

The physical parametrizations in LMDZ

LMDZ Team

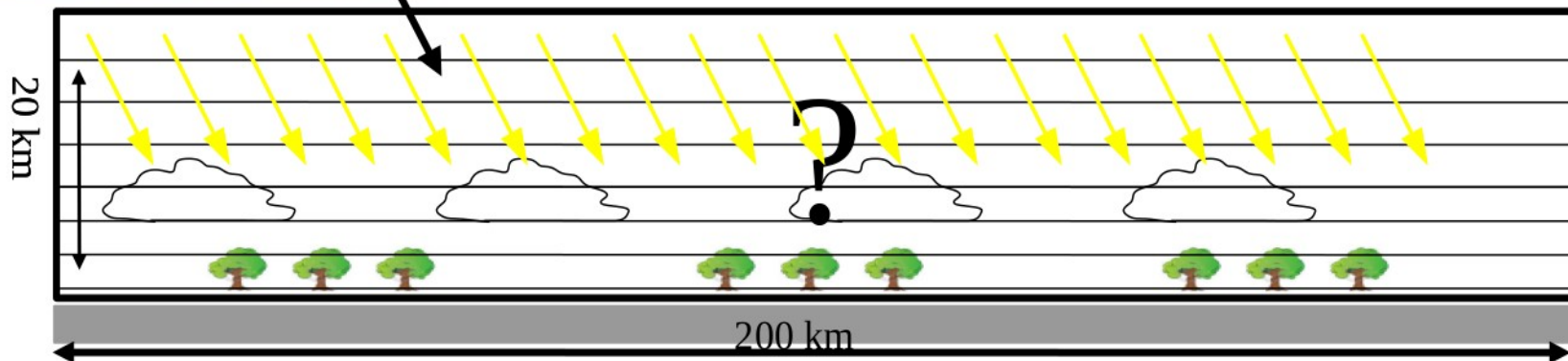
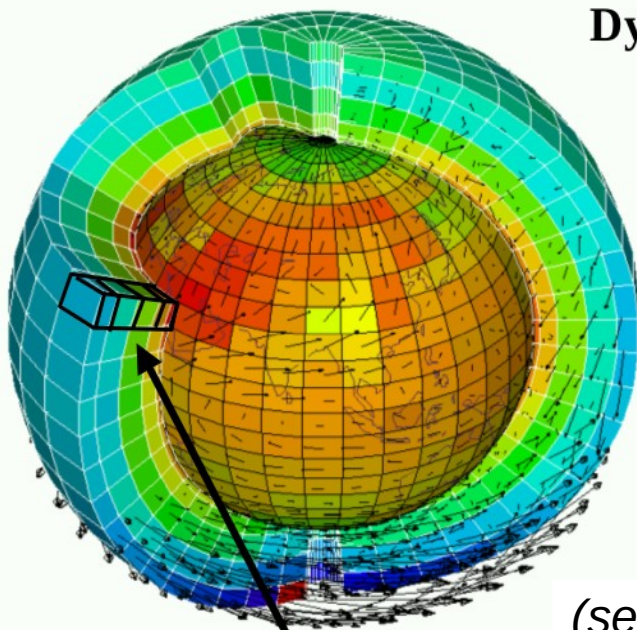
Laboratoire de Météorologie Dynamique
December 2017

Quick reminder : general equations

Dynamical core : primitive equations discretized on the sphere

- Mass conservation
$$D\rho/Dt + \rho \operatorname{div}\underline{U} = 0$$
- Potential temperature conservation
$$D\theta/Dt = Q/C_p (p_0/p)^\kappa$$
- Momentum conservation
$$D\underline{U}/Dt + (1/\rho) \operatorname{grad}p - \underline{g} + 2 \underline{\Omega} \wedge \underline{U} = \underline{F}$$
- Secondary components conservation
$$Dq/Dt = Sq$$

(see yesterday's presentation by F. Hourdin)



Atmospheric GCM equations

Primitive equations in pressure coordinates

$$\left\{ \begin{array}{l}
 \text{Momentum equation :} \\
 \partial_t \vec{v} + \boxed{(\vec{v} \cdot \vec{\nabla}_p) \vec{v}} + \omega \partial_p \vec{v} + \boxed{f \vec{k} \times \vec{v}} + \boxed{\vec{\nabla}_p \Phi} = \boxed{\vec{S}_v} \\
 \text{Continuity equation :} \\
 \vec{\nabla}_p \cdot \vec{v} + \partial_p \omega = 0 \\
 \text{Component conservation :} \\
 \partial_t q + \vec{v} \cdot \vec{\nabla}_p q + \omega \partial_p q = \boxed{S_q} \\
 \text{Thermodynamic equation :} \\
 \partial_t \theta + \vec{v} \cdot \vec{\nabla}_p \theta + \omega \partial_p \theta = \frac{\theta}{c_p T} \boxed{\dot{Q}_{net}}
 \end{array} \right.$$

Sources

$\Phi = gz$ geopotential

$\omega = \partial_t p$ vert. velocity

q = specific humidity

\dot{Q}_{net} = heating rate from all diabatic sources

\vec{S}_v , S_q and \dot{Q}_{net} : source terms determined by the **physical parametrizations** and the **radiative transfer scheme** :

- planetary boundary layer, shallow and deep convection
- scattering and absorption by cloud droplets and crystals
- drag due to topography...

Model tendencies

The integration of a given prognostic variable X ($T, \vec{v}(u, v, w), p, \rho, q_{vap}$) can be written as :

$$X_{t+\Delta t} = X_t + \left(\frac{\partial X}{\partial t} \right)_{\text{dyn}} \Delta t \text{ (dynamical core)} \quad (1)$$

$$+ \left(\frac{\partial X}{\partial t} \right)_{\text{param}} \Delta t \text{ (parameterizations)} \quad (2)$$

Basic facts about parametrizations I

- Each parametrization : (1) works almost independently of the others ;
 (2) depends on vertical profiles of u , v , w , T , q and on some interface variables with the other parametrizations ; (3) ignores the spatial heterogeneities associated with the other processes (except for some processes in the deep convection scheme).
- The total tendency due to sub-grid processes is the sum of the tendencies due to each process :

$$\begin{aligned}
 S_T = (\partial_t T)_\varphi &= (\partial_t T)_{\text{eva}} + (\partial_t T)_{\text{lsc}} + (\partial_t T)_{\text{diff turb}} + (\partial_t T)_{\text{conv}} \\
 &\quad + (\partial_t T)_{\text{wk}} + (\partial_t T)_{\text{Th}} + (\partial_t T)_{\text{ajs}} \\
 &\quad + (\partial_t T)_{\text{rad}} + (\partial_t T)_{\text{oro}} + (\partial_t T)_{\text{dissip}}
 \end{aligned}$$

In the model, the total tendency of T for example is $\partial_t T_{\text{dyn}} + \partial_t T_{\text{param}}$
 $= \text{dtdyn} + \text{dtphy}$, where :

$$\begin{aligned}
 \text{dtphy} = & \text{dteva} + \text{dtlsc} + \text{dtvdf} + \text{dtcon} + \\
 & \text{dtwak} + \text{dtthe} + \text{dtajs} + \\
 & (\text{dtswr} + \text{dtlwr}) + (\text{dtoro} + \text{dtlif}) + (\text{dtdis} + \text{dtec})
 \end{aligned}$$

Basic facts about parametrizations II

- Similarly, the total tendency of a given tracer q writes :

$$S_q = (\partial_t q)_\varphi = (\partial_t q)_{\text{eva}} + (\partial_t q)_{\text{lsc}} + (\partial_t q)_{\text{diff turb}} + (\partial_t q)_{\text{conv}} \\ + (\partial_t q)_{\text{wk}} + (\partial_t q)_{\text{Th}} + (\partial_t q)_{\text{ajs}}$$

In the model, the total tendency of q is therefore

$\partial_t q_{\text{dyn}} + \partial_t q_{\text{param}} = \text{dqdyn} + \text{dqphy}$, where :

$\text{dqphy} = \text{dqeva} + \text{dqlsc} + \text{dqvdf} + \text{dqcon} + \text{dqwak} + \text{dqthe} + \text{dqajs}$

physiq_mod.F90 structure - I

Initialization (once) : *conf_phys*, *phyetat0*,
phys_output_open

Beginning *change_srf_frac*, *solarlong*

Cloud water evap. *reevap*

Vertical diffusion (turbulent mixing) *pbl_surface*

Deep convection *conflx* (Tiedtke) or *concul* (Emanuel)

Deep convection clouds *clouds_gno*

Density currents (wakes) *calwake*

Strato-cumulus *stratocu_if*

Thermal plumes *calltherm* and *ajsec* (sec = dry)

Large scale clouds *calcratqs*

Large scale and cumulus condensation *fisrtilp*

Diagnostic clouds for Tiedtke *diagcld1*

Aerosols *readaerosol_optic*

Cloud optical parameters *newmicro* or *nuage*

Radiative processes *radlwsu*

In blue : subroutines and instructions modifying state
variables

physiq_mod.F90 structure - II

Orographic processes : drag *drag_noro_strato* or
drag_noro

Orographic processes : lift *lift_noro_strato* or
lift_noro

Orographic processes : Gravity Waves *hines_gwd* or
GWD_rando

**Axial components of angular momentum and
mountain torque** : *aaam_bud*

Cosp simulator *phys_cosp*

Tracers *phytrac*

Tracers off-line *phystokenc*

Water and energy transport *transp*

Outputs

Statistics

Output of final state (for restart) *phyredem*

Turbulent diffusion

- Turbulent diffusion or "**turbulent mixing**" : transport by small random movements. Similar to molecular diffusion.

$$Dq/Dt = S_q \quad \text{où} \quad S_q = \frac{\partial}{\partial z} \left(K_z \frac{\partial q}{\partial z} \right)$$

- **Prandtl mixing length** : $K_z = l |w|$
 l : characteristic length of the small movements
 w : characteristic velocity
- **Turbulent kinetic energy (TKE)** : $K_z = l \sqrt{e}$

$$De/Dt = f(dU/dz, d\theta/dz, e, \dots)$$

$$Dl/Dt = \dots$$

Turbulent diffusion : numerics

Process : Turbulent mixing of moisture (q in kg/kg) and potential enthalpy ($H = C_p \theta$).

$$\left\{ \begin{array}{l} \frac{dq}{dt} = \partial_z \phi_q \\ \phi_q = K_z \partial_z q \\ \phi_q|_{\text{srf}} = -\text{Evap} \end{array} \right\} \left\{ \begin{array}{l} \frac{dH}{dt} = \partial_z \phi_\theta \\ \phi_\theta = K_z \partial_z H \\ \phi_\theta|_{\text{srf}} = \phi_{\text{sens}} \left(\frac{p_0}{p_{\text{srf}}} \right)^\kappa \end{array} \right. \quad \begin{array}{l} \text{(Fluxes} \\ \text{positive} \\ \text{downward)} \end{array} \quad (3)$$

Spatial discretization : (moisture)

$$\left\{ \begin{array}{l} m_i \partial_t q_i = \phi_{q,i+1} - \phi_{q,i} \\ \phi_{q,i} = K_i (q_i - q_{i-1}) \\ \phi_{q,1} = -\text{Evap} \end{array} \right. \quad (4)$$

Implicit scheme, yields for the first atmospheric layer :

$$\begin{aligned} q_{1,t+\delta t} &= A + B \phi_{q,1} \delta t \\ \phi_{q,1} &= K_1 (q_{1,t+\delta t} - q_{\text{srf}}) \end{aligned} \quad (5)$$

A and B are coefficient resulting from solving Eq. (4) over the whole atmosphere.

Eqs. (5) are the mixed boundary conditions for the sub-surface model.

Turbulent diffusion : numerics

Process : Turbulent mixing of moisture (q in kg/kg) and potential enthalpy ($H = C_p \theta$).

$$\left\{ \begin{array}{l} \frac{dq}{dt} = \partial_z \phi_q \\ \phi_q = K_z \partial_z q \\ \phi_q|_{\text{srf}} = -\text{Evap} \end{array} \right\} \left\{ \begin{array}{l} \frac{dH}{dt} = \partial_z \phi_\theta \\ \phi_\theta = K_z \partial_z H \\ \phi_\theta|_{\text{srf}} = \phi_{\text{sens}} \left(\frac{p_0}{p_{\text{srf}}} \right)^\kappa \end{array} \right. \quad \begin{array}{l} \text{(Fluxes} \\ \text{positive} \\ \text{downward)} \end{array} \quad (3)$$

Spatial discretization : (moisture)

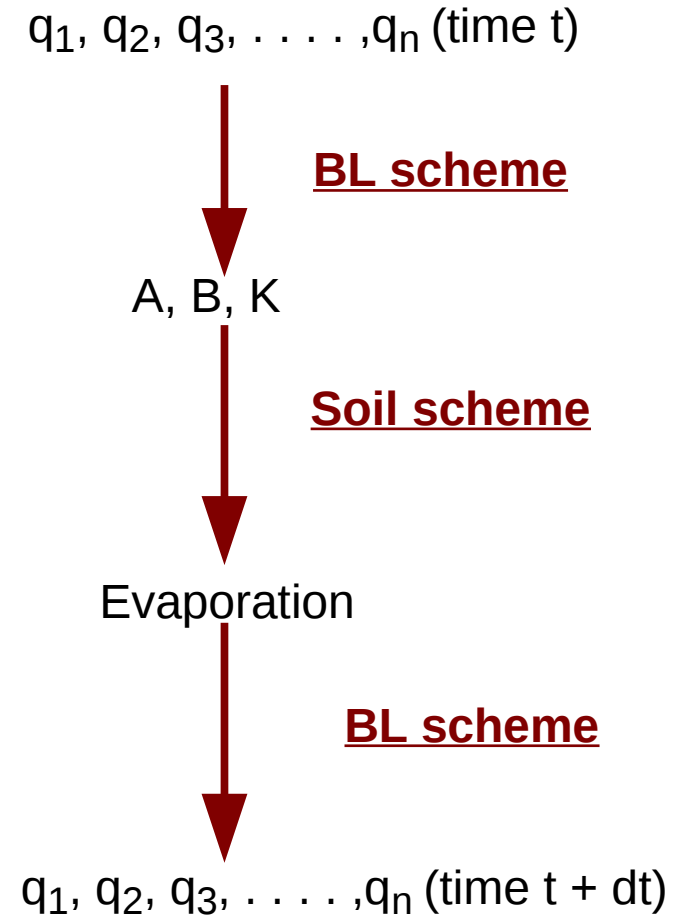
$$\left\{ \begin{array}{l} m_i \partial_t q_i = \phi_{q,i+1} - \phi_{q,i} \\ \phi_{q,i} = K_i (q_i - q_{i-1}) \\ \phi_{q,1} = -\text{Evap} \end{array} \right. \quad (4)$$

Implicit scheme, yields for the first atmospheric layer :

$$\begin{aligned} q_{1,t+\delta t} &= A + B \phi_{q,1} \delta t \\ \phi_{q,1} &= K_1 (q_{1,t+\delta t} - q_{\text{srf}}) \end{aligned} \quad (5)$$

A and B are coefficient resulting from solving Eq. (4) over the whole atmosphere.

Eqs. (5) are the mixed boundary conditions for the sub-surface model.



Vertical diffusion

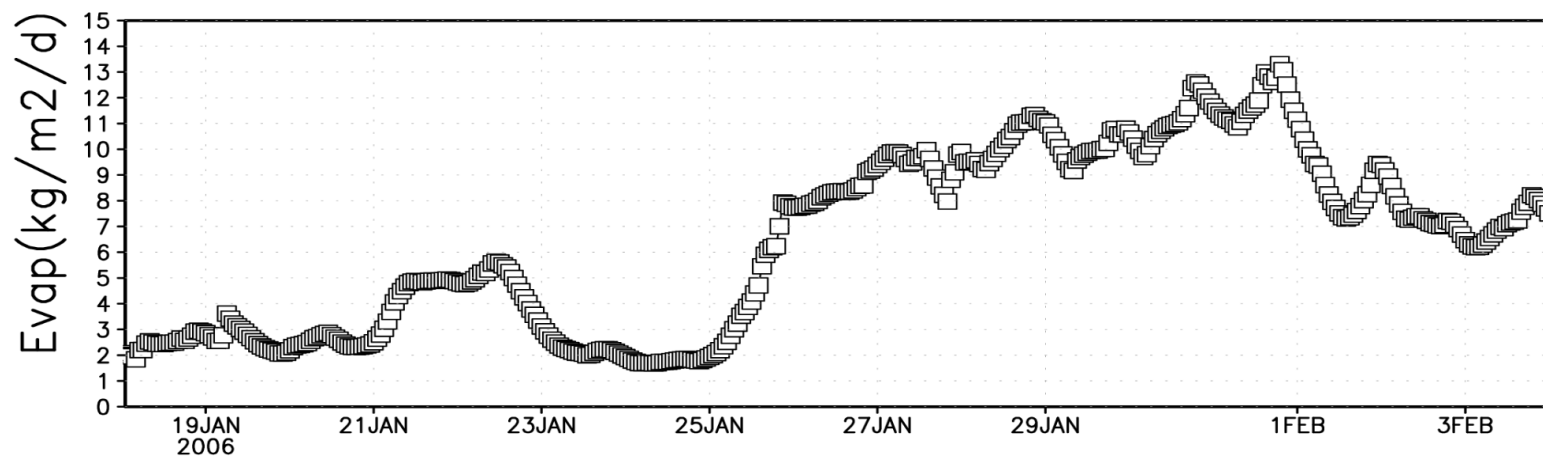
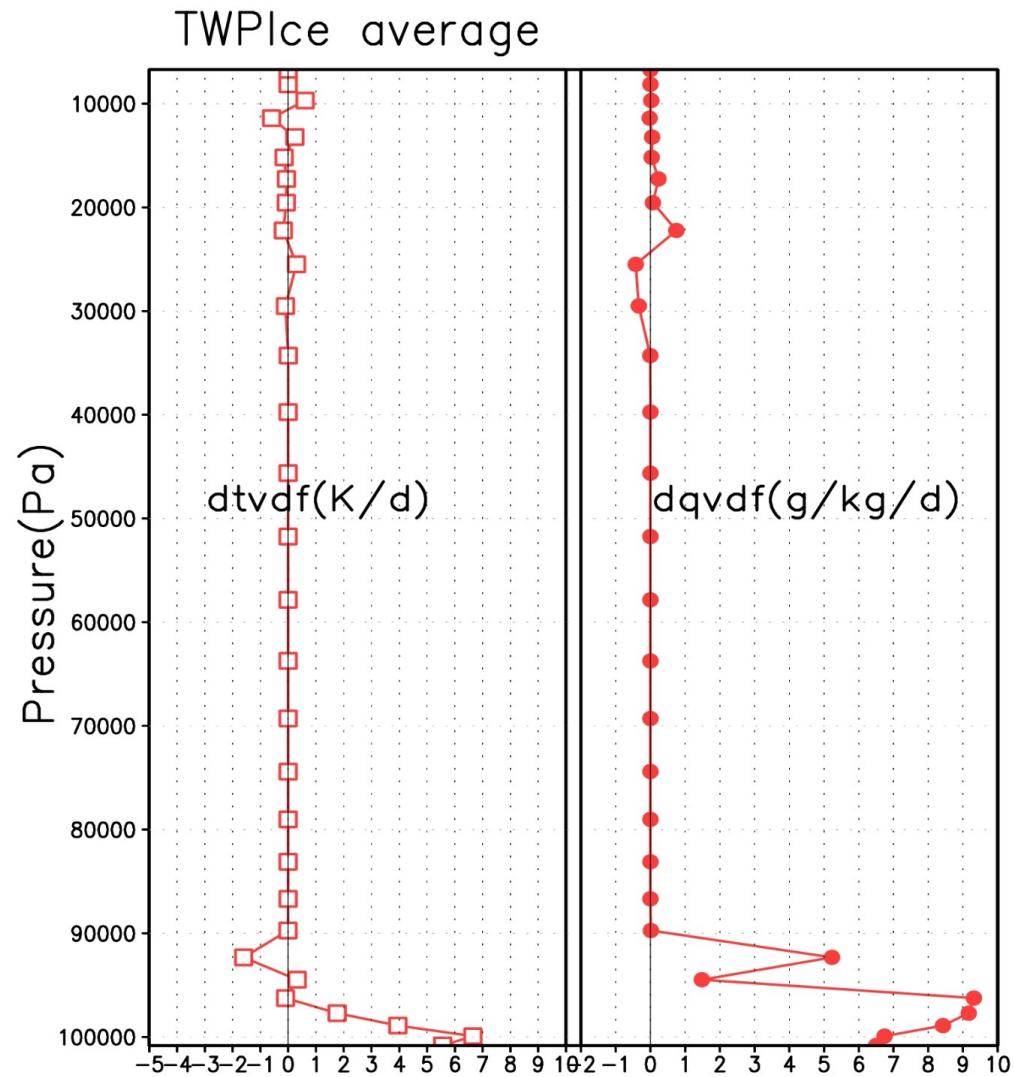
Subroutine : pbl_surface

Tendencies :

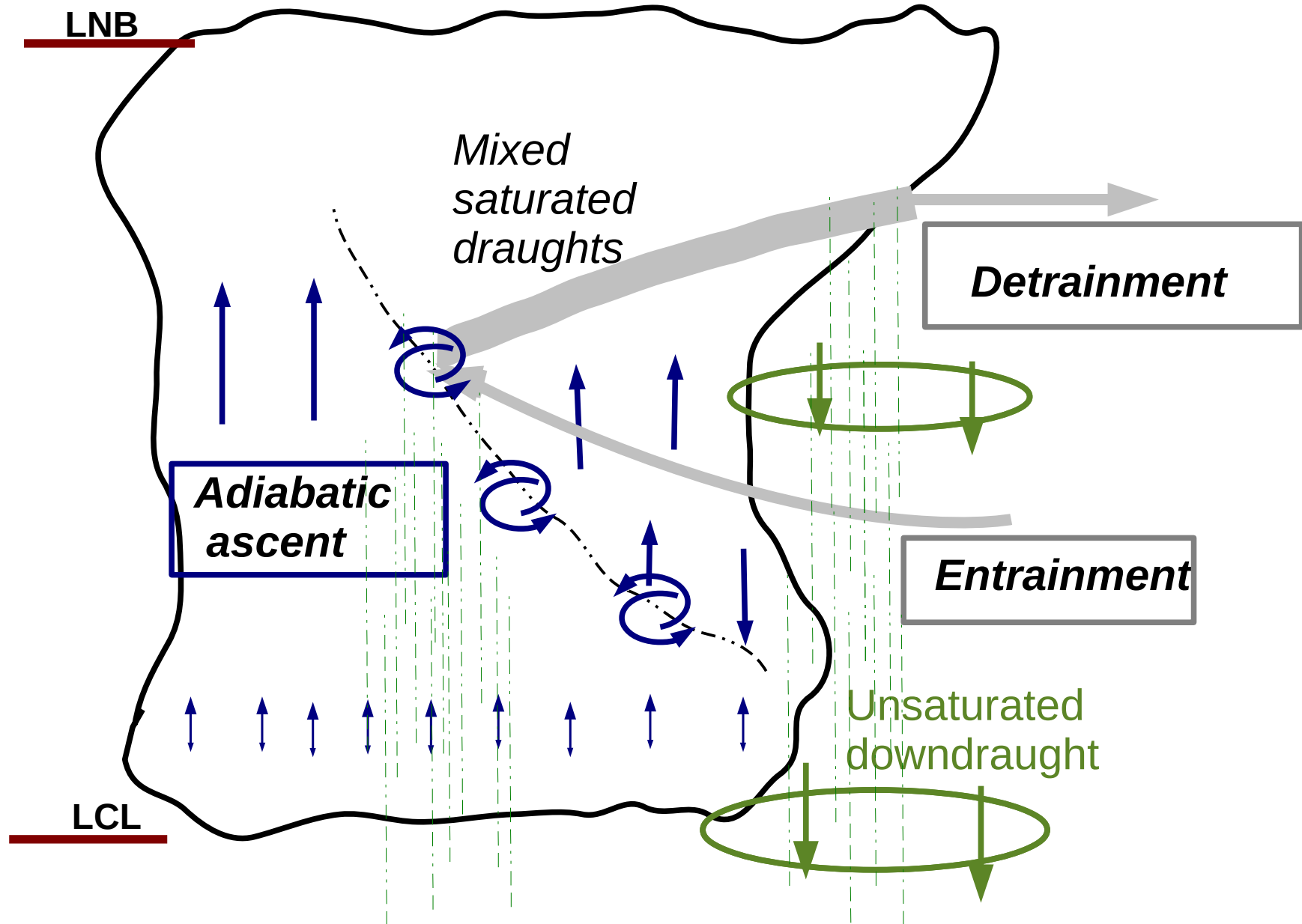
dtvdf, dqvdf, duvdf, dvvdf

Other variables

- sens : sensible heat flux at the surface (positive upward)
- evap : water vapour flux at the surface (positive upward)
- flat : latent heat flux at the surface (positive downward)
- taux, tauy : wind stress at the surface



Emanuel scheme



Deep convection

Subroutine : concvl

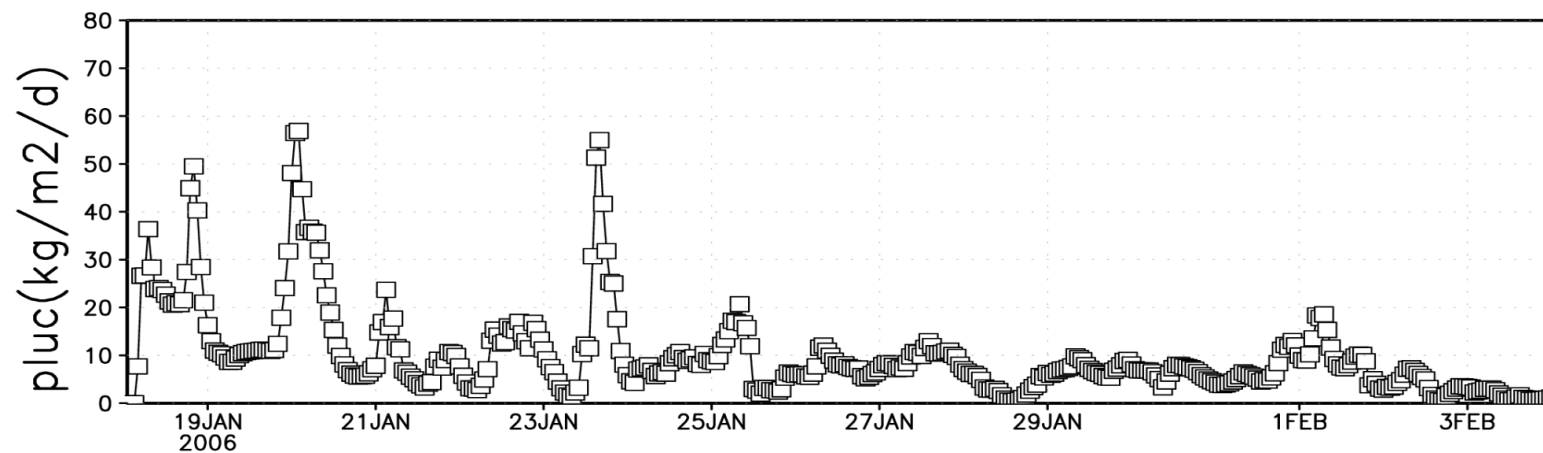
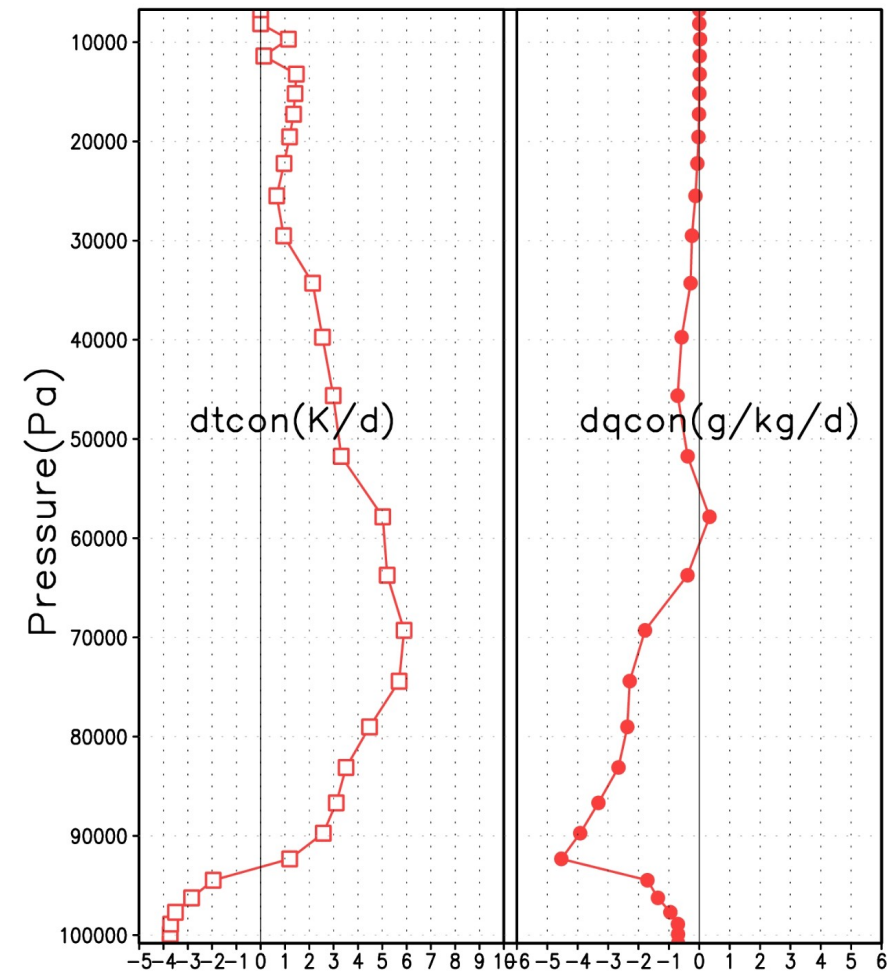
Tendencies :

dtcon, dqcon, ducon, dvcon

Other variables

- pluc : convective precipitation at the surface
- ftd : temperature tendency due to the sole unsaturated downdraughts
- fqd : moisture tendency due to the sole unsaturated downdraughts
- clwcon : condensed water of convective clouds
("in cloud" condensed water content)
- Ma : mass flux of the adiabatic ascent
- upwd : mass flux of the saturated updraughts
- dnwd : mass flux of the saturated downdraughts
- dnwd0 : mass flux of the unsaturated downdraught (precipitating downdraught)
- pr_con_l : vertical profile of convective liquid precipitation
- pr_con_i : vertical profile of convective ice precipitation

TWPlce average



Deep convection

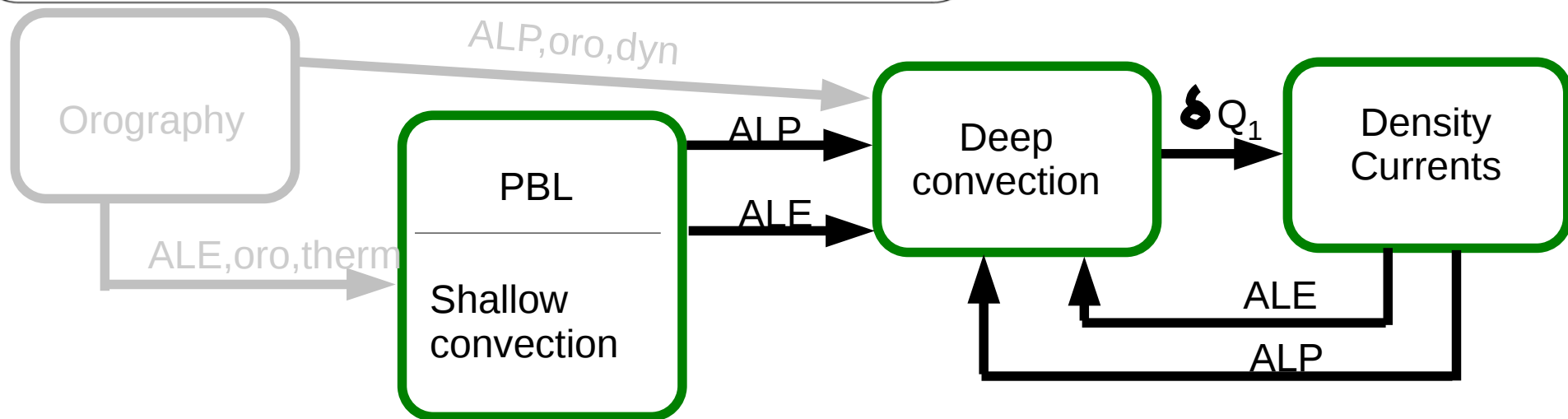
Subroutine : conecv1

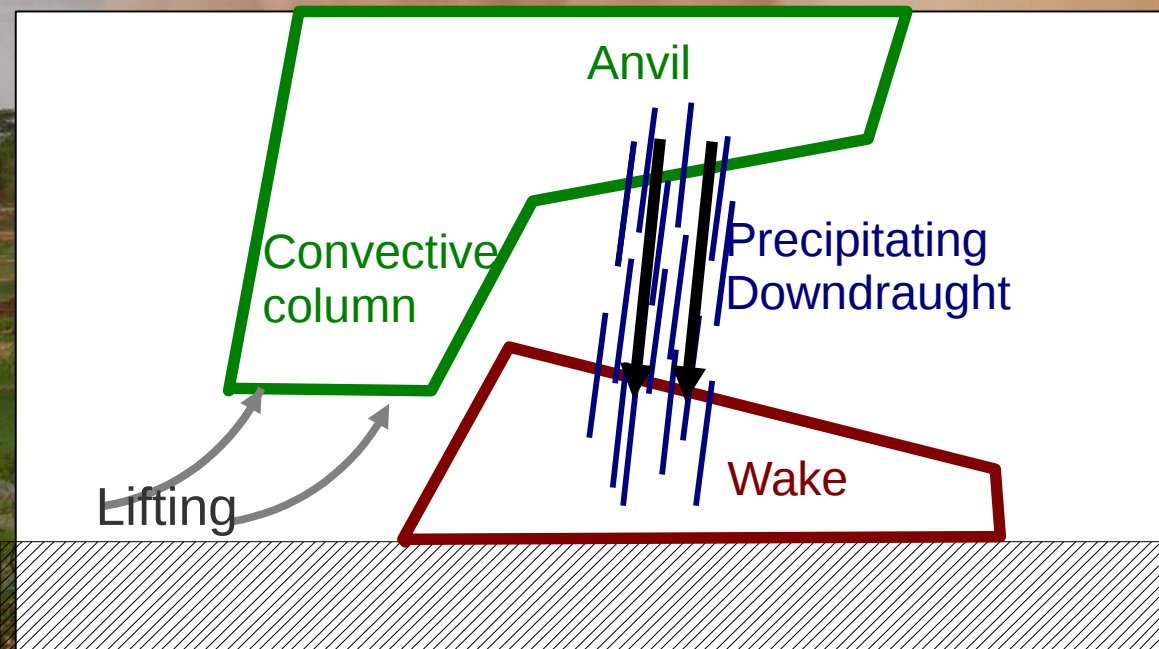
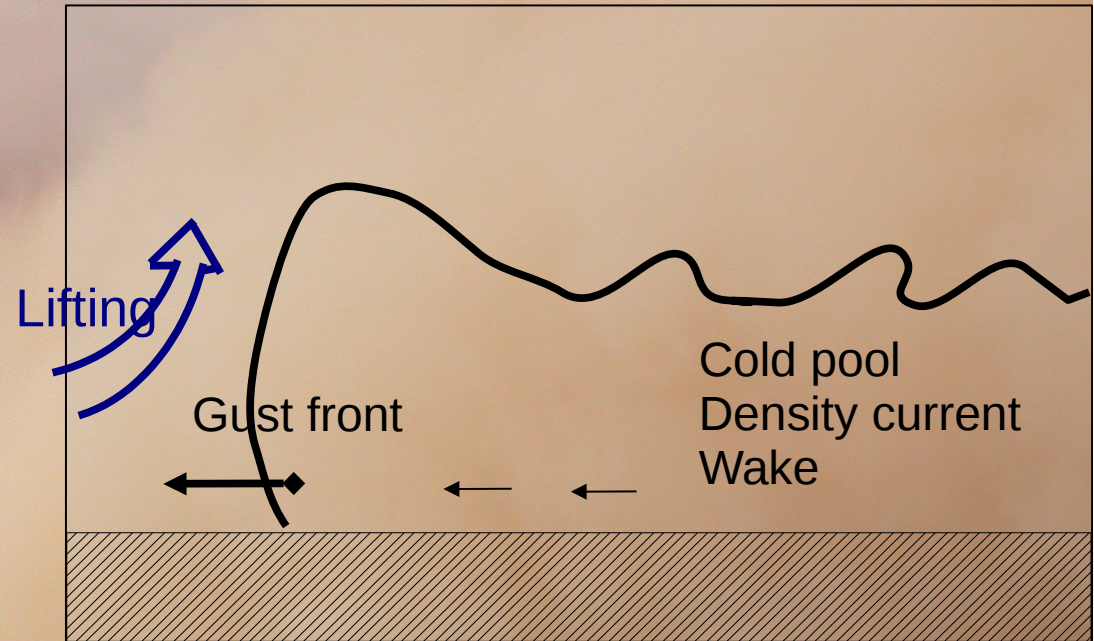
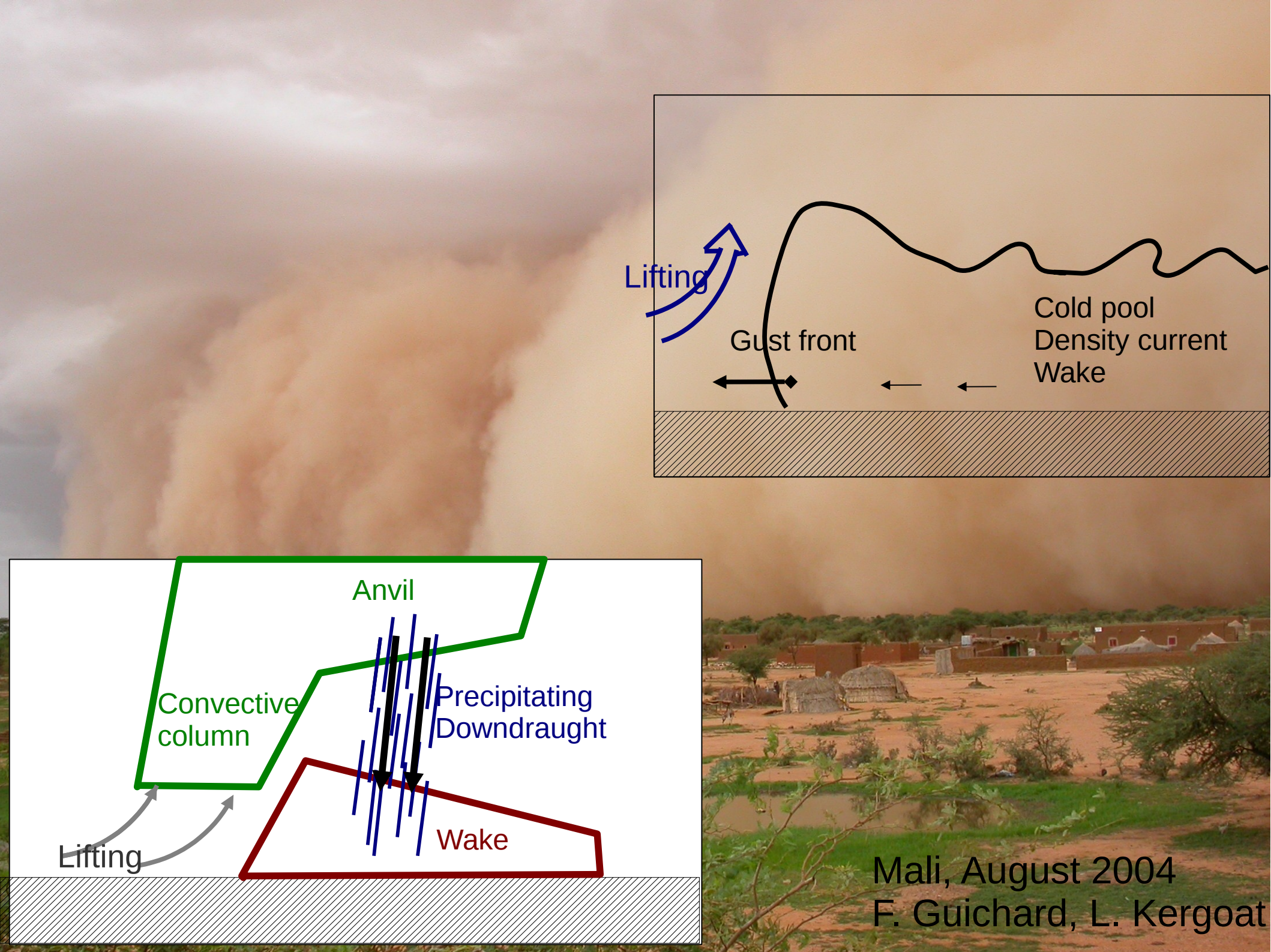
Tendencies :

dtcon, dqcon, ducon, dvcon

Other variables

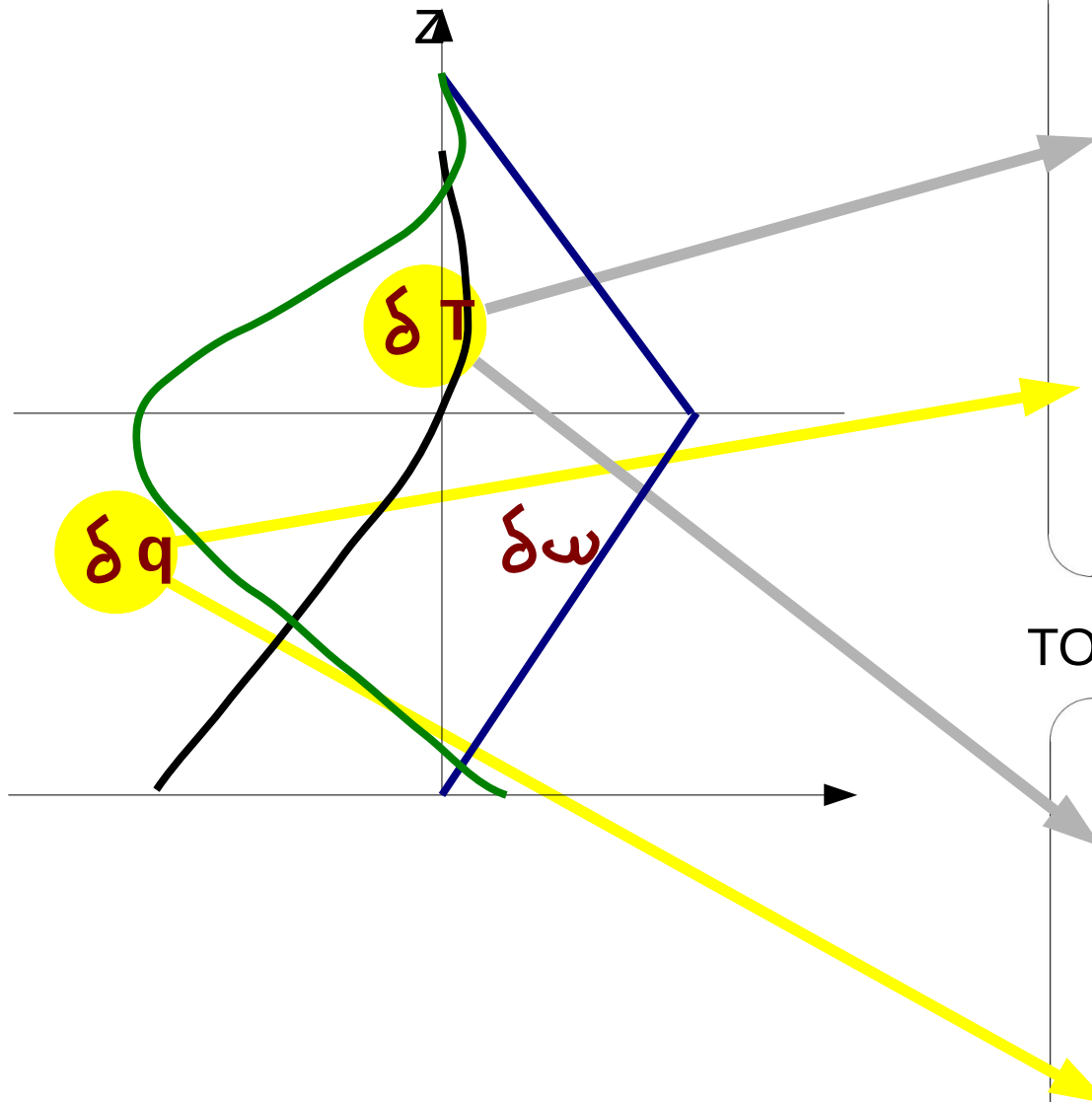
- pluc : convective precipitation at the surface
- ftd : temperature tendency due to the sole unsaturated downdraughts
- fqd : moisture tendency due to the sole unsaturated downdraughts
- clwcon : condensed water of convective clouds
("in cloud" condensed water content)
- Ma : mass flux of the adiabatic ascent
- upwd : mass flux of the saturated updraughts
- dnwd : mass flux of the saturated downdraughts
- dnwd0 : mass flux of the unsaturated downdraught (precipitating downdraught)
- pr_con_l : vertical profile of convective liquid precipitation
- pr_con_i : vertical profile of convective ice precipitation



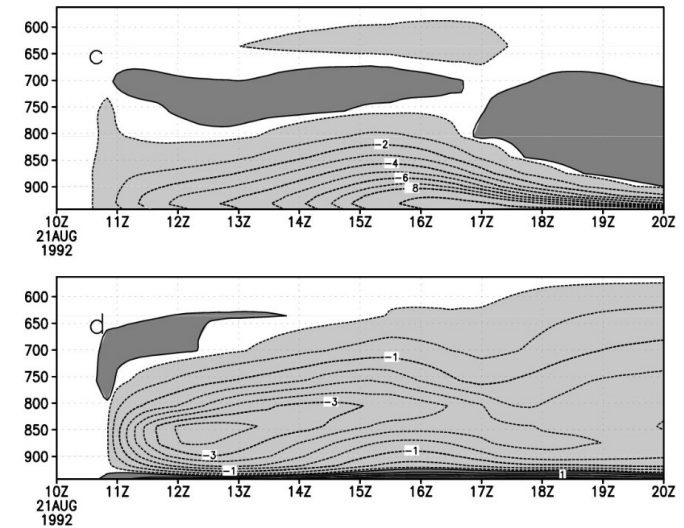


Mali, August 2004
F. Guichard, L. Kergoat

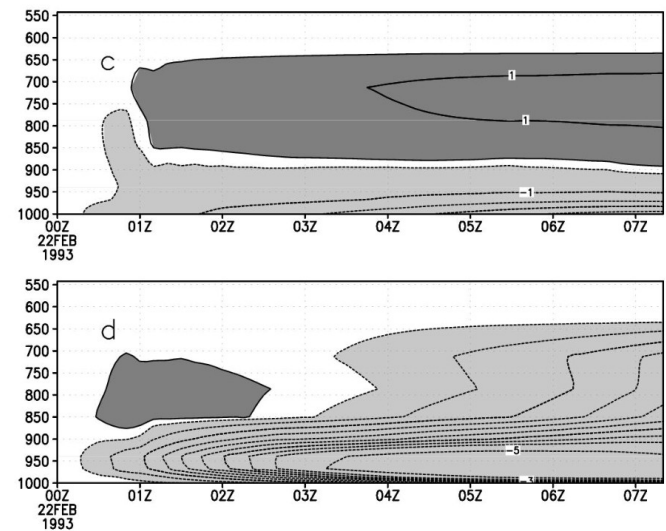
Simulated wake properties



HAPEX92: 21 Aug 1992 squall line case



TOGA-COARE: 22 Feb 1993 squall line case



Cold pools (wakes)

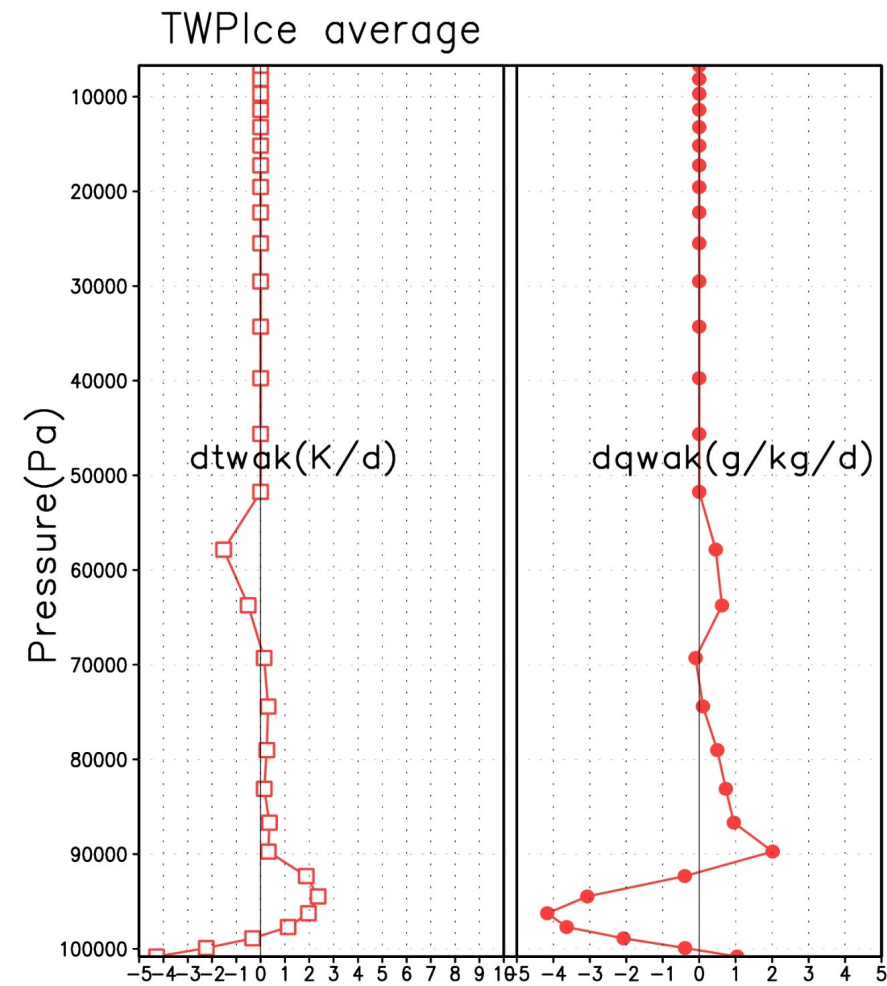
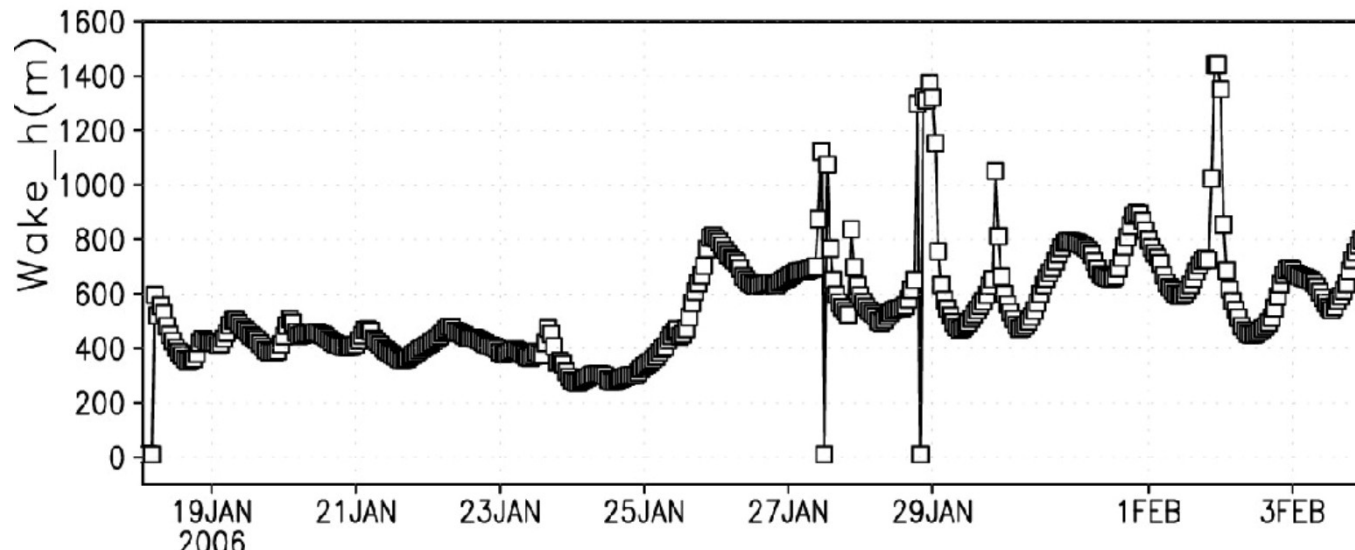
Subroutine : calwake

Tendencies :

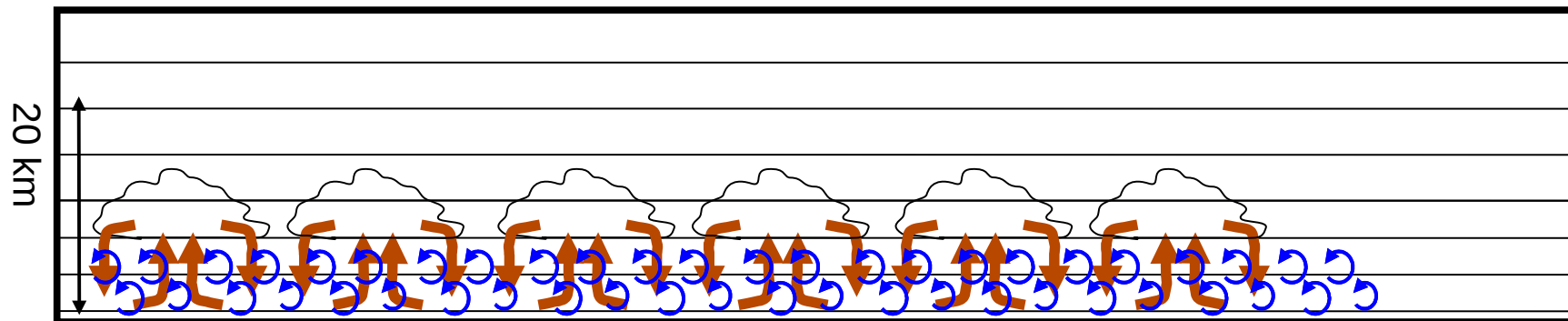
dtwak, dqwak

Other variables

- Alp_wk : lifting power due to cold pools
- Ale_wk : lifting energy due to cold pools
- wake_s : fractional area of cold pools
- wake_h : cold pool height
- wape : WAKE Potential Energy
- wake_deltat : vertical profile of temperature difference $T_w - T_x$
- wake_deltaq : vertical profile of humidity difference $q_w - q_x$
- wake_omg : vertical profile of vertical velocity difference $\omega_w - \omega_x$



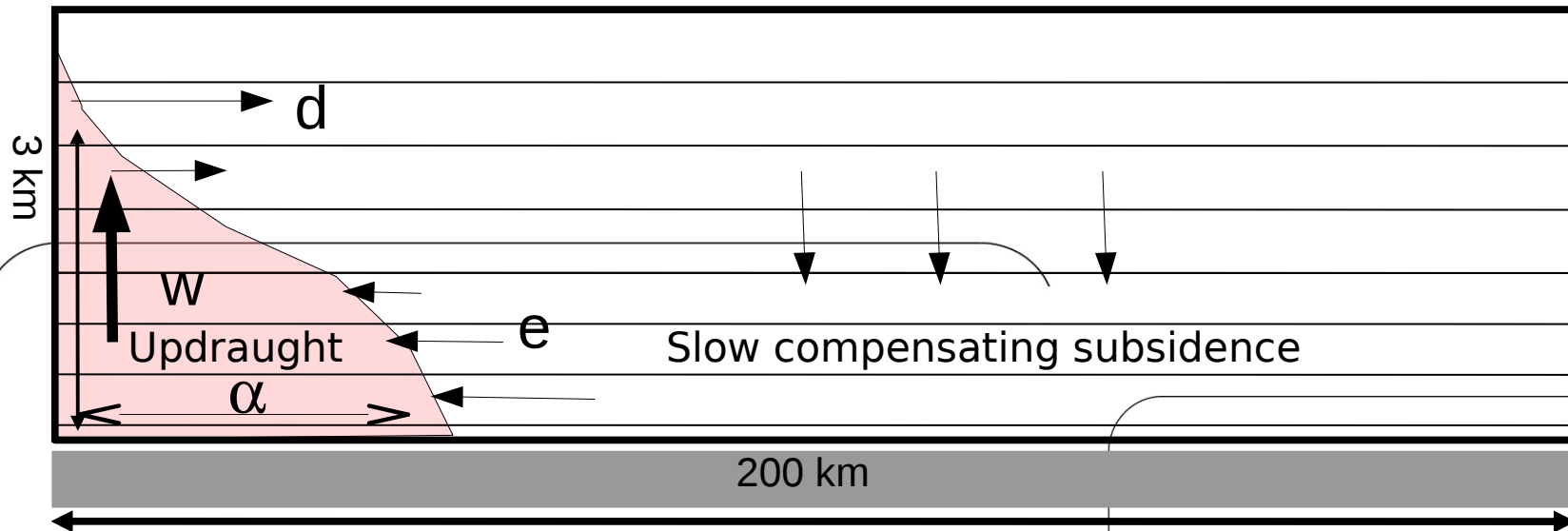
In a model column there are structures
of boundary layer scale



“The Thermal Model”:

Each column is split in two parts:
Ascending air from the surface and
subsiding air around it.

The model represents a mean
plume (the thermal) and a mean
cloud.



Internal variables of the parametrization :

- w = mean vertical velocity of ascending plumes
- α = fractionnal area covered by the updraughts
- e = lateral input rate of air into the plume (entrainment)
- d = output rate of air from the plume (detrainment)
- q_a = concentration of constituent q in the updraughts

Source term for the explicit equations :

$$S_q = -\frac{1}{\rho} \frac{\partial}{\partial z} \overline{\rho w' q'} = \frac{1}{\rho} \frac{\partial}{\partial z} \left[\rho K_z \frac{\partial q}{\partial z} \right] - \frac{1}{\rho} \frac{\partial}{\partial z} [f(q_a - q)]$$

Turbulent Diffusion

Transport by the thermal plume model

- Mass conservation

$$\frac{\partial f}{\partial z} = e - d \quad \text{where } f = \alpha \rho w$$

- Mass conservation of constituent q

$$\frac{\partial f q_a}{\partial z} = e q - d q_a$$

- Equation of movement

$$\frac{\partial f w}{\partial z} = -d w + \alpha \rho B$$

- where B is the buoyancy :

$$B = g \frac{\theta_{va} - \theta_v}{\theta_v}$$

- and the complex part lies in the expression of e and d :

$$e = f \max \left(0, \frac{\beta}{1+\beta} (a_1 \frac{B}{w^2} - b) \right)$$

$$d = \dots$$

Etc ...

Thermals and dry adjustment

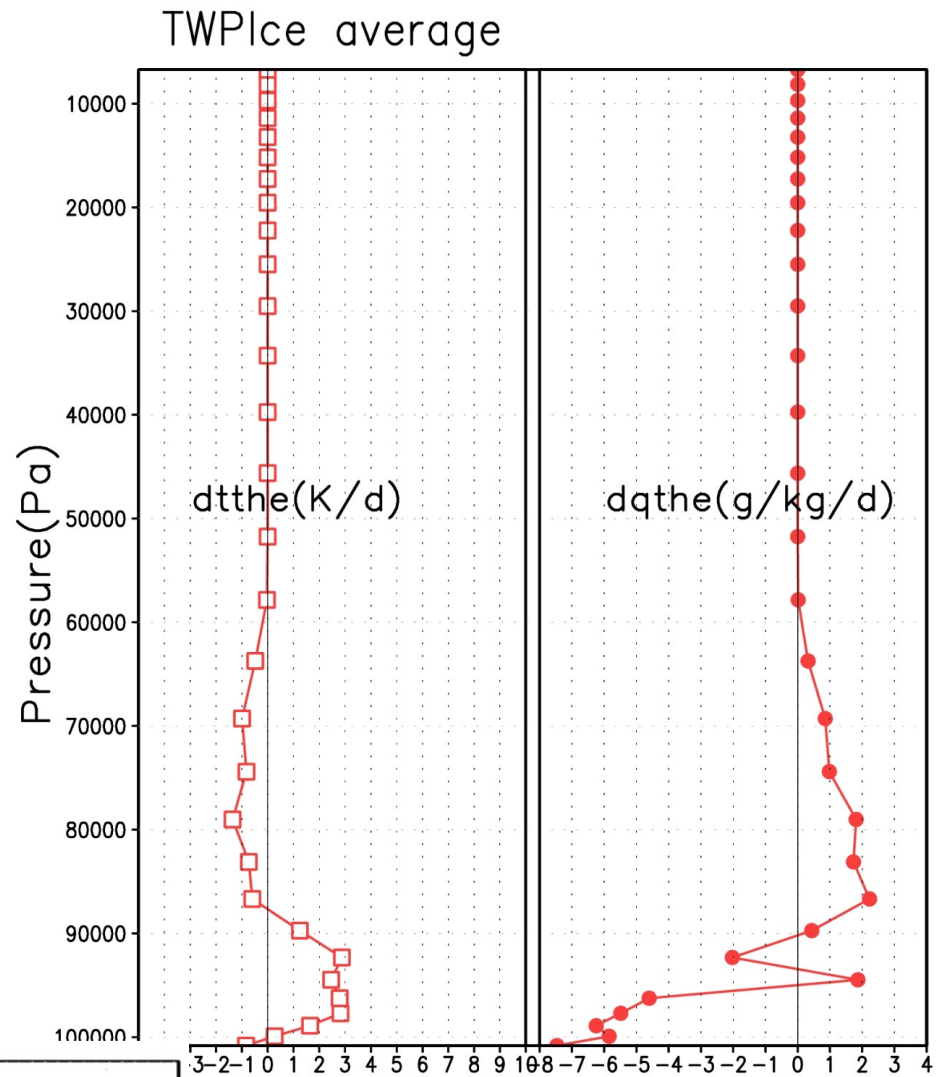
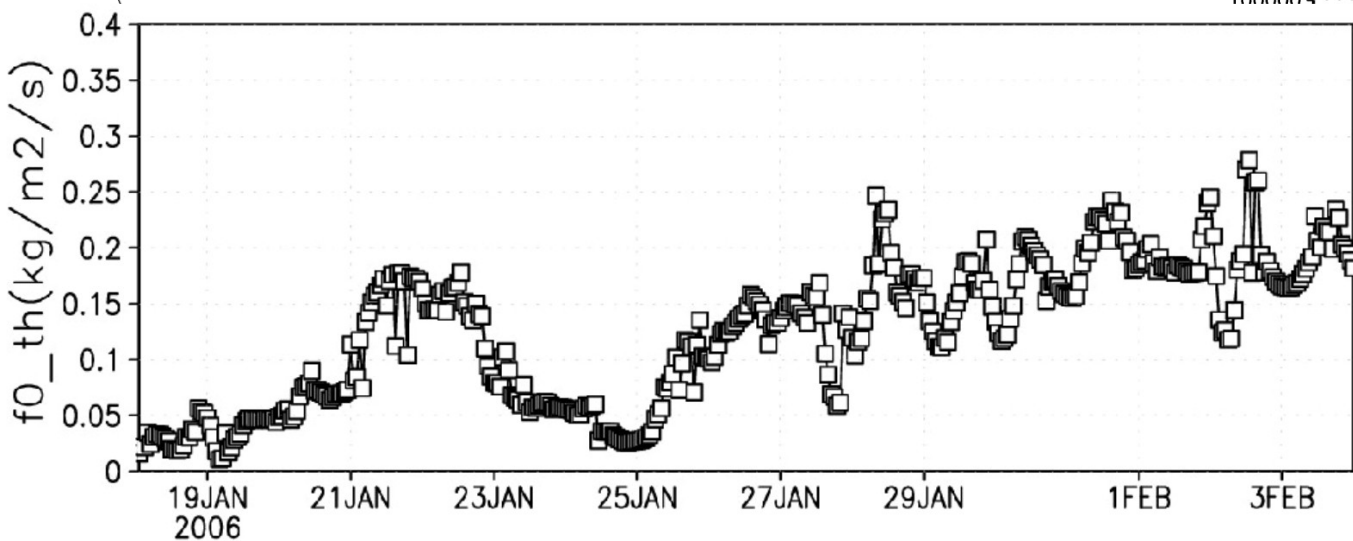
Subroutine : calltherm

Tendencies :

dtthe, dqthe, duthe, dvthe

Other variables

- dtajs : temperature tendency due to the sole dry adjustment
- dqajs : humidity tendency due to the sole dry adjustment
- a_th : fractional area of thermal plumes
- d_th : detrainment
- e_th : entrainment
- f_th : mass flux
- w_th : vertical velocity in the thermal plume (m/s, positive upward)
- q_th : total water content in the thermal plume
- zmax_th : altitude of the top of the thermal plume (m)



Large scale condensation (evap & lsc)

Subroutines : reevap & firtilp

Tendencies :

dteva, dqeva : tendencies due to cloud water evaporation

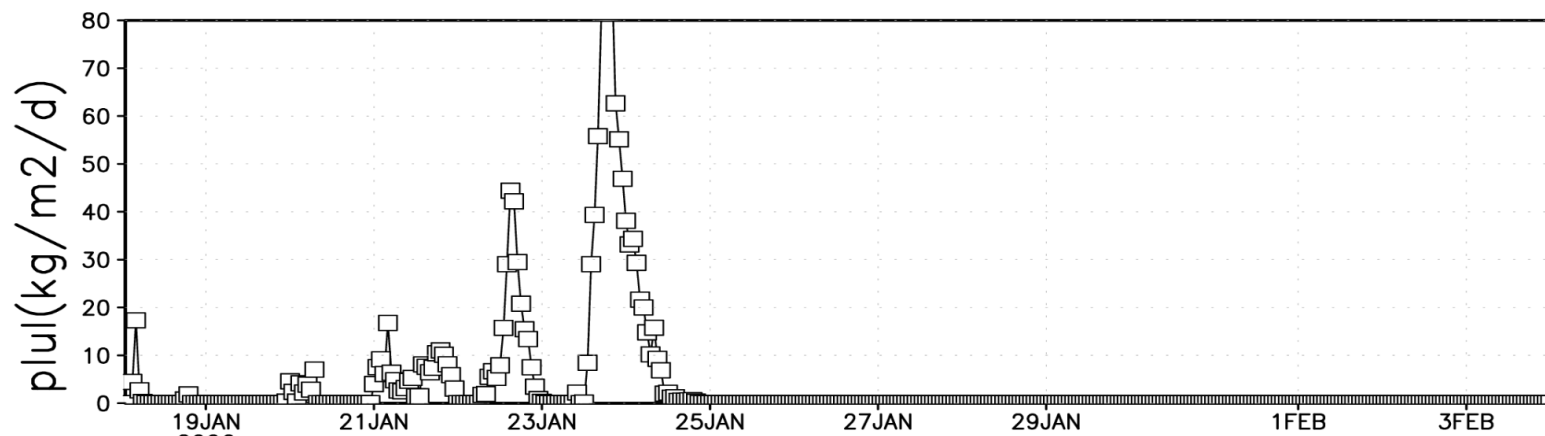
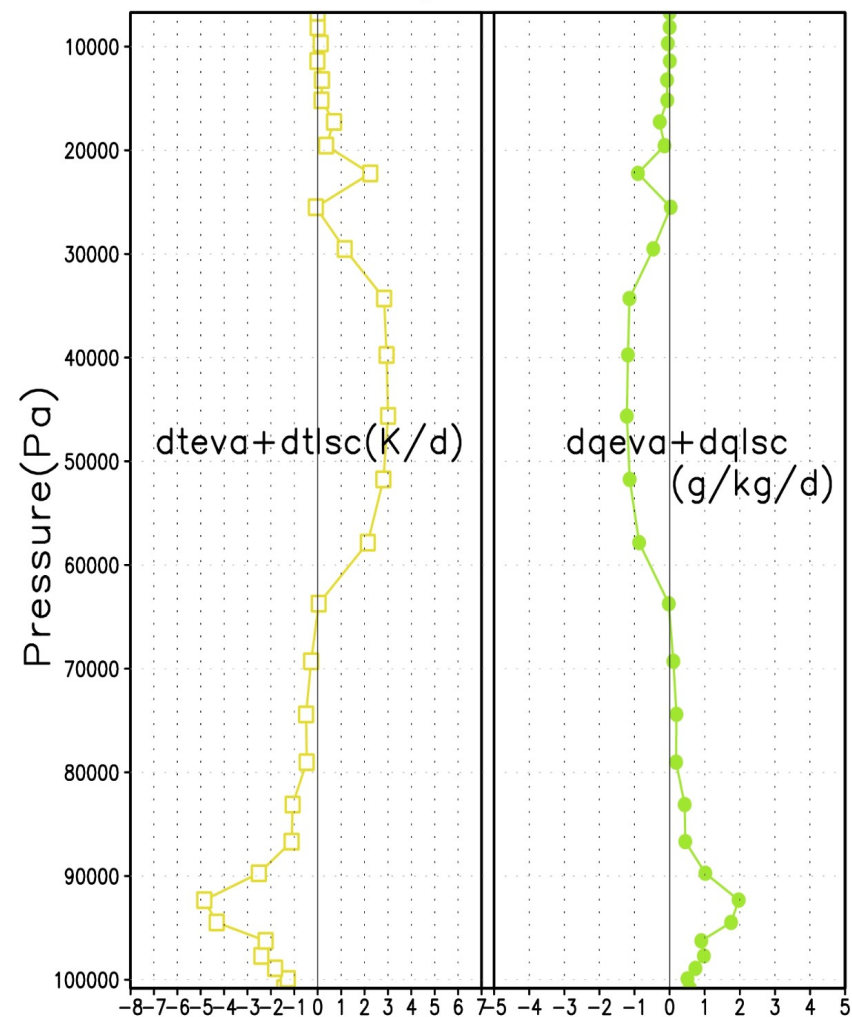
dtlsc, dqlsc : tendencies due to cloud water condensation

Total tendencies are the sums of the evaporation and condensation tendencies.

Other variables

- plul : so called "large scale" or "stratiform" precipitation; encompasses both stratiform precipitation and boundary layer cumulus precipitation.
- rneb : cloud cover
- pr_lsc_l : vertical profile of large scale liquid precipitation
- pr_lsc_i : vertical profile of large scale ice precipitation

TWPlce average



Radiation I

Subroutine : radlsw

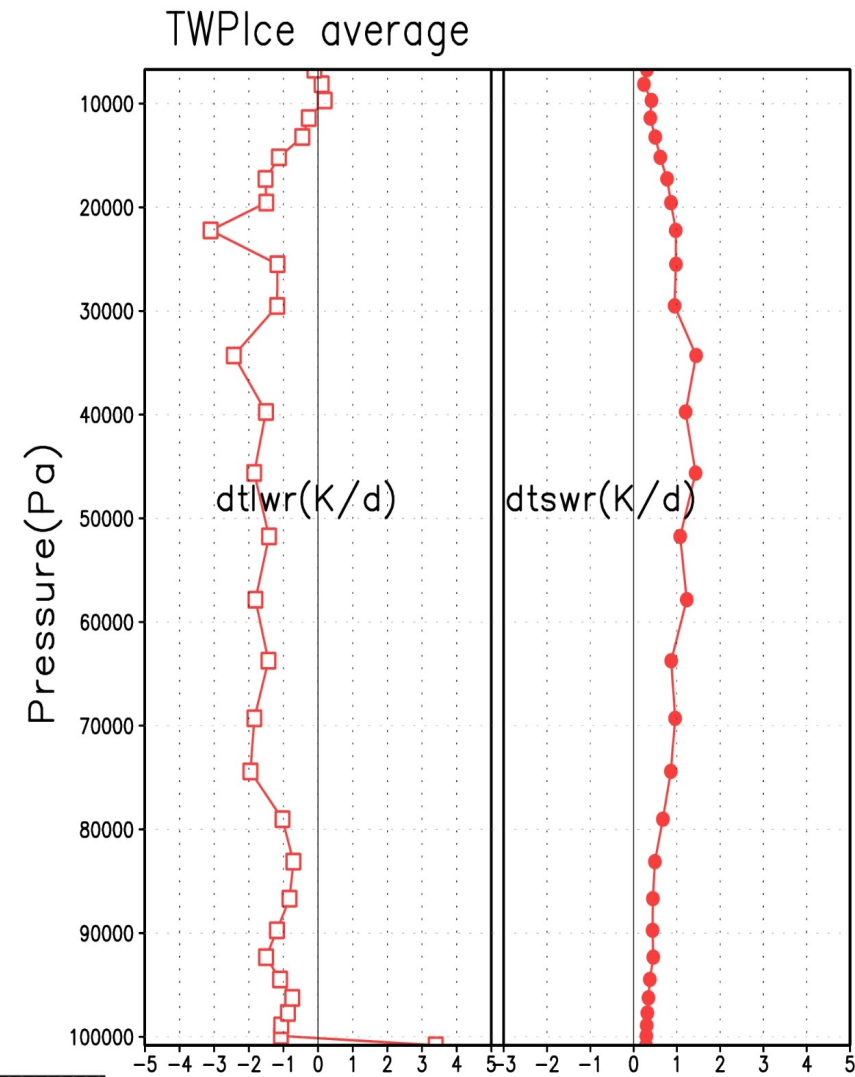
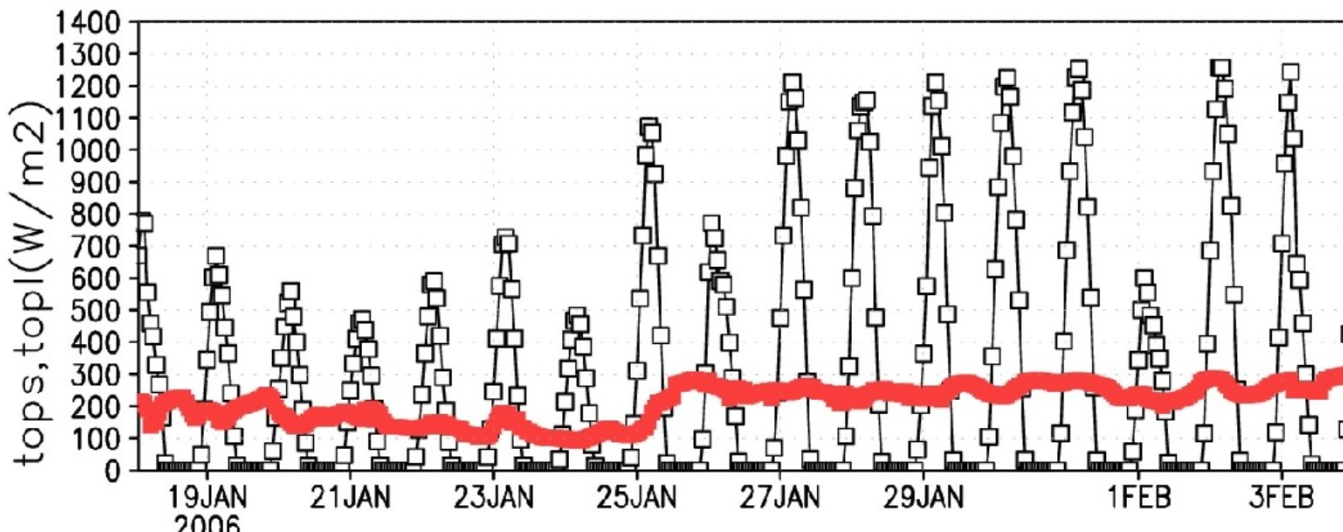
Tendencies :

dtswr, dtlwr Temperature tendencies due to solar radiation (SW = short wave) and thermal infra-red (LW = long wave)

The total radiative tendency is the sum of the SW and LW tendencies.

Other variables

- dtsw0 : clear sky SW tendency
- dtlw0 : clear sky LW tendency
- tops : net solar radiation at top of atmosphere (positive downward)
- topl : net infra-red radiation at top of atmosphere (positive upward)
- tops0, topl0 : same for clear sky
- sols : net solar radiation at surface (positive downward)
- soll : net infra-red radiation at surface (positive downward)
- sols0, soll0 : same for clear sky



Radiation II : Energy budget

Energy budget at the top of the atmosphere :

$$\text{nettop} = \text{tops-topl} = (\text{SWdn}-\text{SWup}) - (\text{LWup}-\text{LWdn})$$

Energy input (received solar energy minus reflected solar and emitted LW energy)

Positive in the tropics, negative at the poles

Surface energy budget (from the atmosphere to the surface) :

$$\text{bils} = \text{soll} + \text{sols} + \text{sens} + \text{flat}$$

$$\text{soll} = \text{ldnsfc}-\text{lwpsfc} \text{ (same for } \text{sols} \text{)}$$

flat : latent heat flux (from the atmosphere to the surface)

Negative when there is surface evaporation

sens : sensible heat flux (from the atmosphere to the surface)

Positive when the atmosphere heats the surface (polar regions)

Negative when the atmosphere is heated by the surface (continents & oceans)

In the model, this would be (- sens)

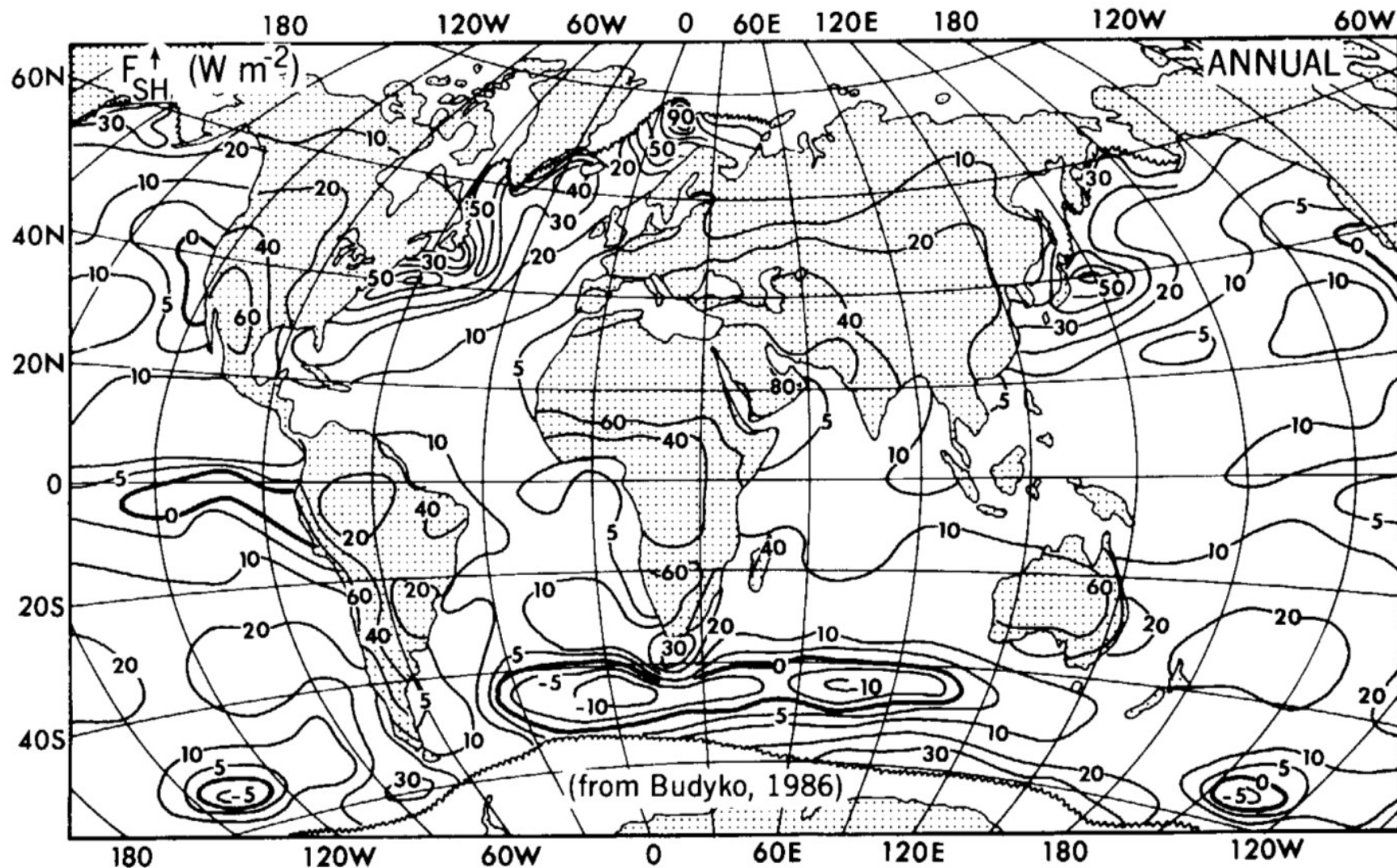


FIGURE 10.8. Global distribution of the sensible heat flux from the earth's surface into the atmosphere in W m^{-2} for annual-mean conditions after Budyko (1986).

hines_gwd

⇒ Parametrization of the momentum flux deposition due to a broad band spectrum of gravity waves.

Sources d'ondes de gravité: Convective, fronts, relief.

Wave mean flow interaction equations:

$$\frac{\partial \bar{u}_g}{\partial t} - \bar{f}_0 \bar{v} = -\frac{1}{\rho_0} \frac{\partial}{\partial y} (\rho_0 \overline{u'_g v'_g})$$

$$\frac{\partial \bar{T}}{\partial t} + N^2 \frac{H}{R} \bar{w} = -\frac{1}{\rho_0} \frac{\partial}{\partial y} (\rho_0 \overline{v'_g T'}) + \frac{\bar{J}}{C_p}$$

Transformed Eulerian mean Equations:

$$\frac{\partial \bar{u}_g}{\partial t} - \bar{f}_0 \bar{v}^* = \frac{1}{\rho_0} \vec{\nabla} \cdot \vec{F} + \bar{X}$$

$$\frac{\partial \bar{T}}{\partial t} + N^2 \frac{H}{R} \bar{w}^* = \frac{\bar{J}}{C_p}$$

Avec (\bar{v}^*, \bar{w}^*) : "residual mean circulation"

$$\bar{v}^* = \bar{v} - \frac{1}{\rho_0} \frac{R}{H} \frac{\partial}{\partial z} (\rho_0 \frac{\overline{v'_g T'}}{N^2})$$

$$\bar{w}^* = \bar{w} + \frac{1}{\rho_0} \frac{R}{H} \frac{\partial}{\partial y} (\rho_0 \frac{\overline{v'_g T'}}{N^2})$$

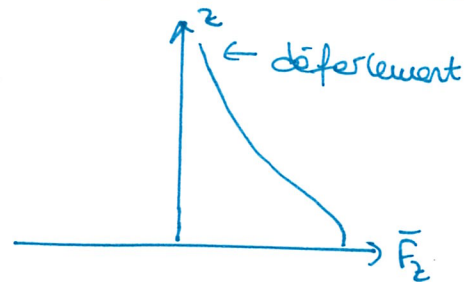
\vec{F} flux d'Eliassen-Palm $\vec{F} = \vec{C}_g \cdot A$

Pour Ondes de gravité:-

- Niveau critique de déferlement: $|\hat{\omega}| = |\omega - k u| \rightarrow 0$
 $\hat{\omega}$ fréquence intrinsèque
- Sign $(\bar{F}_z) = -\text{sign}(\hat{\omega})$

$$\boxed{\begin{matrix} \hat{\omega} < 0 \\ R > 0 \end{matrix}} \left\{ \begin{matrix} \hat{C}_\varphi < 0 \\ \bar{F}_z > 0 \end{matrix} \right.$$

Propagation de la Phase vers l'Ouest

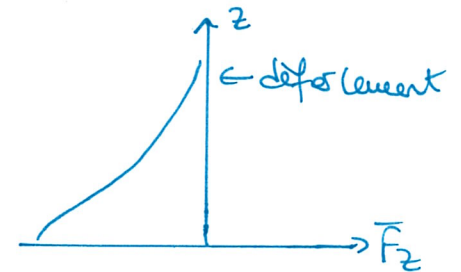


$$\frac{d\bar{F}_z}{dz} < 0 \Rightarrow \frac{\partial \bar{u}}{\partial t} < 0$$

Freine le vent moyen

$$\boxed{\begin{matrix} \hat{\omega} > 0 \\ R > 0 \end{matrix}} \left\{ \begin{matrix} \hat{C}_\varphi > 0 \\ \bar{F}_z < 0 \end{matrix} \right.$$

Propagation de la Phase vers l'Est



$$\frac{d\bar{F}_z}{dz} > 0 \Rightarrow \frac{\partial \bar{u}}{\partial t} > 0$$

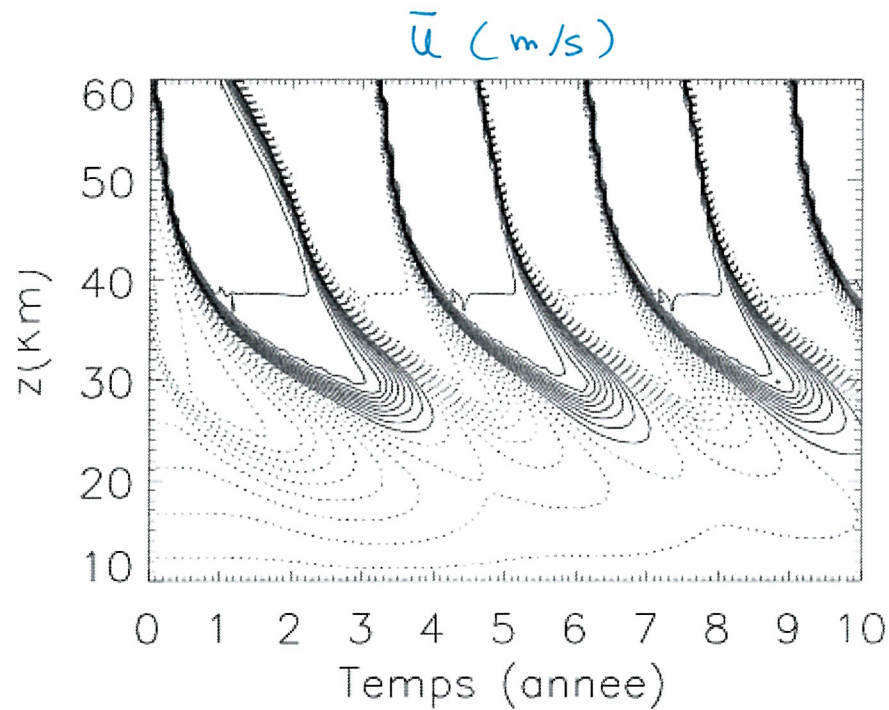
Accélère le vent moyen

=> Quasi-Biennial Oscillation

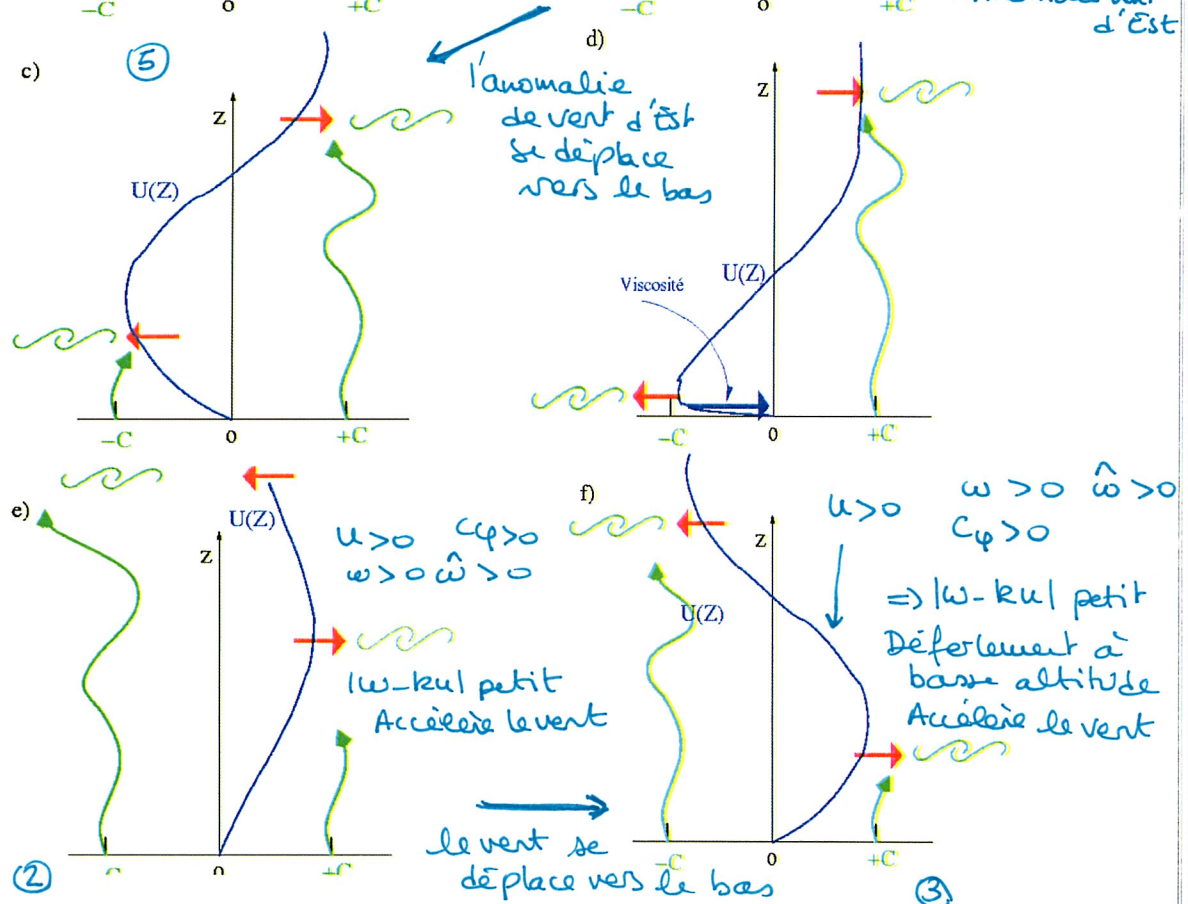
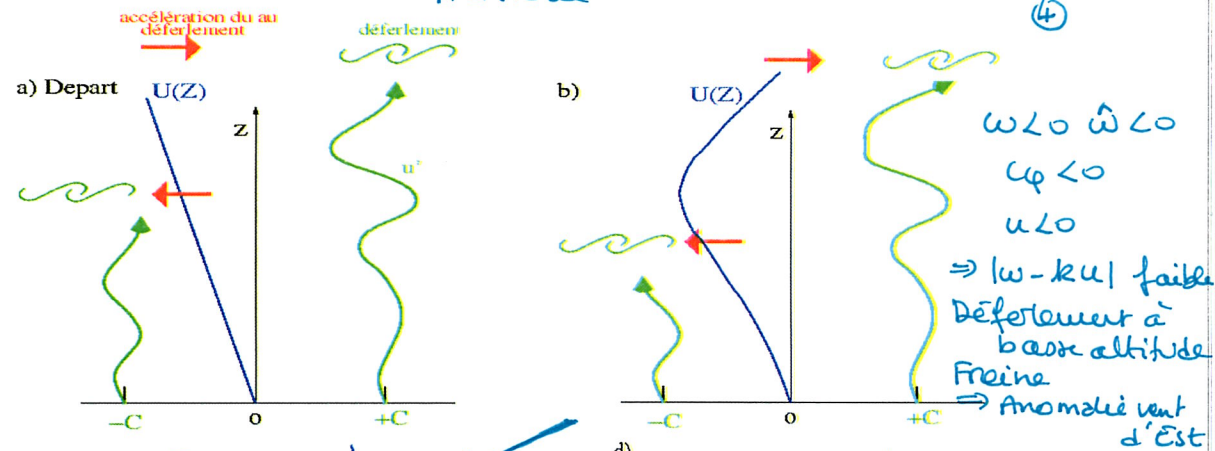
Altitude de déferlement des ondes de gravité:-

$$z = 2H \ln \left(\frac{|\omega - ku|}{|m| W_0} \right)$$

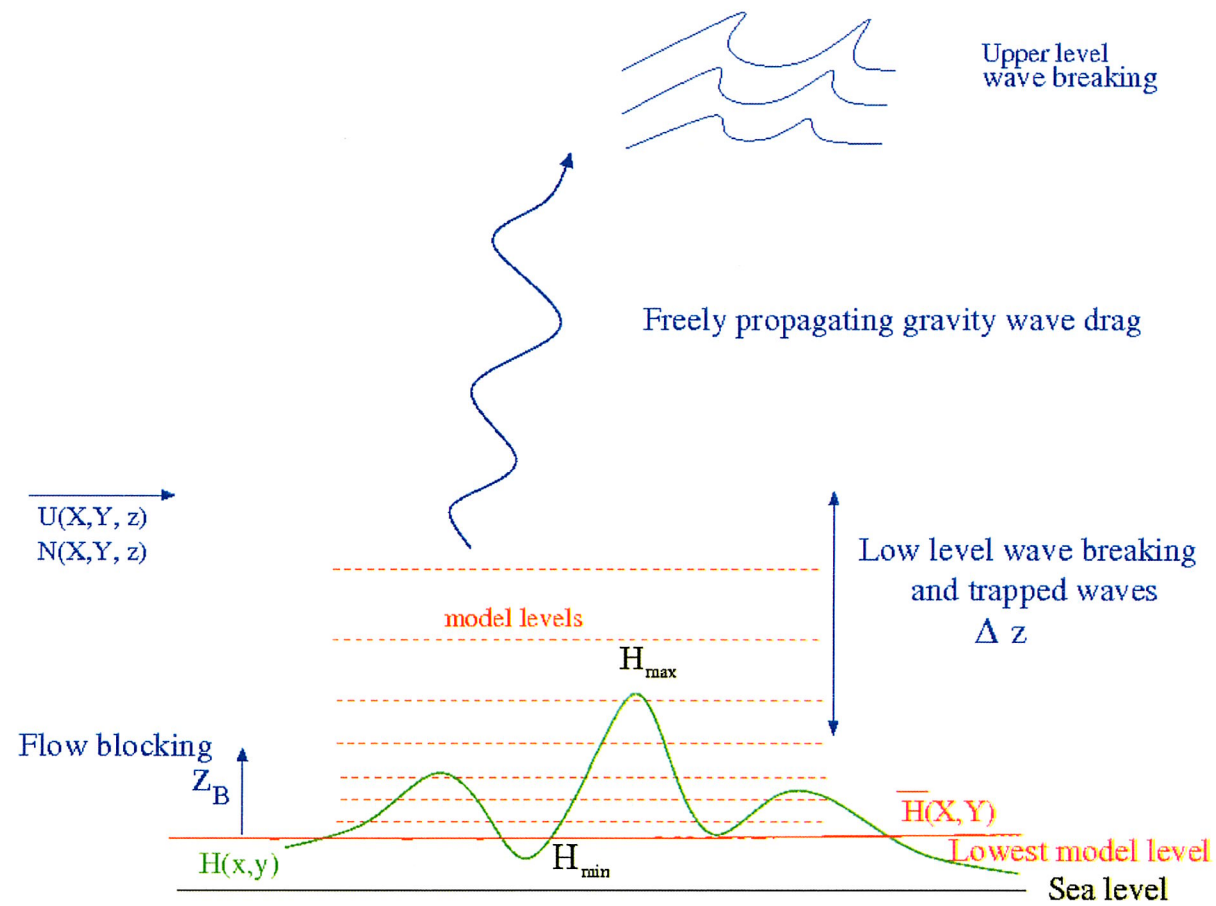
- $\hat{\omega} > 0$ Accélère le vent moyen
- $\hat{\omega} < 0$ Freine le vent moyen



① $\omega > 0$ $\hat{\omega} > 0$ $c_p > 0$
 $u < 0 \Rightarrow |u - ku|$ grand
 Pénétration de l'onde à haute altitude



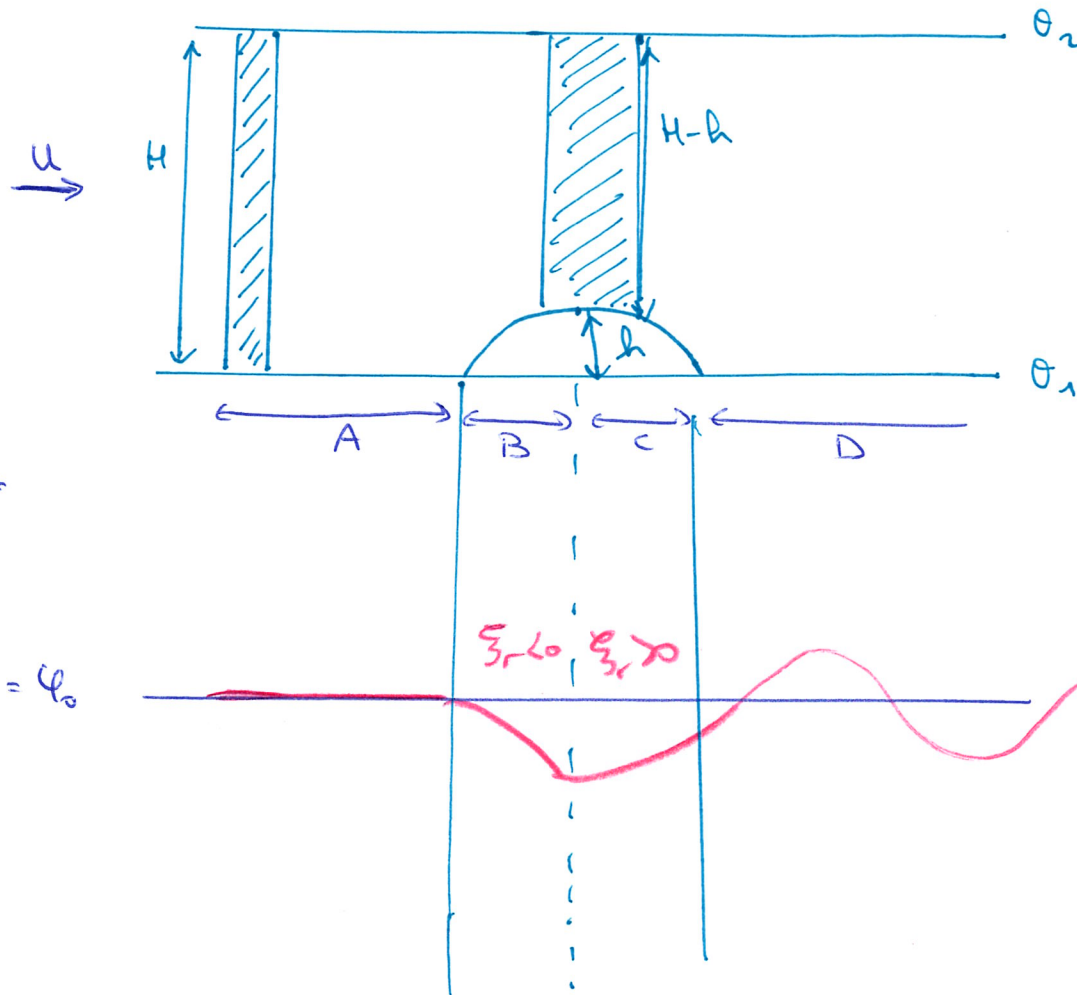
drag_noro



lift_noro

Au cours d'un mouvement adiabatique
On conserve la vorticité potentielle :
 $PV = \frac{1}{\sigma} \oint \partial \theta$

PV = $\frac{1}{e} \sum_a \frac{\partial \theta}{\partial z}$ avec $\sum_a = f + \sum_r$



Vue de haut :

$$\psi = \psi_0$$

$$PV = \frac{f_0}{e} \frac{(\theta_2 - \theta_1)}{h}$$

$$\bullet \frac{E_n B}{\frac{\partial \theta}{\partial z}} \rightarrow \text{Dmc } \xi_a \downarrow$$

$$\Rightarrow \xi_r < 0$$

Déviatiôn vers le Sud ($v < 0$)

• $\frac{\partial \epsilon_c}{\partial z} =$ \rightarrow Donc $\xi_a \nearrow \xi_r \searrow$

Déviation vers le Nord

- En D = $\frac{\partial \theta}{\partial z}$ retrouve sa valeur initiale $\theta_a = f_0$

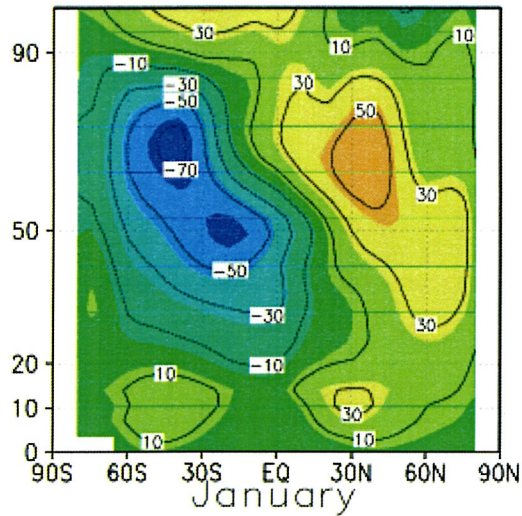
Mais quand la colonne atteint la latitude ϕ_0 elle vient du sud \Rightarrow trajectoire oblique.

La colonne traverse le Parallèle.
Au Nord $f = f_0 + \beta y$ $\Sigma r = -\beta y$
 \Rightarrow on retourne vers le Sud.

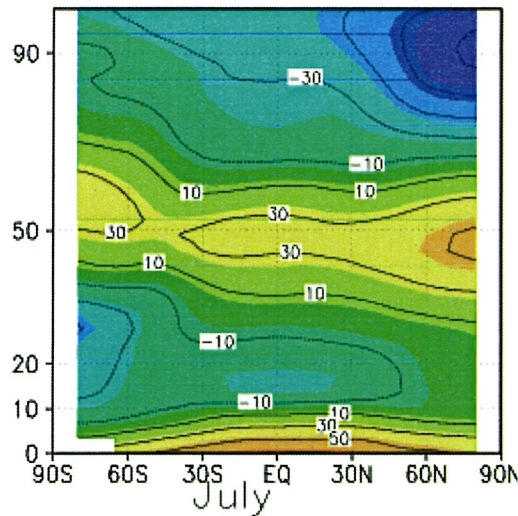
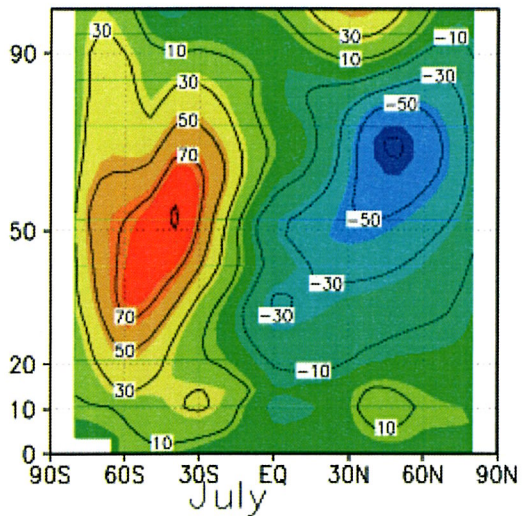
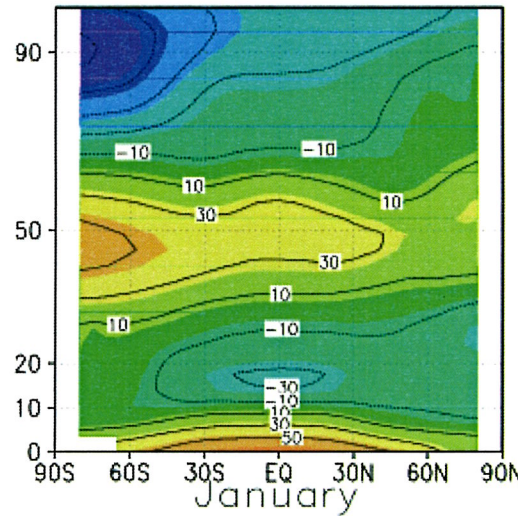
Etc.

⇒ Importance des ondes de Rossby stationnaires créées par le relief pour la circulation stratosphérique.

U (m/s)

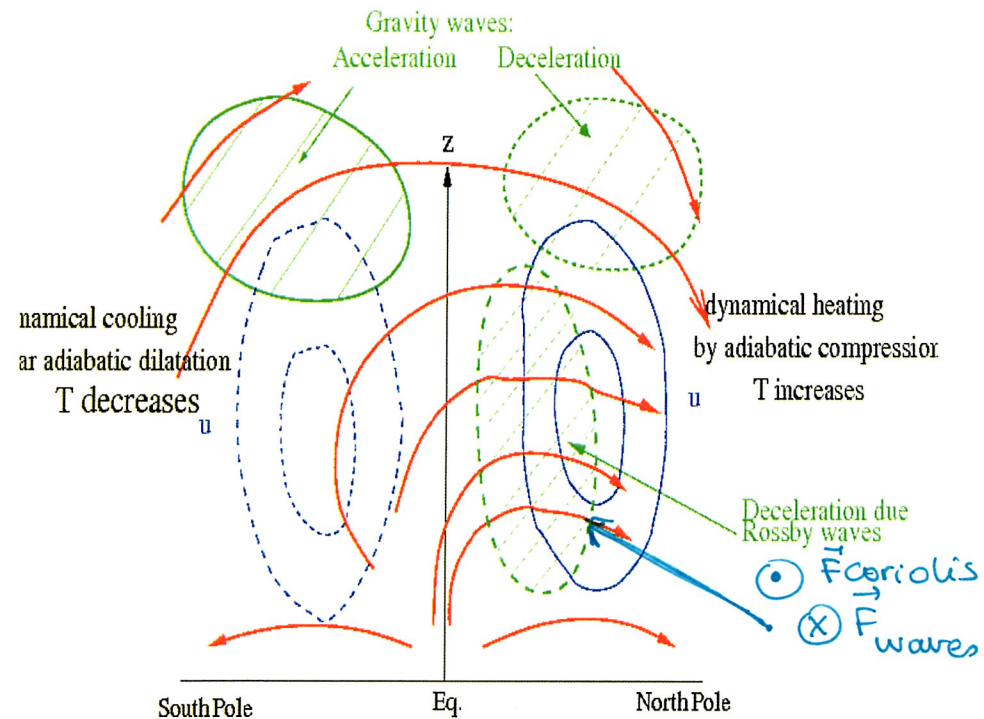


Température (°C)



- A 50 km max de Température au Pôle d'Été
- En Janvier $U > 0$ de l'Hém Nord
 $U < 0$ de l'Hém Sud
- Le gradient de Température n'est pas aussi fort que s'il était déterminé radiativement uniquement.

En Janvier :



\Rightarrow Diminution du gradient horizontal de Température obtenu par les termes radiatifs.

Relation de dispersion des ondes de Rossby :

$$c - u_0 = \frac{-\beta}{k^2 + l^2 + \frac{f_0^2}{N^2} \left(m^2 + \frac{1}{4H^2} \right)}$$

ondes stationnaires: $c = 0 \Rightarrow u_0 > 0$

$$m^2 = \frac{N^2}{f_0^2} \left[\frac{\beta}{u_0} - (k^2 + l^2) \right] - \frac{1}{4H^2}$$

Propagation verticale des ondes de Rossby pour $m^2 > 0$

$$\Rightarrow \boxed{0 < u_0 < u_c}$$

$$\begin{array}{l} z \\ \uparrow \\ u = u_c \quad \vec{F} = 0 \\ 0 < u < u_c \quad \uparrow \vec{F} = \vec{G} A \end{array} \left. \vphantom{\begin{array}{l} z \\ \uparrow \\ u = u_c \quad \vec{F} = 0 \\ 0 < u < u_c \quad \uparrow \vec{F} = \vec{G} A \end{array}} \right\} \frac{\partial F_z}{\partial z} < 0$$

En Janvier de l'IN

$u > 0 \Rightarrow$ propagation verticale jusqu'à z tel $u = u_c$

TEK équation: $\frac{\partial \bar{u}}{\partial t} - \frac{1}{f_0} \bar{v}^* = \frac{1}{\rho} \vec{\nabla} \cdot \vec{F}$

$\frac{\partial F_z}{\partial z} < 0 \Rightarrow$ freinage

Stationnaire: $-\frac{1}{f_0} \bar{v}^* = \frac{1}{\rho} \vec{\nabla} \cdot \vec{F}$

Orography

Subroutines : drag_noro (or drag_noro_strato)
& lift_noro (or lift_noro_strato)

Tendencies :

dtoro, duoro, dvoro : tendencies of temperature and velocity due to the drag
dtlif, dulif, dvlif : tendencies of temperature and velocity due to the lift

Total tendencies are the sums of the drag and lift tendencies.