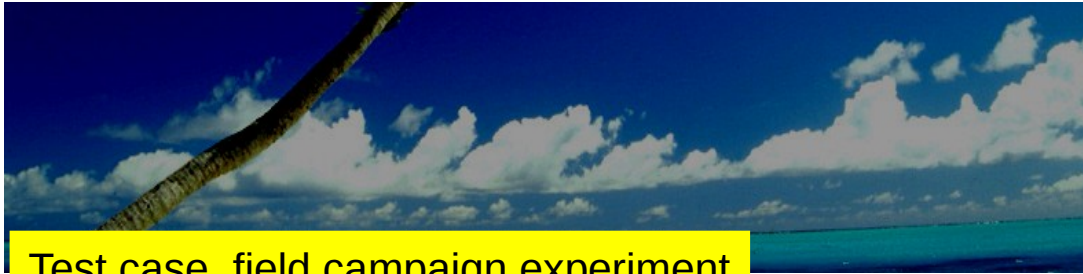


# LMDZ Single Column Model

- + what is it ?
- + why is it interesting ?
- + List of 1D cases
- + how to install and run it ?

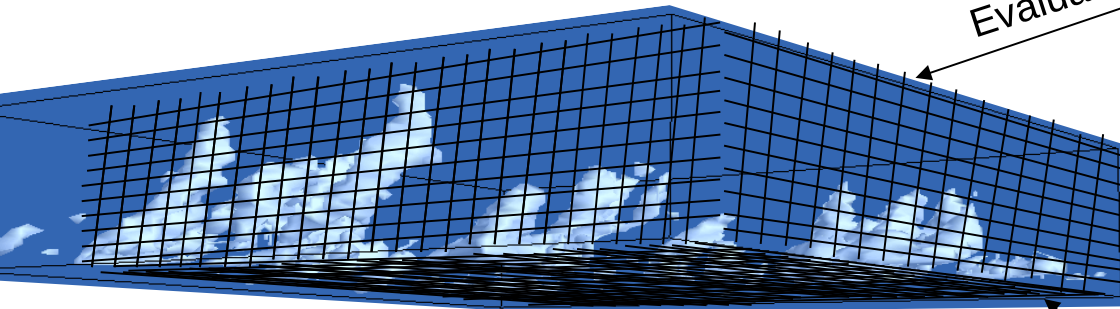
# What is it ?



Observation



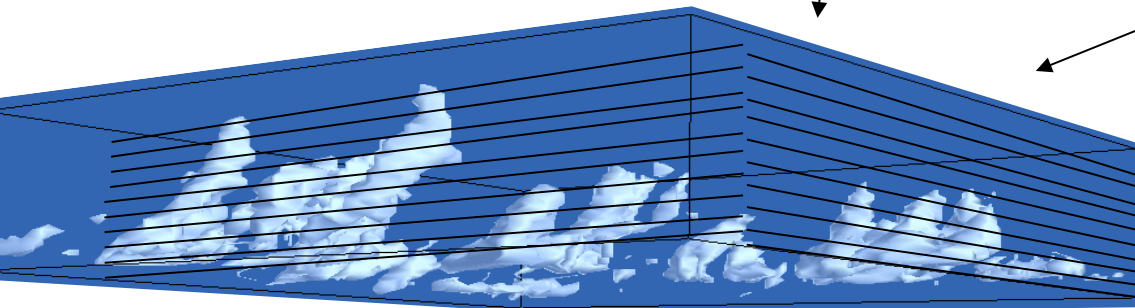
Evaluation



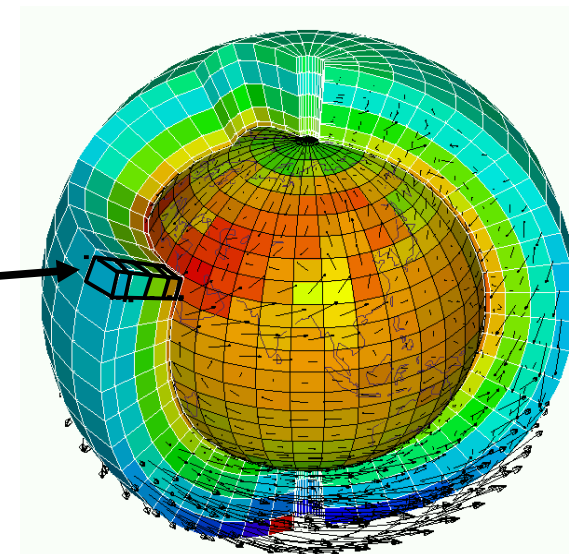
Explicit simulations, Grid cell, 20-100 m

Evaluation

« Large scale »  
conditions  
imposed



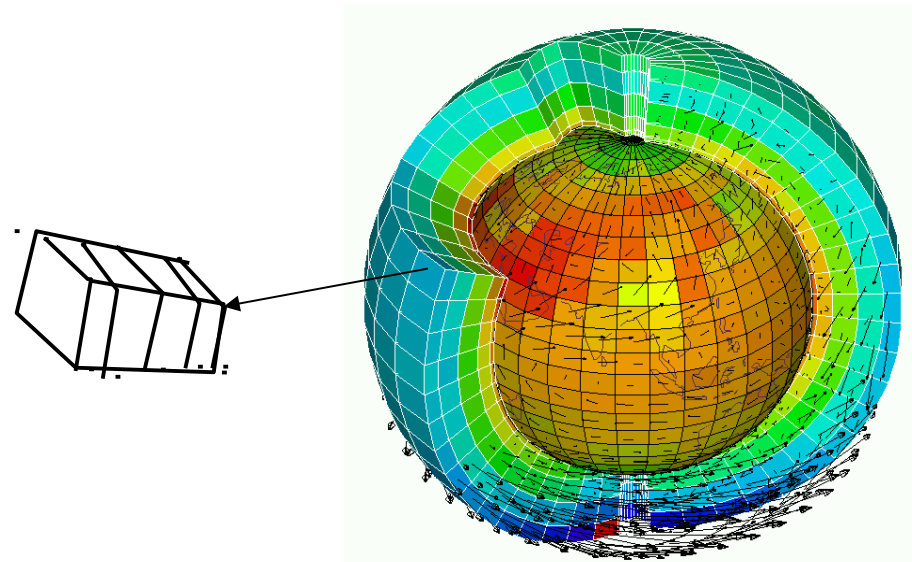
Climate model, parameterizations, « single-column » mode



Courtesy F.Hourdin

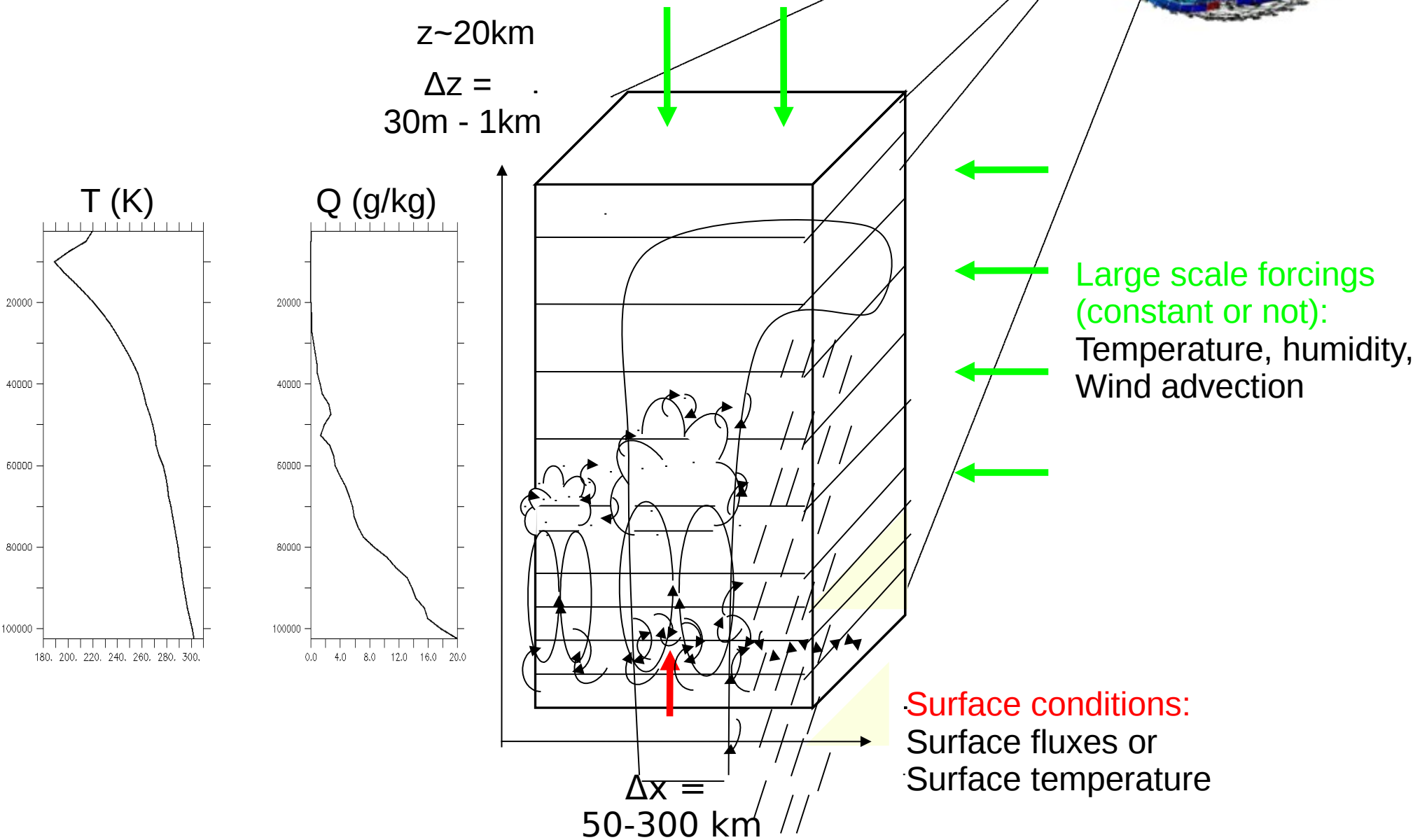
3D is a collection of many “single column models”, covering earth and interaction with each other through a set of rules known as “large scale dynamics”.

In a 1D model, there is no dynamics. We use observations or model output or idealized forcings at the boundaries of the column.



# LMDZ model in 1D mode

- We impose large scale conditions.
- Duration of the case varies from few hours to few months
- We study parameterizations in a given environment.



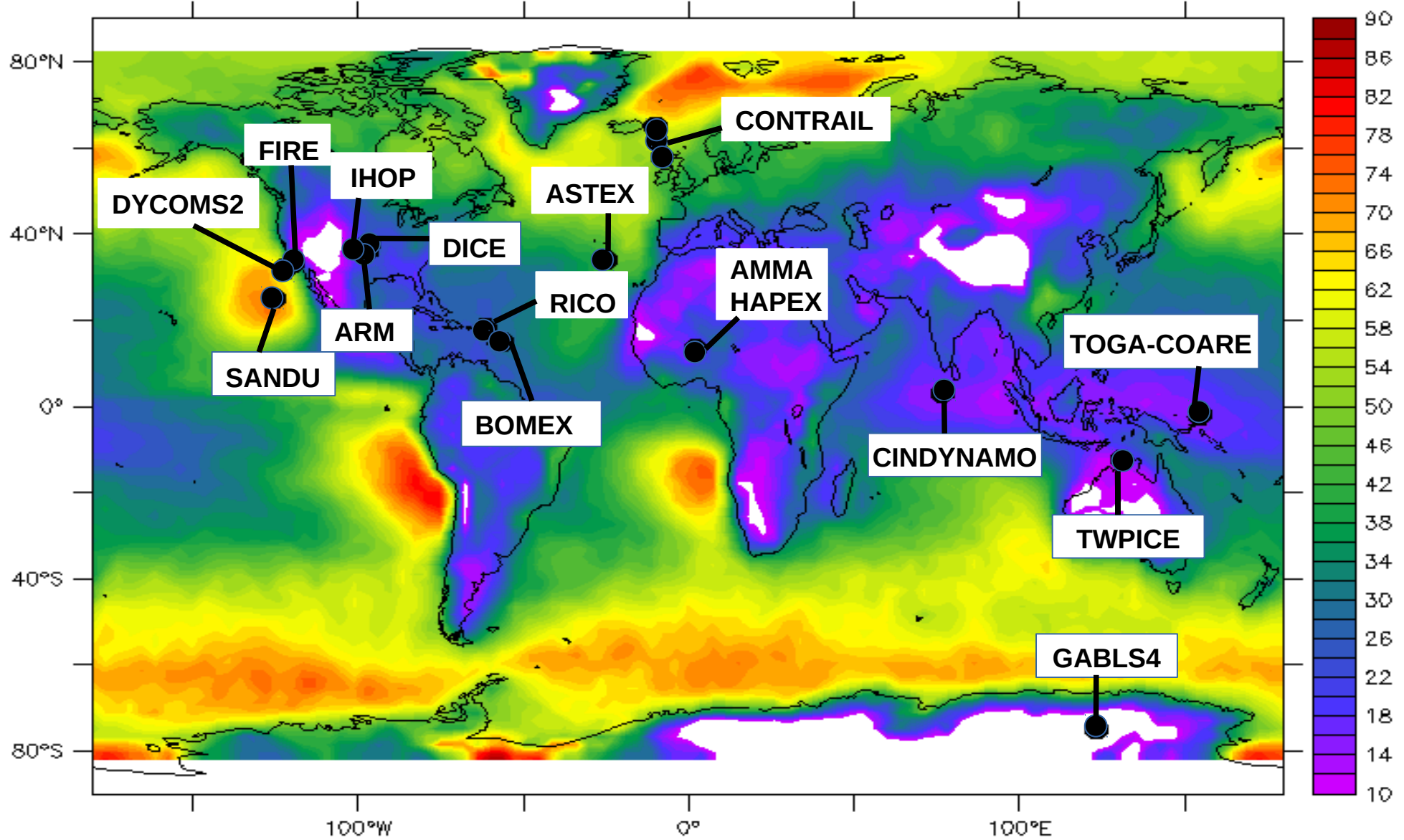
## Why is it interesting ?

- + **simple tool**: technical and understanding, usable on any laptop
- + it's a useful tool for **parameterization development** in different meteorological situation: shallow convection, deep convection, transition from stratocumulus to cumulus, stable boundary layer, radiation...
- + we can evaluate behavior of physical parameterizations **comparing results to observations or to explicit simulations** (CRM, LES)
- + then we go back to GCM: test and debug new parameterizations
- + we can anticipate the effect of new development in 3D simulations
- + we have hierarchy of models: SCM, LAM, AGCM, GCM ...

## **List of 1D cases**

The cases are located in different places of the world  
and represent various meteorological situations

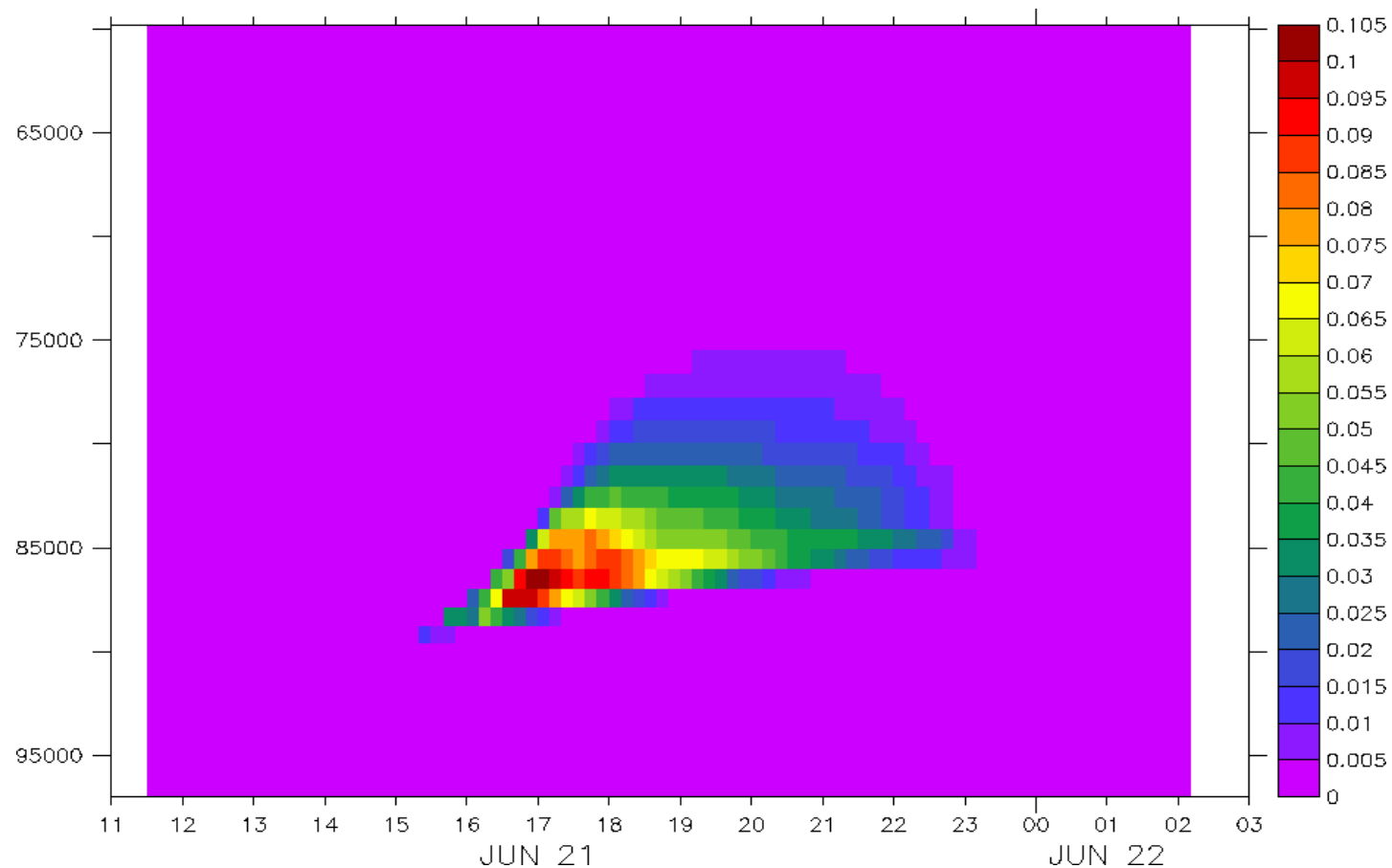
## Where are located all these cases ?



*Background : low cloud cover from Calipso (Chepfer et al. 2008)*

## Dry and shallow convection

- **Arm\_cu** (diurnal cycle of shallow cumulus over land)
- **Rico** (Rain In Cumulus over Ocean, shallow precipitating cumulus over sea)
- **Ayotte** (convective boundary layer, sky clear )

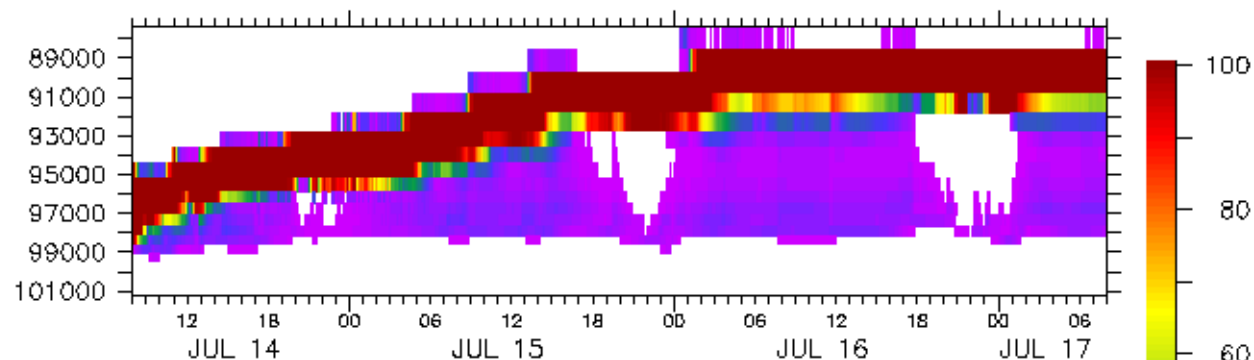


Arm\_cu case - Cloud fraction

# Stratocumulus and transition to cumulus

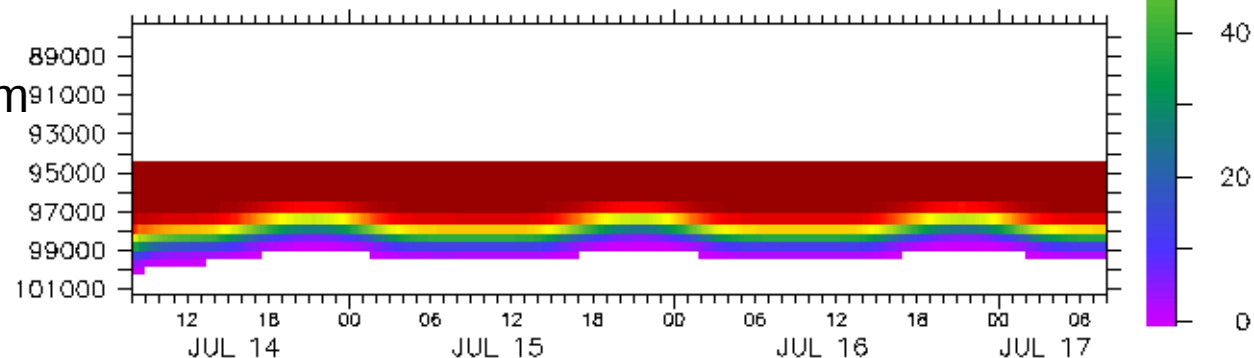
- **Sandu** (transition case with 3 options according to variation of SST)
- **Fire** (diurnal cycle of stratocumulus)

Fire case:  
Cloud fraction (%)



Top: standard version  
Bad representation because not  
Enough entrainment at the cloud top

Bottom: version developped by A.Jam

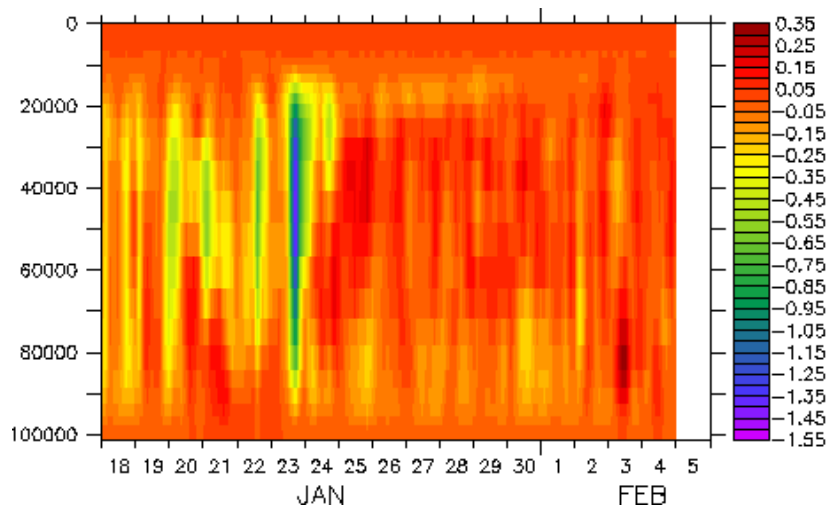


## Deep convection:

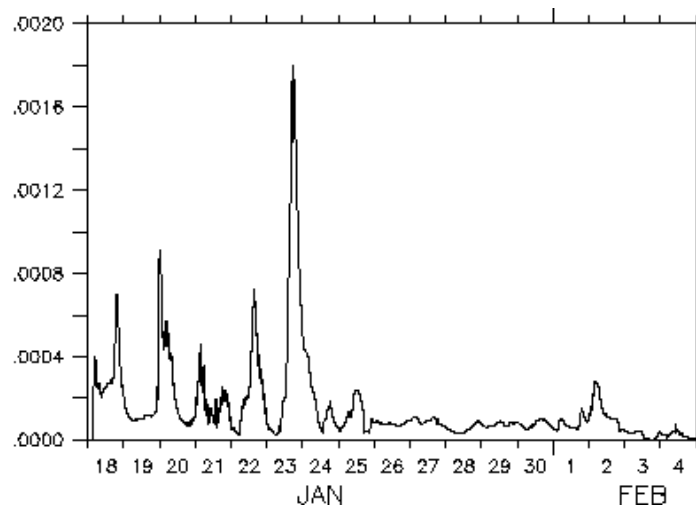
### Over ocean:

- **Toga**
- **case\_e** (part of Toga)
- **TWPICE** : off the coast of Darwin
- **Cindy Dynamo** : MJO study

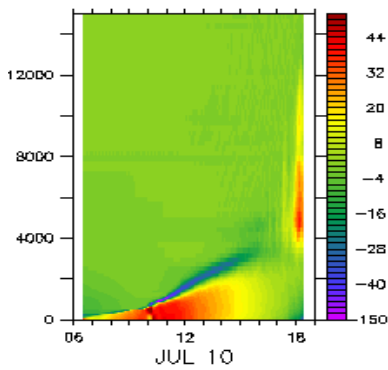
## **TWPICE Case (2 weeks)**



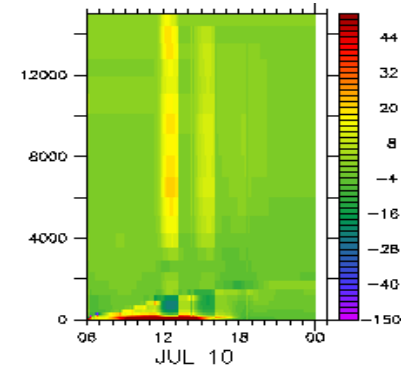
Forcings vertical speed  
(m/s)



Precipitation (mm/j)



LES from 6h **to 18h**



LMDZ\_AR4\_L39  
From 6h **to 00h**

AMMA case (10h july 2006):  
Thetal tendencies due to all schemes (K/j)

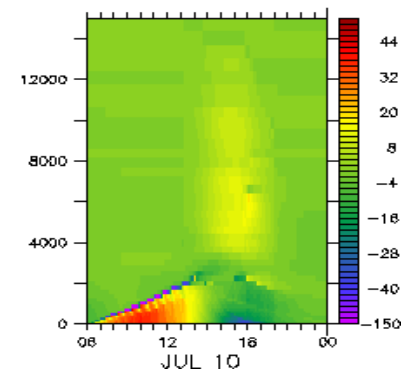
### Deep convection:

#### Over land:

- **Hapex**
- **AMMA**

#### Idealized case:

- **eq\_rad\_conv** (RCE) : radiative and convection scheme active



LMDZ\_NP\_L70 from 6h **to 00h**

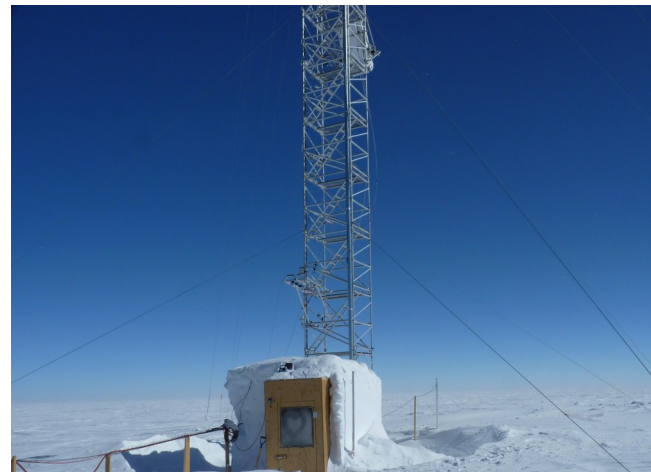
## Stable boundary layer:

- **Dice**
- **GABLS4**

We can run these cases with  
atmosphere forced or coupled with  
Orchidee model



**DICE** case : characterize boundary layer  
In the site of SGP during 3 days/night



**GABLS4** case : interaction of a very  
stable boundary layer with a snow  
surface

## How to install and run it ?

+ install LMDZ 3D with **install\_lmdz.sh**

+ install LMDZ 1D :

- \* **wget** <http://www.lmd.jussieu.fr/~lmdz/pub/1D/1D.tar.gz>

- \* **extract** 1D directory → creates 1D directory

- \* **cd** 1D; **./run.sh**

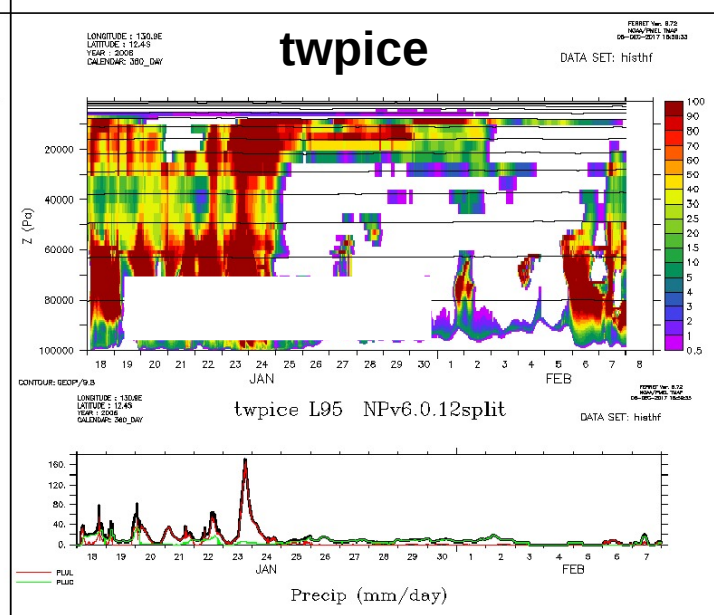
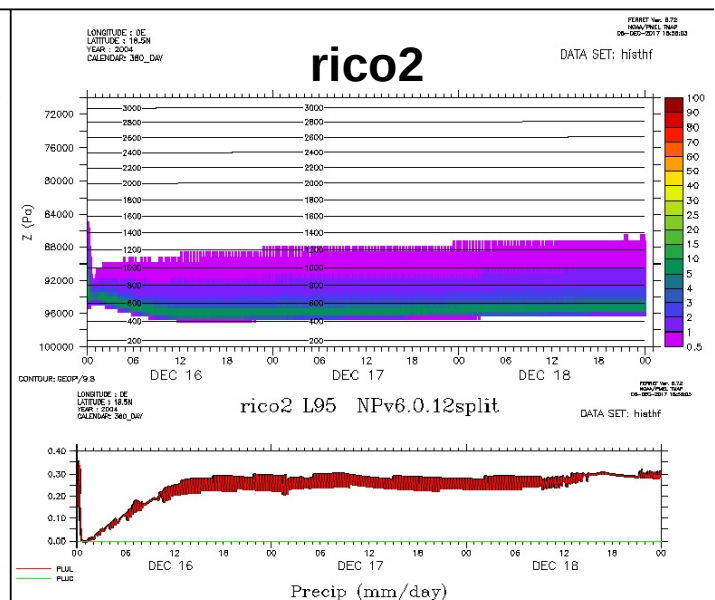
- \* **runs automatically 6 cases** ( dice\_bucket

arm\_cu2 rico2 fire sandufast twpice) with 1 physical package

- \* **shows some output:** cloud cover + precipitation

plot

## Cloud cover



# What can you do in each case directory ?

- + **compile & run** with run.sh: choose case, physical package and level number
- + Look at **initial profiles and forcings** (ascii files or netcdf files)
- + Modify **\*def files**
- + Read **readme** file
- + **get results** in netcdf files (hist\*.nc)
- + **compare** to LES results if available

**Where is  
located 1D ?**

~/LMDZ20171130.trunk

1D/run.sh

OUTPUT\_COMMUNS

DOC

bin

CAS

DIAG

INPUT

RESU

compile  
lmdz1d.e

*bin contains  
compile script  
and different  
versions of  
executable file*

amma  
lmdz1d.def  
run.def  
readme  
forcing files  
arm\_cu  
LES.nc  
case\_e  
fire  
etc ...

*CAS contains .def  
files specific for  
each case and LES  
results to compare  
with LMDZ outputs*

dice\_bucket  
gabls4

*DIAG contains  
scripts to plot  
diagnostics*

DEF  
config.def  
gcm1d.def  
traceur.def  
PHYS  
physiq.def\_NPv3.2  
physiq.def\_NPv6.0.12  
etc...  
VERT  
L79  
L79.def  
L130  
L130.def  
etc...

*INPUT contains 3  
generic .def files and  
.def files specific for  
each physical  
package, vertical  
discretization*

all.pdf  
NPv3.2L39  
amma  
arm\_cu  
NPv6.0.12L79  
amma  
dice  
SAVE5438

*RESU directory  
contains results and  
is created at the first  
run*

modipsl

modeles

LMDZ5

libf

dyn3d

phylmd

dyn1d

lmdz1d.F90

**Code  
to modify**

There are two ways to run the model :

- Either in « **operationnal mode** » with ~1D/run.sh (several cases and physical packages)
- Or « **by hand** » in ~/RESU/Npxxx/case/ with compile.sh then lmdz1d.e

# How to run a case or compile after modifications : With run.sh

## Which case(s) ?

```
listecas="dice ihop arm_cu rico sandufast sanduref sanduslow fire  
toga ayotte twpice case_e amma " # testes
```

```
listecas="arm_cu rico sandufast fire twpice amma"
```

```
listecas="amma"
```

## Which physics ?

```
listedef="SP NPv3.1 NPv3.2 NPv4.12 NPv5.00 NPv5.10"
```

```
listedef="NPv5.00 "
```

```
listedef="NPv6.1 "
```

## Number of levels ?

```
case $DEF in  
  SP|NPV3.1|NPv3.2) L=39 ;;  
  NPv4.12) L=59 ;;  
  H2002) L=130  
  *) L=79  
esac
```

## Where are the results ?

In LMDZtesting/1D/RESU/NPv6.1L95/amma

All the files necessary to 1D run are copied + output files

lrwxrwxrwx	1	...	36	3	déc.	11:38	amma.nc	← forcings
-rw-r--r--	1	...	285452	3	déc.	11:38	histhf.nc	↙ Result files
-rw-r--r--	1	...	279088	3	déc.	11:38	hourly.nc	↘
-rw-r--r--	1	...	15292	3	déc.	11:38	limit.nc	
-rw-r--r--	1	...	652	3	déc.	11:38	lmdz1d.def	*def files
-rw-r--r--	1	...	4247	3	déc.	11:38	config.def	
-rw-r--r--	1	...	85	3	déc.	11:38	gcm.def	
-rw-r-----	1	...	3765	3	déc.	11:38	physiq.def	
-rwxr-xr--	1	...	692	3	déc.	11:38	run.def	
-rw-r--r--	1	...	42	3	déc.	11:38	traceur.def	
-rw-r--r--	1	...	7728	3	déc.	11:38	paramLMDZ_phy.nc	
-rw-r--r--	1	...	16532	3	déc.	11:38	startphy.nc	

### CAUTION !

You can modify \*def files in ~LMDZtesting/1D/RESU and quickly rerun the model because lmdz1d.e is in this directory. **BUT BE CAREFULL**

The « original » files are either under ~LMDZtesting/CAS or ~LMDZtesting/INPUT  
And will be replaced at each run of run.sh

# About 1D output files

Keep only histhf file with the maximum of data

phys_out_filekeys=	n	y	n	n	n	
phys_out_filenames=	hourly	histhf	day	histins	histLES	filehf
phys_out_filelevels=	5	10	10	10	10	0
phys_out_filetypes=	ave(X)	inst(X)	ave(X)	inst(X)	inst(X)	inst(X)
phys_out_filetimesteps=	1hr	1ts	1day	1hr	6hr	1ts

**To get all variables names :**

```
ncdump -h histhf.nc|grep long_name|sort
```

**To get the names of all temperature tendencies :**

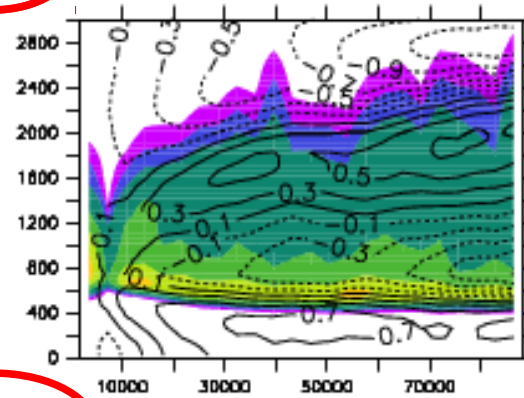
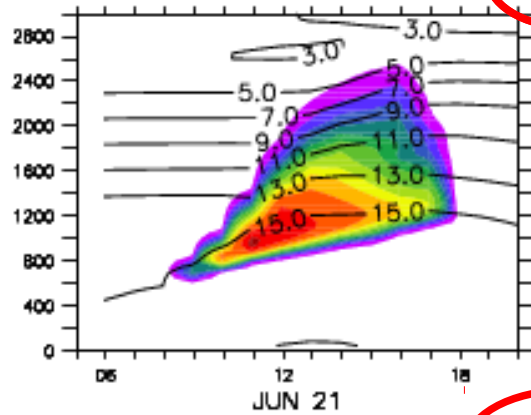
```
ncdump -h histhf.nc|grep long_name|grep dt
```

Arm\_cu

LES

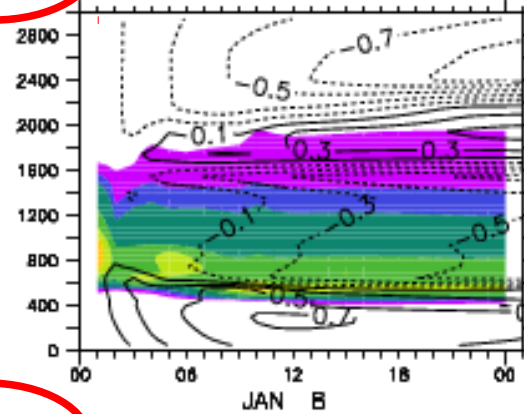
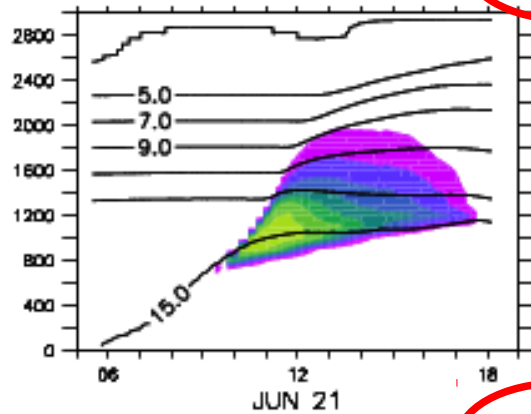
Rico

Z (m)



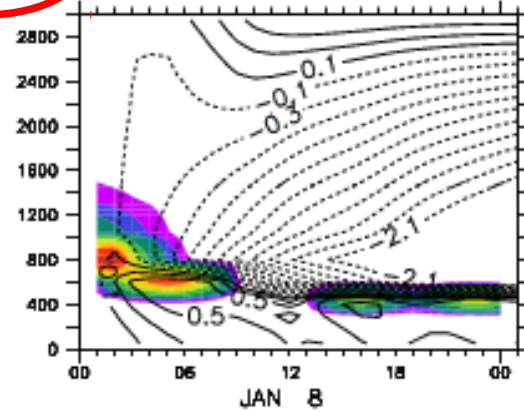
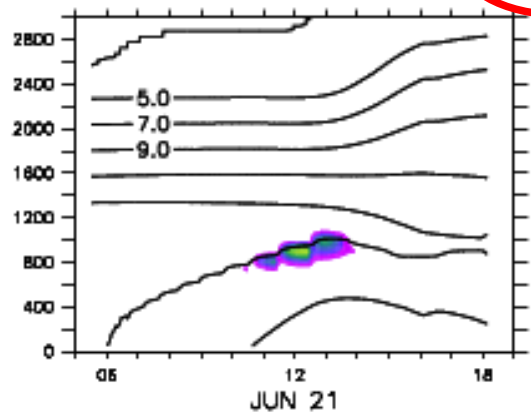
NPv3

Z (m)



SP

Z (m)



**In LMD, we used these cases to develop New Physics version.**

For 2 cases, Arm\_cu and Rico, we compare results of « standard physics » (CMIP3), « new physics » (CMIP5) and LES model.

Shade= cloud cover  
Contour= specific humidity (g/kg)



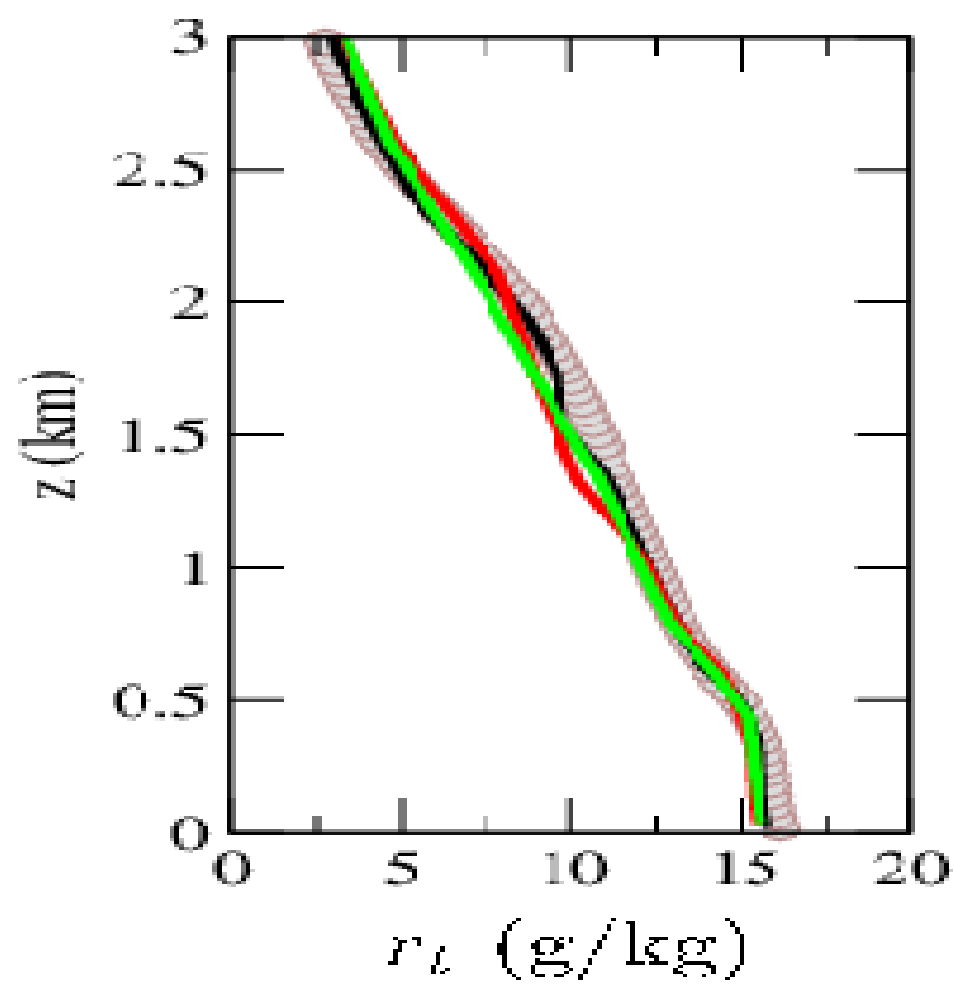
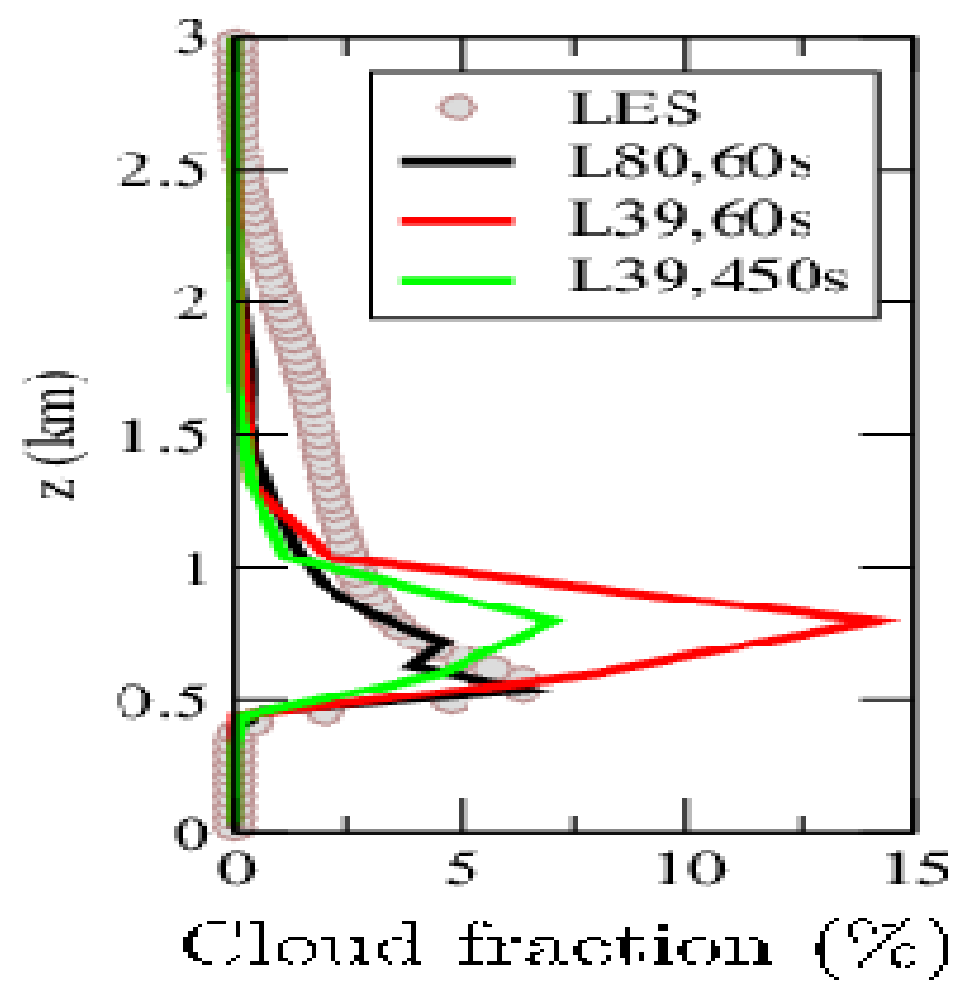
1 2 3 4 5 6 8 10 12 15 20 25 30



0.2 0.5 1 3 5 7 10 15 20

**Rico case :**

Sensitivity Tests to vertical discretization and time step



Thank you !!