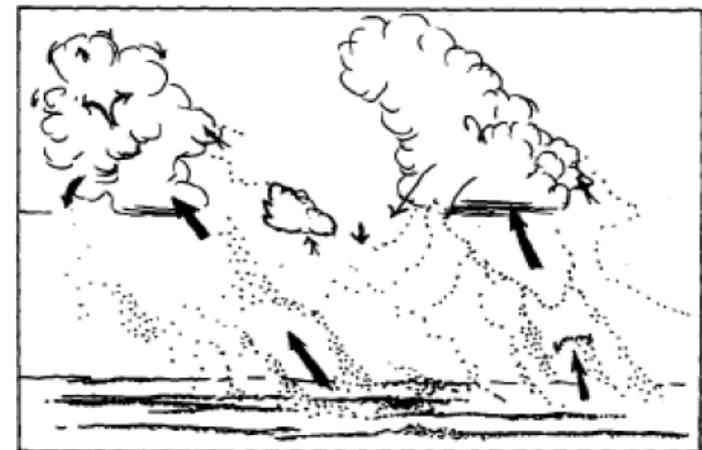


Around 8 hours of simulation by a **Cloud Resolving Model (CRM)** – C. Muller, LMD

## Thermals in a Large-Eddy Simulation (LES)

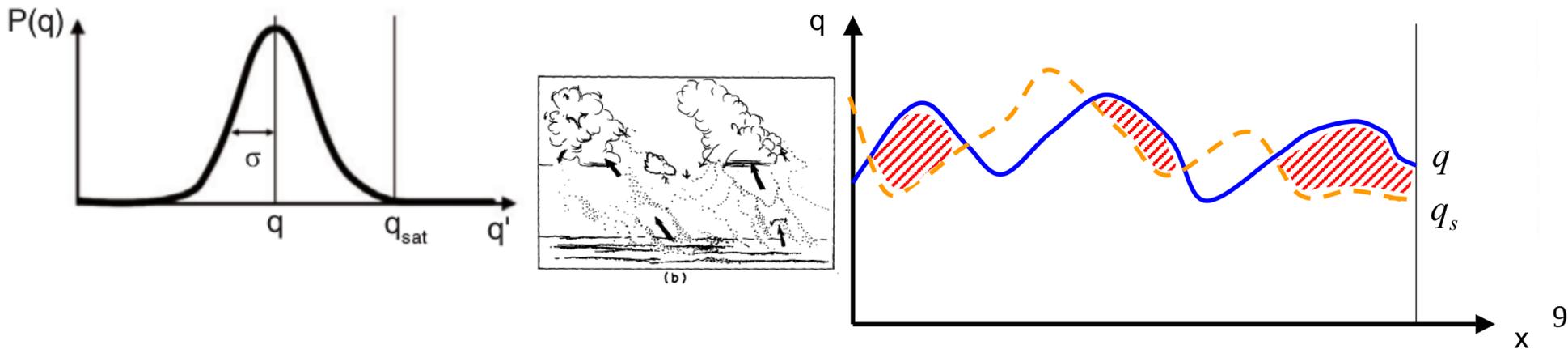
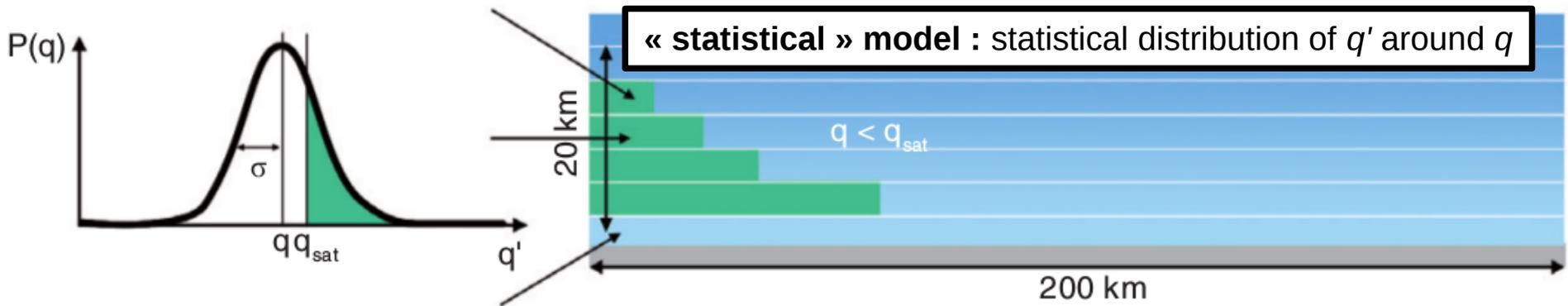
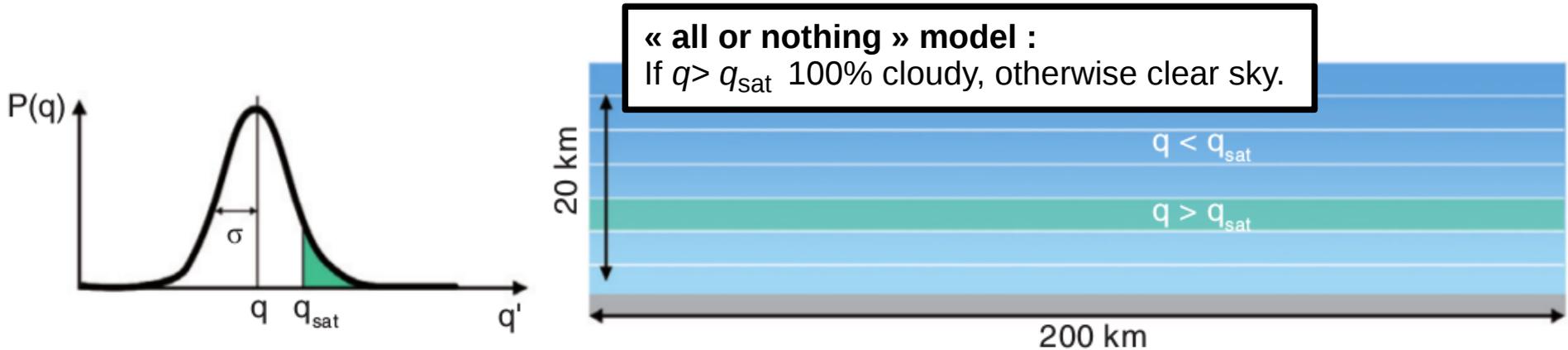


Conditional sampling of thermals based on a tracer emitted at the surface.

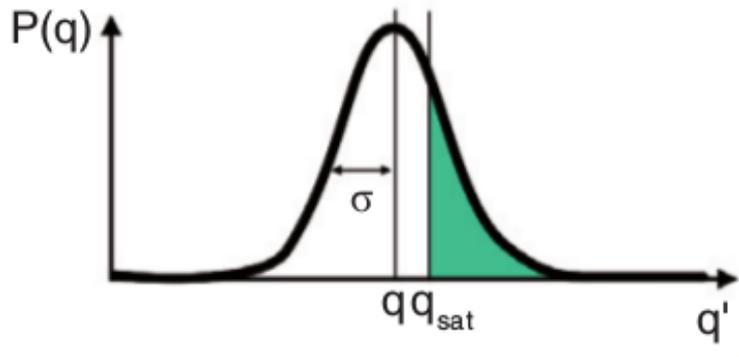
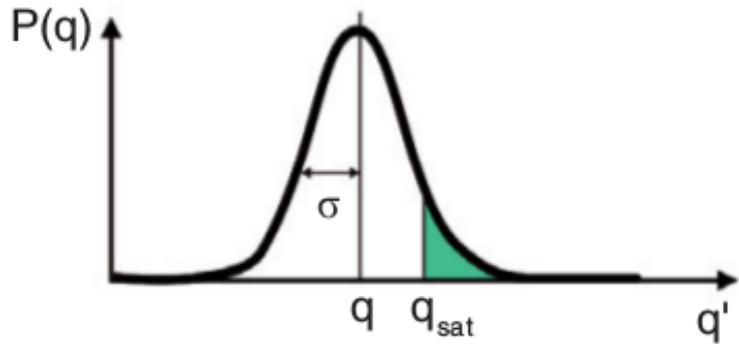


Lemone et Pennell, MWR, 1976

# Statistical cloud scheme



# Statistical cloud scheme 2/2



Mean total water content :

$$\bar{q} = \int_0^{\infty} q P(q) dq$$

Domain-averaged condensed water content :

$$q_c = \int_{q_{sat}}^{\infty} (q - q_{sat}) P(q) dq$$

Cloud fraction :

$$\alpha_c = \int_{q_{sat}}^{\infty} P(q) dq$$

In-cloud condensed water content :

$$q_c^{in} = \frac{q_c}{\alpha_c}$$

The goal of a cloud scheme is therefore to compute  $q_c^{in}$  and the cloud fraction based on the different physical parameterizations.

## physiq\_mod.F90 structure - I

Initialization (once) : *conf\_phys*, *phyetat0*,  
*phys\_output\_open*

Beginning *change\_srf\_frac*, *solarlong*

Cloud water evap. *reevap*

Vertical diffusion (turbulent mixing) *pbl\_surface*

1. Deep convection *conflx* (Tiedtke) or *concvl* (Emanuel)

Deep convection clouds *clouds\_gno*

Density currents (wakes) *calwake*

Strato-cumulus *stratocu\_if*

2. Thermal plumes *calltherm* and *ajsec* (sec = dry)

Thermal plume clouds *calcratqs*

3. Large scale condensation *fisrtlp*

Diagnostic clouds for Tiedtke *diagcld1*

Aerosols *readaerosol\_optic*

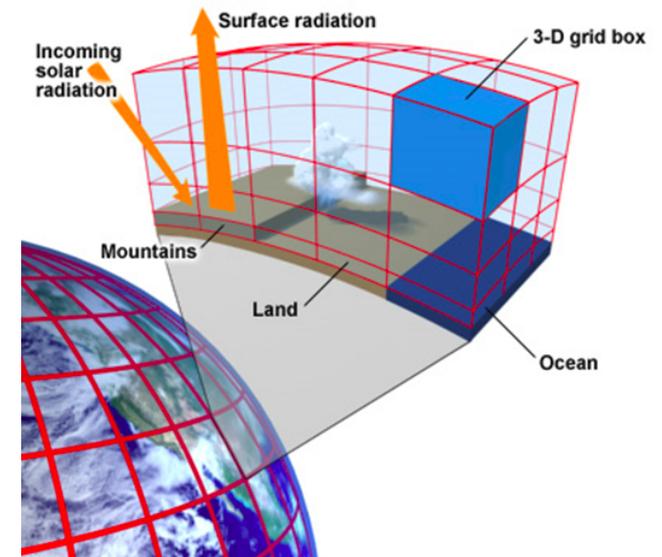
Cloud optical parameters *newmicro* or *nuage*

Radiative processes *radlwsu* (bis)

In blue : subroutines and instructions modifying state variables

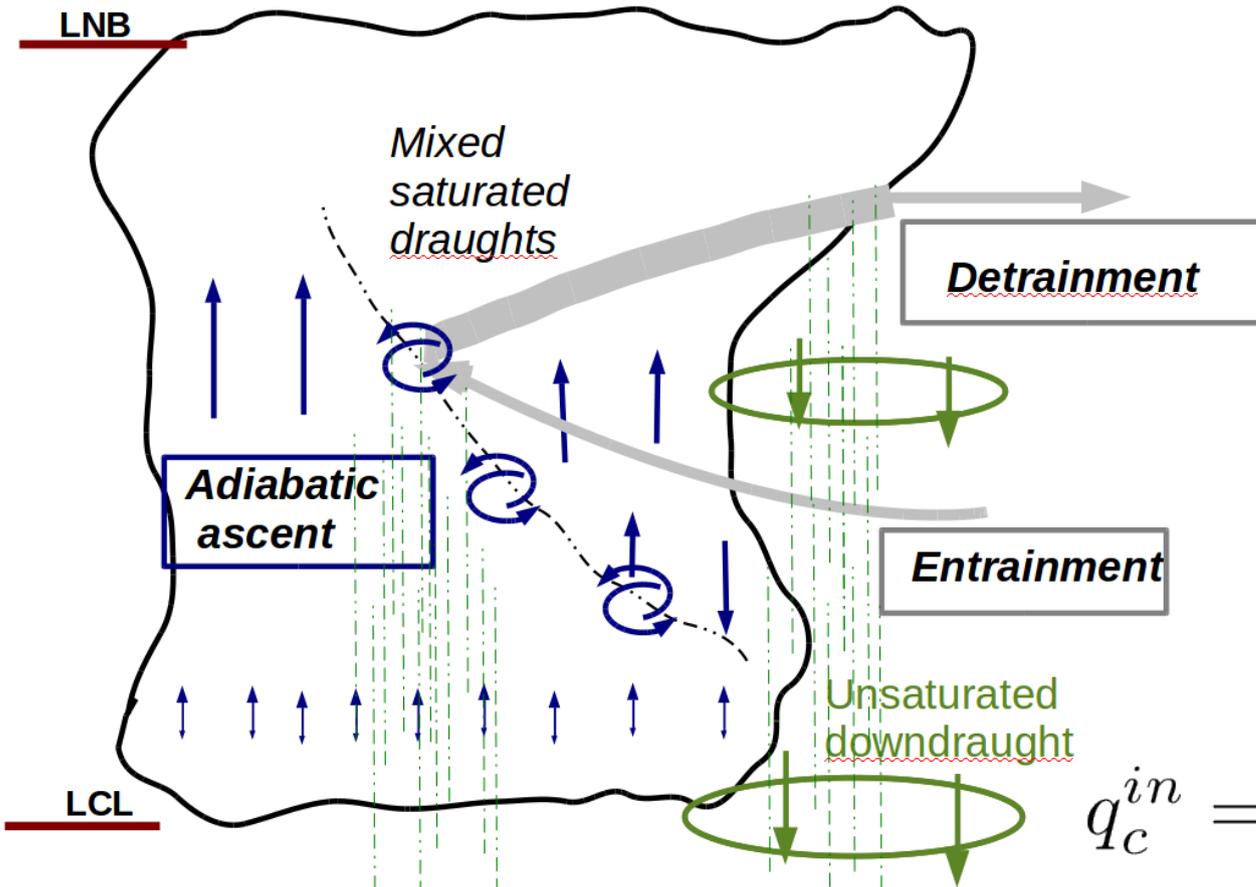
**CAREFUL** : clouds are evaporated/sublimated at the beginning of each time step (~15 min), but vapor, droplets and crystals are prognostic variables. In other words, **clouds can move but can't last for more than one timestep** (meaning that for example, crystals can't grow over multiple timesteps).

# LMDz physics parameterizations

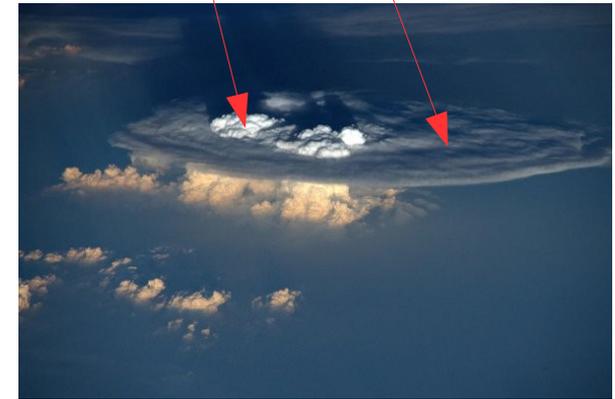


# 1. Deep convection

Emanuel scheme



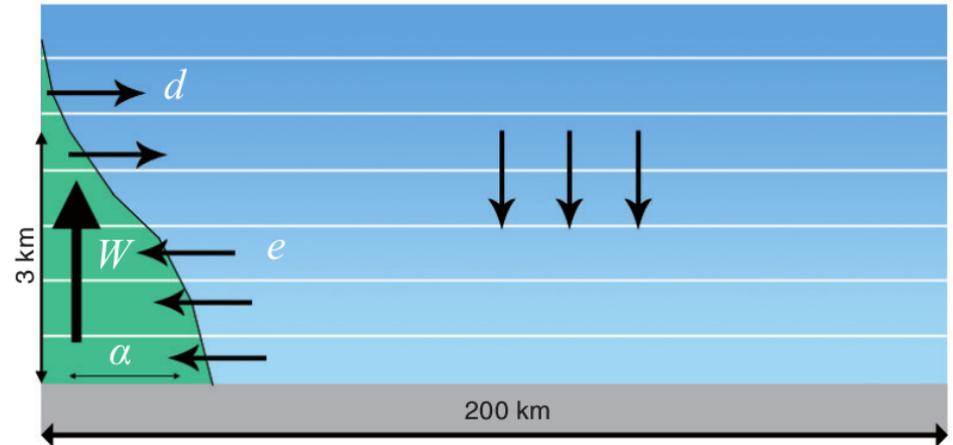
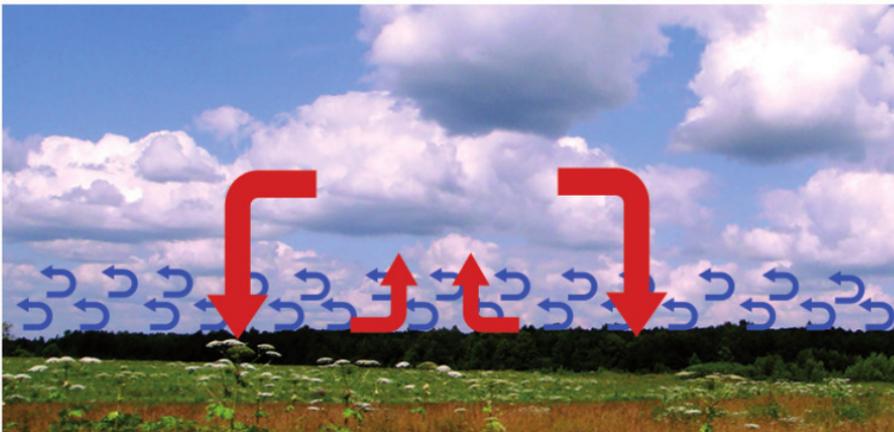
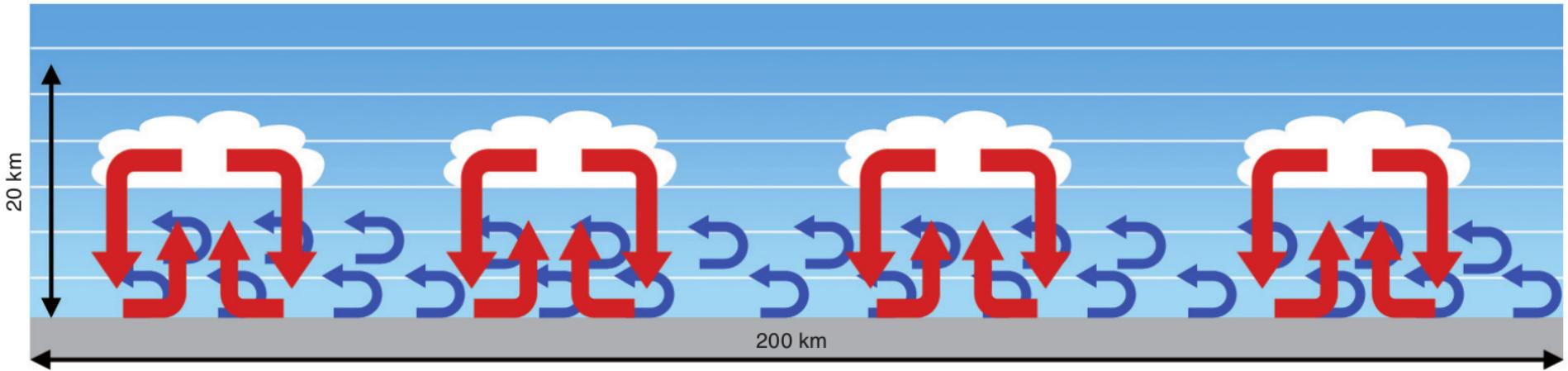
$$q_c^{in} = \frac{\sigma_a q_{ca} + \sigma_m q_{cm}}{\sigma_a + \sigma_m}$$



$$q_c^{in} = \frac{\frac{M_a}{\rho w_a} q_{ca} + \frac{\tau M_t g}{\delta p} q_{cm}}{\frac{M_a}{\rho w_a} + \frac{\tau M_t g}{\delta p}}$$

$q_c^{in}$  is computed by the deep convection scheme and  $\bar{q}$  is known  $\rightarrow$  cloud fraction is found

## 2. Shallow convection 1/2



# 2. Shallow convection 2/2

Bi-Gaussian distribution of saturation deficit  $s$ :

$$s = a_1(q_t - q_{\text{sat}}(T_1))$$

- One mode associated with thermals

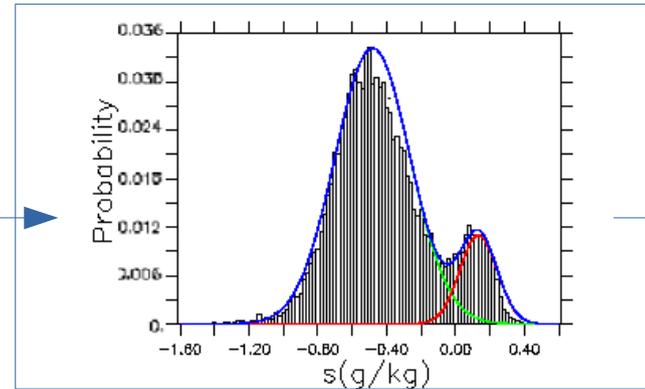
$$s_{\text{th}}, \sigma_{\text{th}}$$

- One mode associated with their environment:

$$s_{\text{env}}, \sigma_{\text{env}}$$

$$s_{\text{env}}, \sigma_{\text{env}}, s_{\text{th}}, \sigma_{\text{th}}, \alpha$$

Shallow convection



$q_c, CF$

*Jam & al., BLM, 2013*

We know:

Mean state:  $s_{\text{env}}$

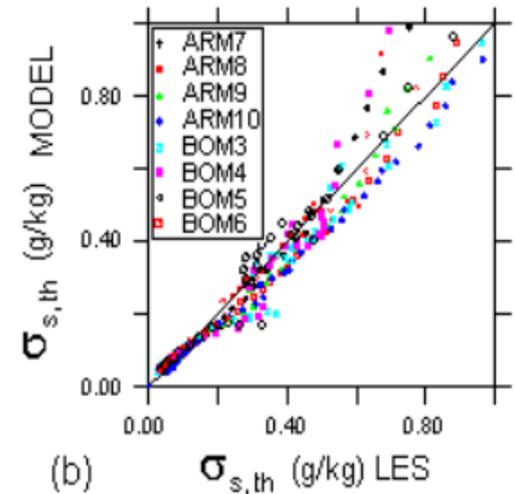
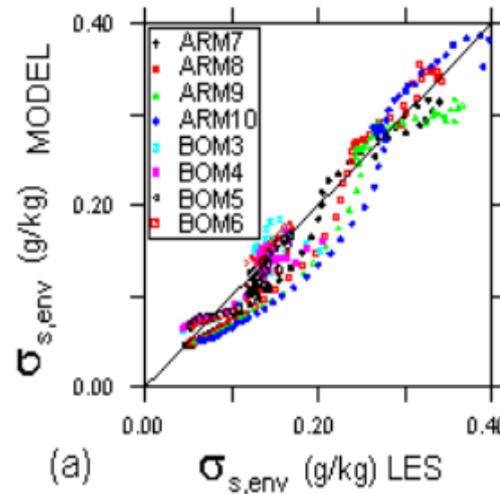
Thermal properties:  $s_{\text{th}}, \alpha$

Parameterization of the variances:

$$\sigma_{s,\text{env}} = c_{\text{env}} \frac{\alpha^{\frac{1}{2}}}{1 - \alpha} (\bar{s}_{\text{th}} - \bar{s}_{\text{env}}) + b \bar{q}_{t_{\text{env}}}$$

$$\sigma_{s,\text{th}} = c_{\text{th}} \alpha^{-\frac{1}{2}} (\bar{s}_{\text{th}} - \bar{s}_{\text{env}}) + b \bar{q}_{t_{\text{th}}}$$

$q_c^{\text{in}}$  is deduced from the mean water content of the environment and thermals and the parameterized spreads of the two gaussian distributions



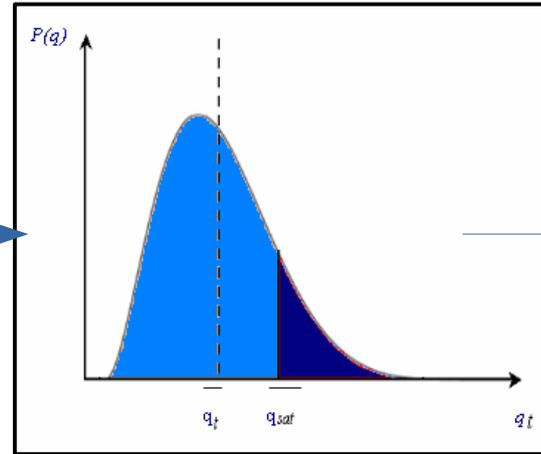
# 3. Large scale condensation

Log-normal distribution of total water  $q_t$  (Bony & Emanuel, JAS, 2001)

Grid-cell mean state

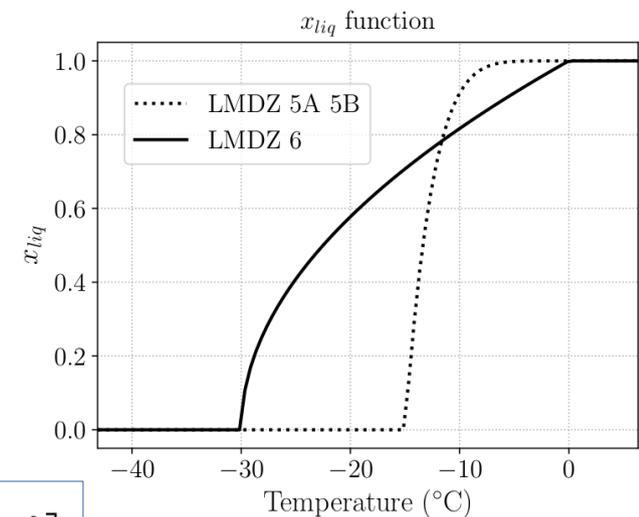
→  $q, q_{sat}$

$\sigma/q$  imposed (called “ratqs”)



$$\alpha_c = \int_{q_{sat}}^{\infty} P(q) dq$$

$$q_c = \int_{q_{sat}}^{\infty} (q - q_{sat}) P(q) dq$$



- condensate: liquid/ice partitioning (function of the temperature) :

$$^a \text{Cloud liquid fraction} = \left( \frac{T - T_{\min}}{T_{\max} - T_{\min}} \right)^n, \text{ for } T_{\min} \leq T \leq T_{\max}$$

- A fraction of the condensate falls as rain (parameters controlling the maximum water content of clouds and the auto-conversion rate) :

$$\frac{dq_{lw}}{dt} = -\frac{q_{lw}}{\tau_{convers}} \left[ 1 - e^{-(q_{lw}/clw)^2} \right]$$

- The rain is partly evaporated in the grid below (parameter controlling the evaporation rate) :

$$\frac{\partial P}{\partial z} = \beta [1 - q/q_{sat}] \sqrt{P}$$

$$\frac{dq_{iw}}{dt} = \frac{1}{\rho} \frac{\partial}{\partial z} (\rho w_{iw} q_{iw})$$

$$w_{iw} = \gamma_{iw} w_0$$

$$w_0 = 3.29 (\rho q_{iw})^{0.16}$$

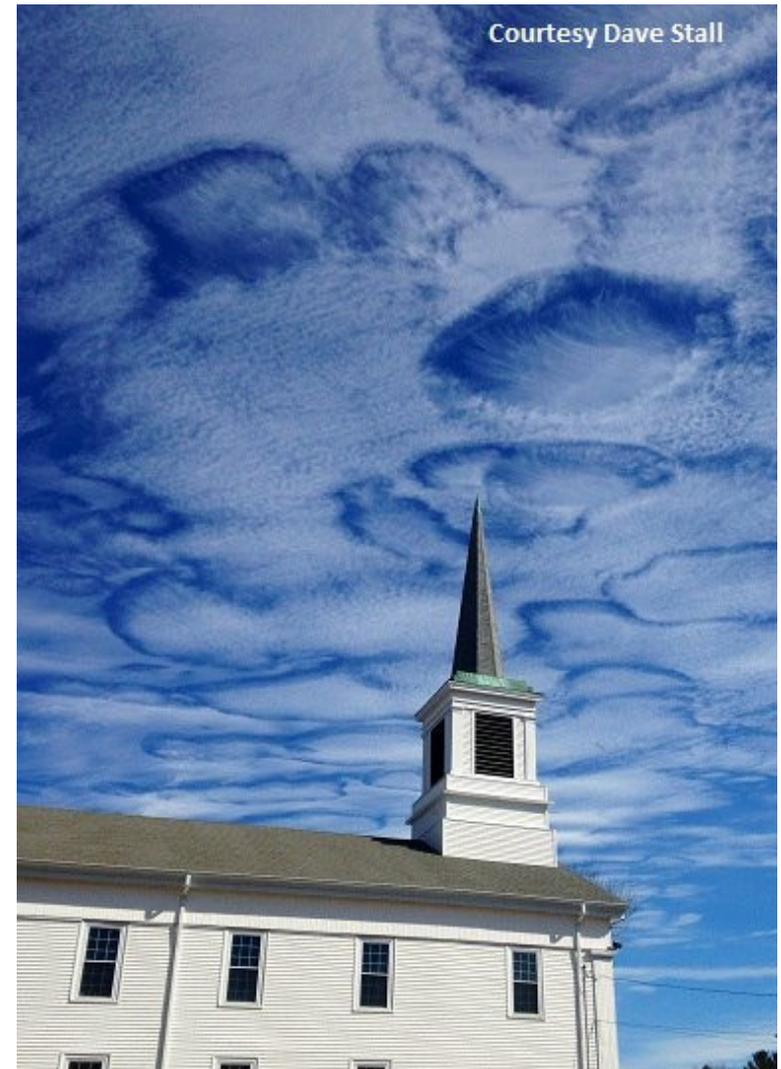
# 3. “Bergeron” effect



Growth of an ice crystal at the expense of surrounding supercooled water drops  
[Wallace, 2005]

$$\frac{dq_{lw}}{dt} = -\frac{q_{lw}}{\tau_{convers}} \left[ 1 - e^{-(q_{lw}/clw)^2} \right]$$

- Can occur under negative temperatures → the resulting liquid precipitation **is converted to ice**
- when freezing, rain releases latent heat, which can potentially bring the temperature back to above freezing. If this is the case, a small amount of rain remains liquid to stay below freezing and stabilize the numerical scheme.

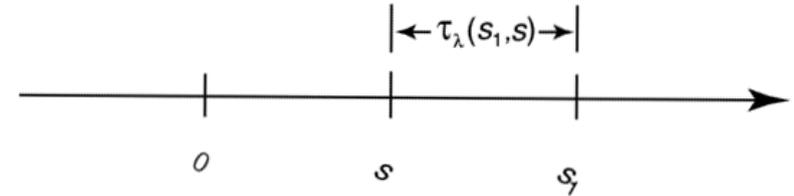


Fallstreak hole (also known as hole punch clouds), Rhode Island, USA

# Radiative transfer

**Radiative transfer equation :**

$$-\mu \frac{\partial I_\lambda}{\partial \tau_\lambda}(\tau_\lambda, \mu, \Phi) = -I_\lambda(\tau_\lambda, \mu, \Phi) + S_\lambda(\tau_\lambda, \mu, \Phi) + \frac{w_{0\lambda}}{4\pi} \int_0^{2\pi} \int_{-1}^1 P_\lambda(\mu, \mu', \Phi, \Phi') I_\lambda(\tau_\lambda, \mu', \Phi') d\mu' d\Phi'$$



$$\mathbf{q_{c,tot}} = \mathbf{q_c^{in} (thermals) \times CF (thermals)} \longleftarrow \mathbf{bi-gaussian PDF}$$

$$+ \mathbf{q_c^{in} (convection) \times CF (convection)}$$

$$+ \mathbf{q_c^{in} (large-scale) \times CF (large-scale)} \longleftarrow \mathbf{Lognormal PDF}$$

$$\mathbf{CF_{tot}} = \mathbf{\min( CF (thermals) + CF (convection) + CF (large-scale), 1.)}$$

Solving the radiative transfer equation requires :

- $\mathbf{q_{c,tot}}$  to compute the optical depth ;
- **Cloud droplet and crystal sizes** to compute the optical properties ;
- $\mathbf{CF_{tot}}$  to compute the heating rates in the clear-sky ( $1 - \mathbf{CF_{tot}}$ ) and cloudy ( $\mathbf{CF_{tot}}$ ) columns.

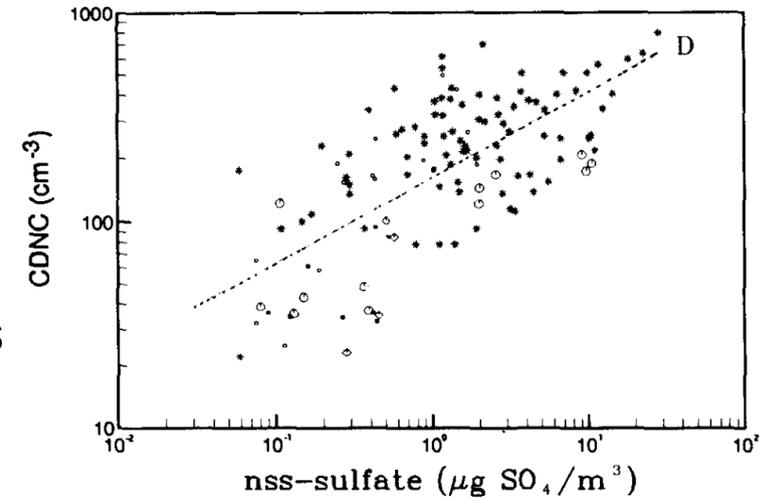
# Optical properties of liquid clouds

(see O. Boucher's talk)

ok\_cdnc = y  
bl95\_b0 = 1.3  
bl95\_b1 = 0.2

$$\text{CDNC} = 10^{b0 + b1 \log(m \text{ SO}_4)}$$

Link cloud droplet number concentration to aerosol mass concentration (Boucher and Lohmann, Tellus, 1995)  
 $m\text{SO}_4 \rightarrow$  Now uses mass of all soluble species.



$$N = \text{CDNC}$$

$$r_3 = \left( \frac{l \rho_{\text{air}}}{(4/3) \pi \rho_{\text{water}} N} \right)^{1/3}$$

$$r_e = \frac{\int r^3 n(r) dr}{\int r^2 n(r) dr}$$

$$r_e = 1.1 r_3$$

Size-dependent computation of cloud optical properties (Fouquart [1988] in the SW, Smith and Shi [1992] in the LW)

# Optical properties of ice clouds

Optical properties are computed using Ebert and Curry [1992], based on the computed crystal sizes.

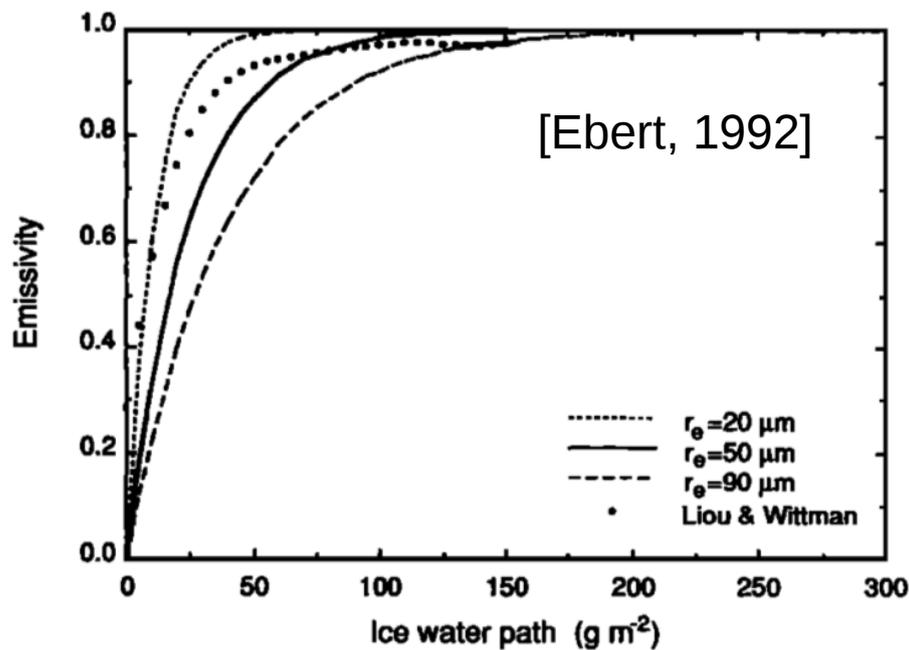
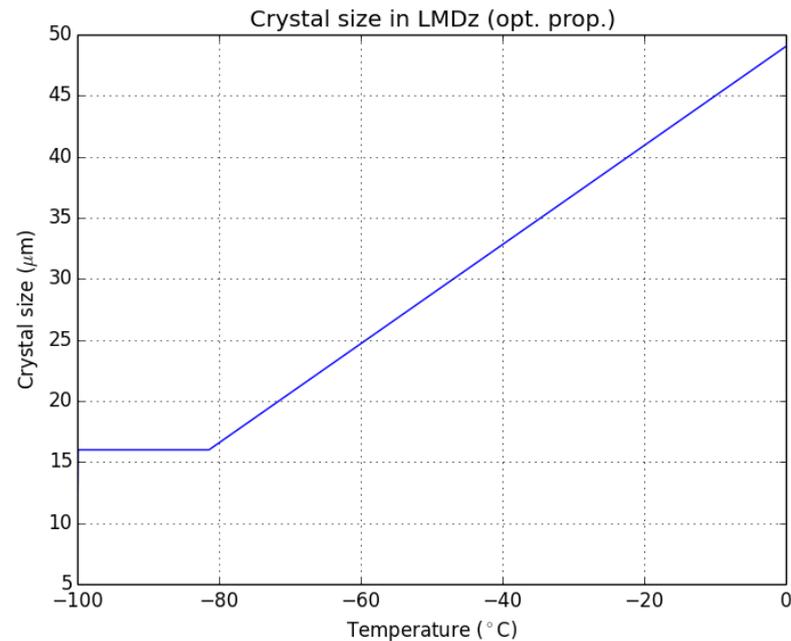


Fig. 5. Cirrus infrared emissivity for  $r_e = 20, 50, \text{ and } 90 \mu\text{m}$  as a function of ice water path. The solid circles represent values computed using the parameterization of *Liou and Wittman* [1979].



**Crystal sizes follow**

$$r = 0.71T + 61.29 \text{ in } \mu\text{m}$$

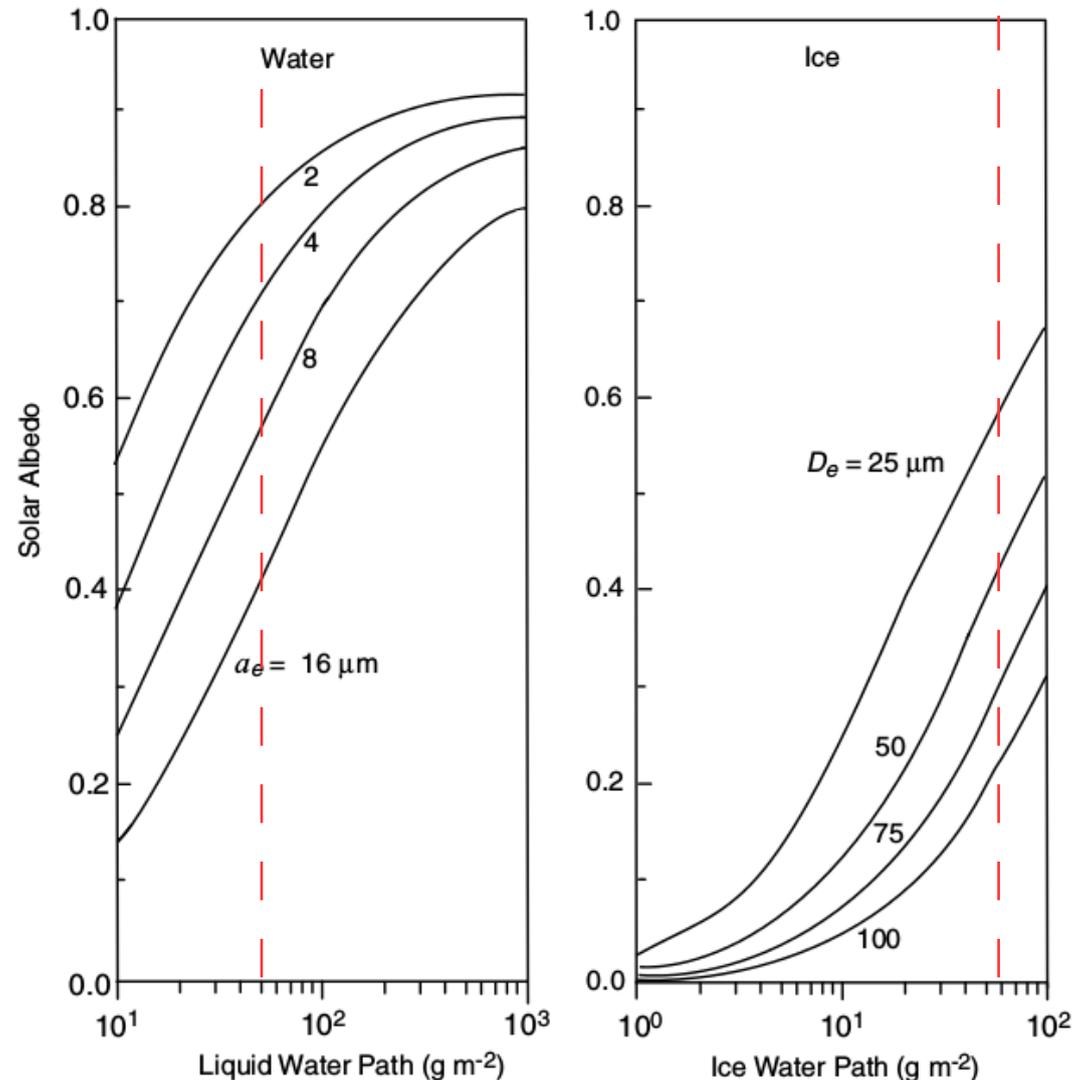
[*Iacobellis et Somerville 2000*]

with  $r_{\min} \sim 10 \mu\text{m}$  (tuneable)

for  $T < -81.4^\circ\text{C}$  [*Heymsfield et al. 1986*]

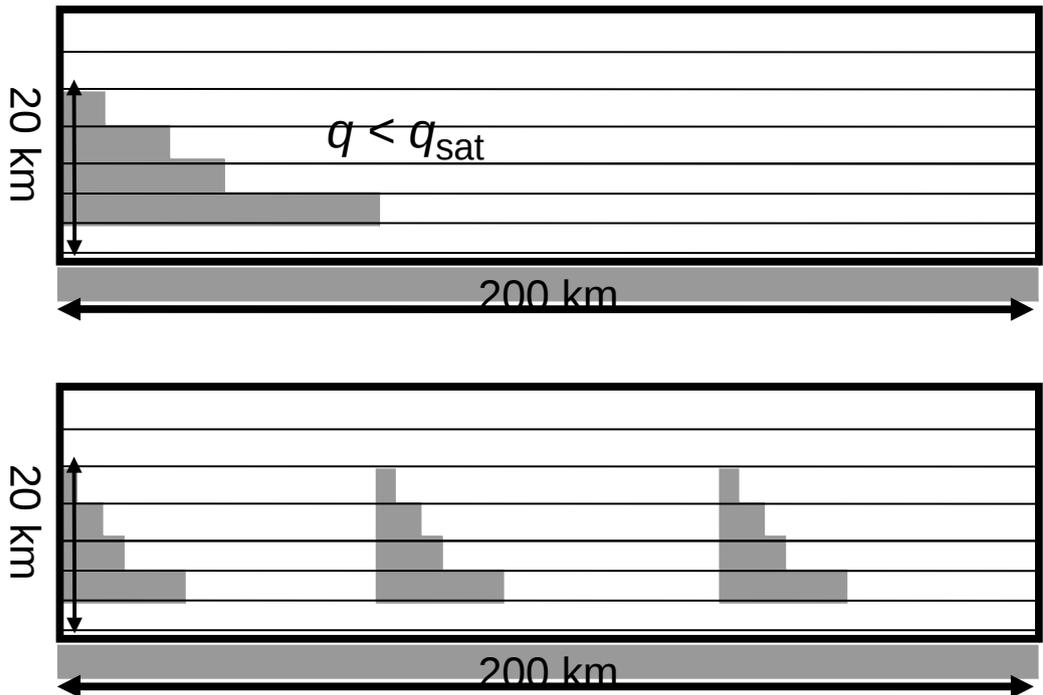
# Importance of cloud phase

- Clouds reflect sunlight (negative forcing, cooling) and emit in the infrared (positive forcing, warming) ;
- For the same water content, liquid clouds reflect more sunlight than ice clouds ;
- For liquid clouds : if the cloud water content increases, there is a negative forcing (reflection dominates) ;
- For ice clouds : if the cloud water content increases, the forcing depends on the size of the crystals.



[Liou 2002]

CF versus height is known, but radiation also needs to know the total cloud **cover** ; we therefore parameterize the **cloud overlap**

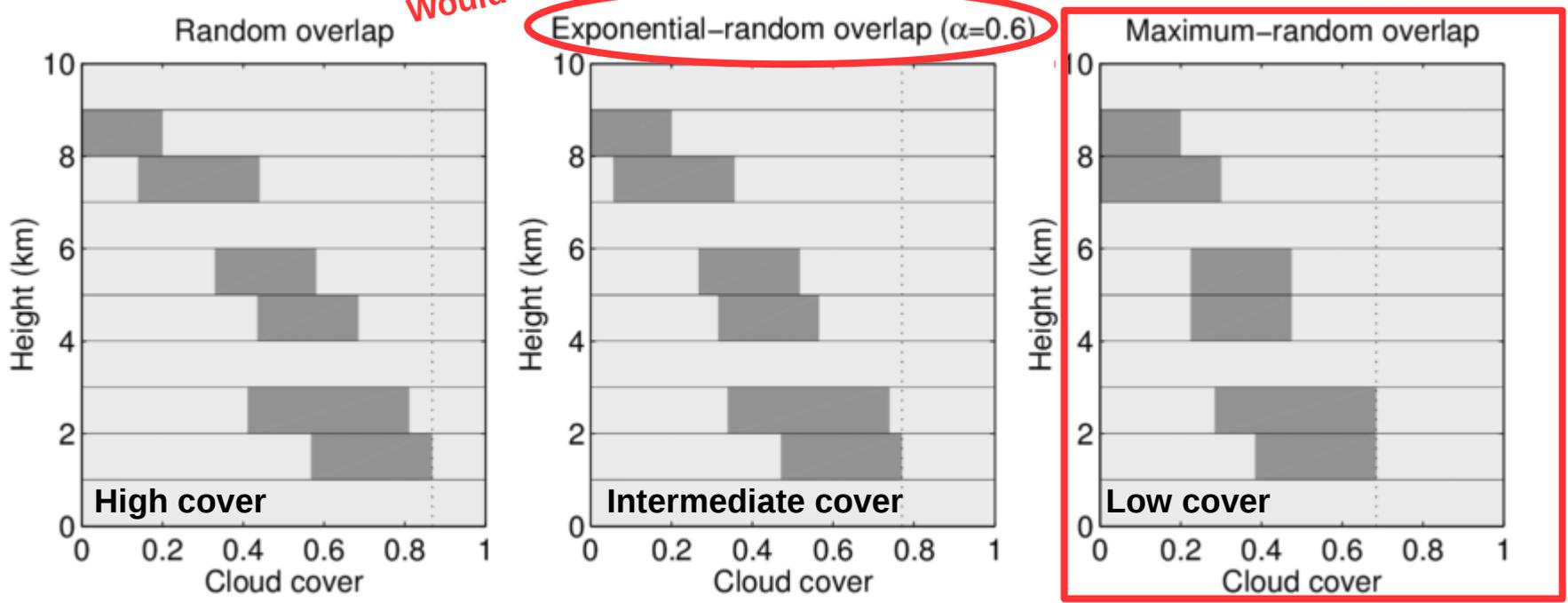


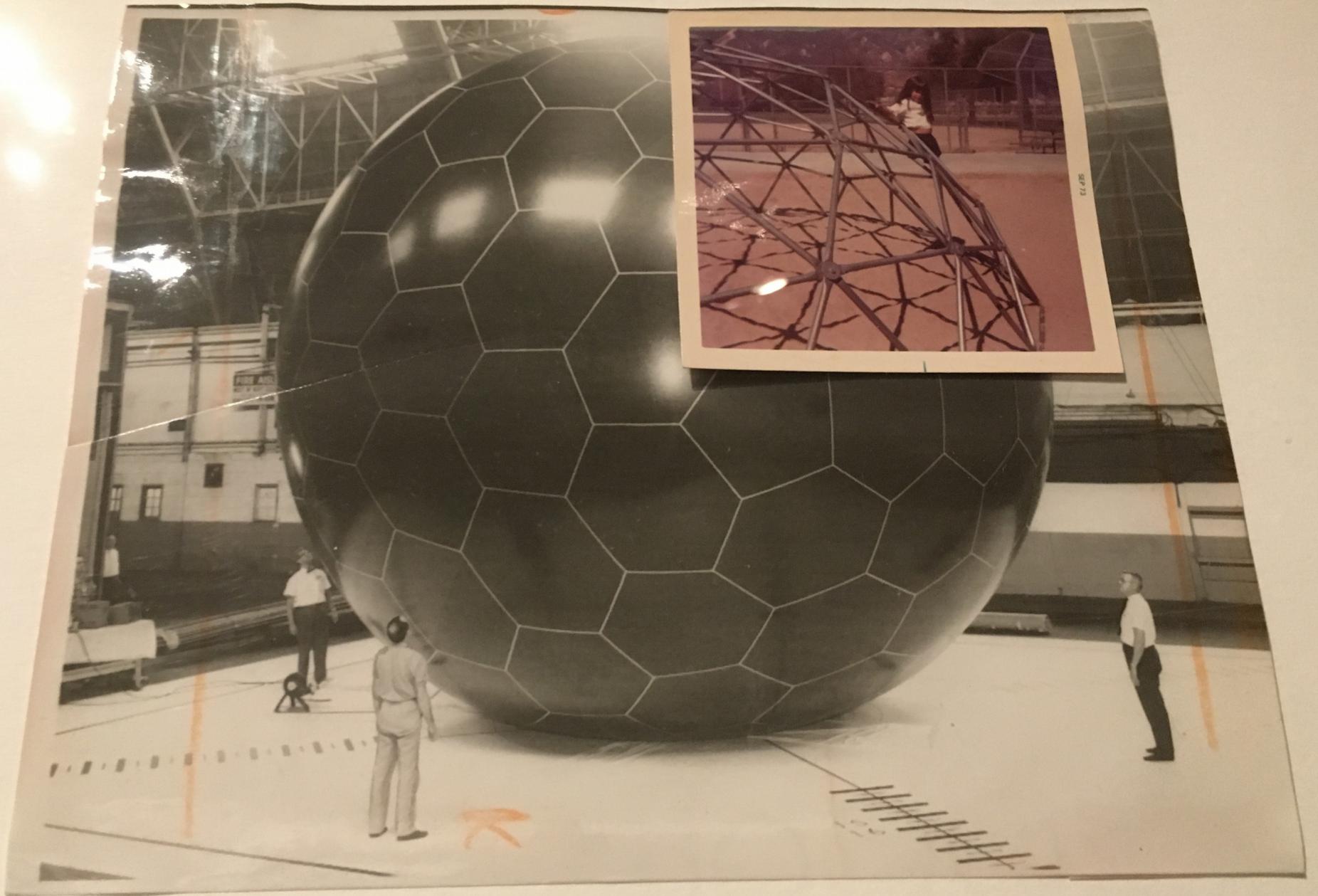
**LMDz : Maximum random overlap**  
 For the GCM, these two scenes are identical ;



*Would be better !*

Used in LMDz





**Welcome to the LMDz team !**