

Projet CEFIPRA « NFIM »
(F. Hourdin / O.P. Sharma)

ANR « HEAT »

Projet IPSL « DYNAMICO »

Projet G8 « ICOMEX »
(Y. Meurdesoif)

2011

2012

2013

2014

2015

2016

2017

Thèse Sarvesh Dubey (LMD & IIT Delhi)

Post-doc L. Fita

Thèse Marine Tort (LMD)

Post-doc A. Traore

Post-doc E. Kritsikis

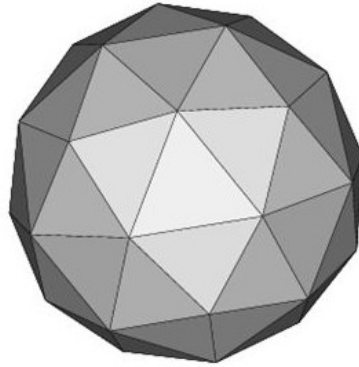
Thèse C. Colavolpe (CNRM)

Mesh partitioning for parallel computing

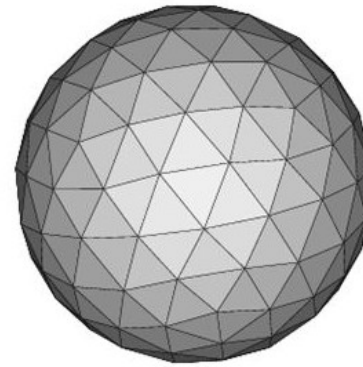
nbp=1



nbp=2



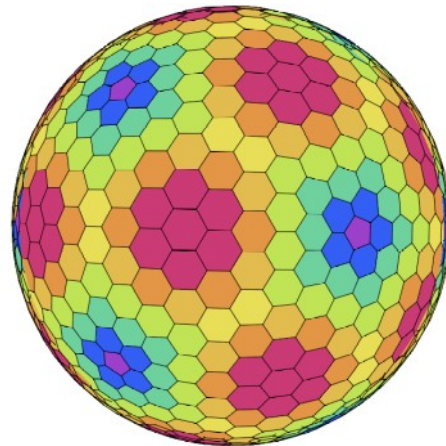
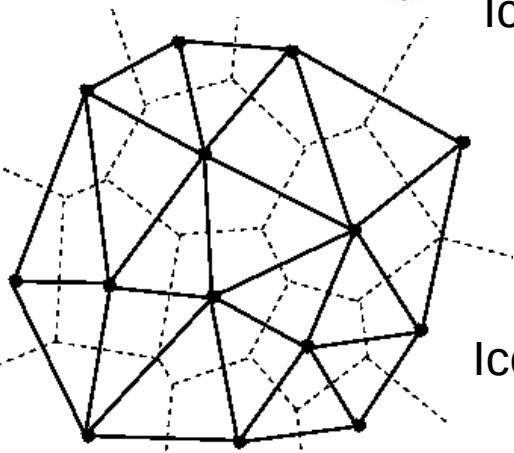
nbp=4



Icosahedral-triangular mesh
 $10 \cdot nbp^2 + 2$ vertices

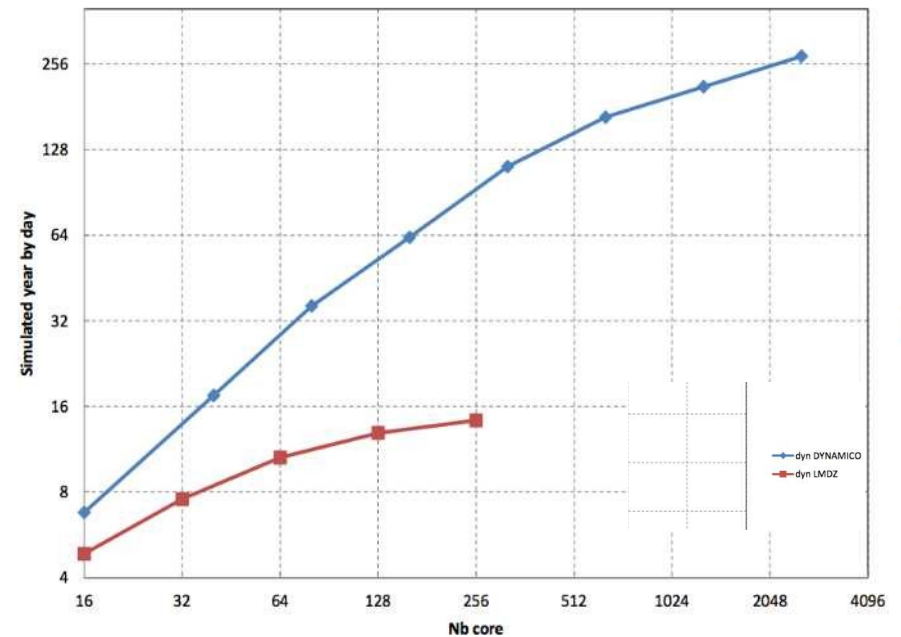
Voronoi
dual

Icosahedral-hexagonal mesh
 $10 \cdot nbp^2 + 2$ cells



- Easy to partition into $10 \times nsplit^2$ domains
- About $(nbp/nsplit)^2$ cells per domain = MPI process
- $Nbp/nsplit > 10$ for performance

Dynamico : 32x32x10x39lvl Vs LMDZ 96x95x39



Y. Meurdesoif, 2013

Physical space

$$\lambda, \varphi, \Phi, g$$

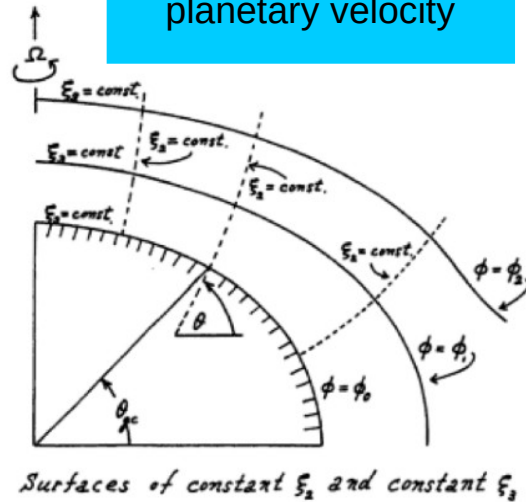
$$u, v, w$$

$$\alpha = 1/\rho, s, r$$

Thermodynamics

$$p, T, \chi$$

Geopotential
Coordinates
Metric, gravity,
planetary velocity



Computational space $S^2 \times [0,1]$

$$\mathbf{v} = \mathbf{G}(\mathbf{n}, \Phi) \cdot \dot{\mathbf{n}} + \mathbf{R}(\mathbf{n}, \Phi)$$

$$\Phi(\mathbf{n}, \eta, t), \mu,$$

\mathbf{n}

η

Horizontal mesh
Icosahedral C-grid

Vertical mesh
Lorenz

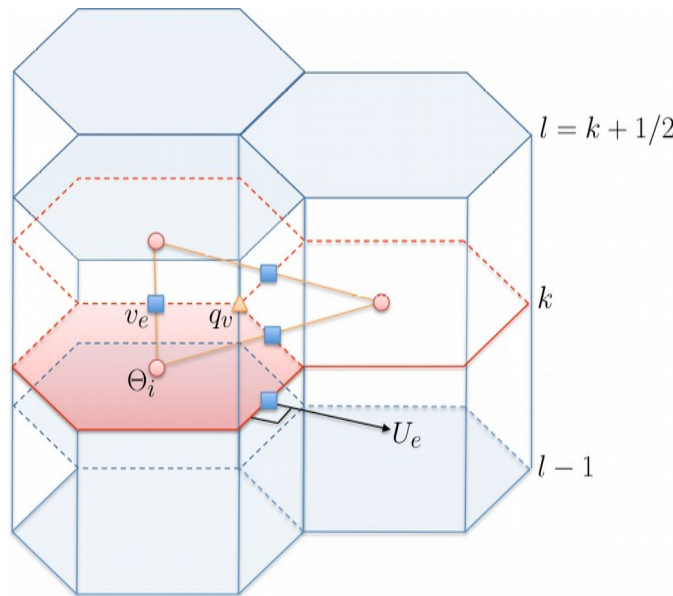
Discrete space

$$m_{ik} = \int \int \int \mu d\mathbf{n}d\eta$$

$$W_{il} = \int \int \mu \eta d\mathbf{n}$$

$$v_{ek} = \int \mathbf{v} \cdot d\mathbf{n}$$

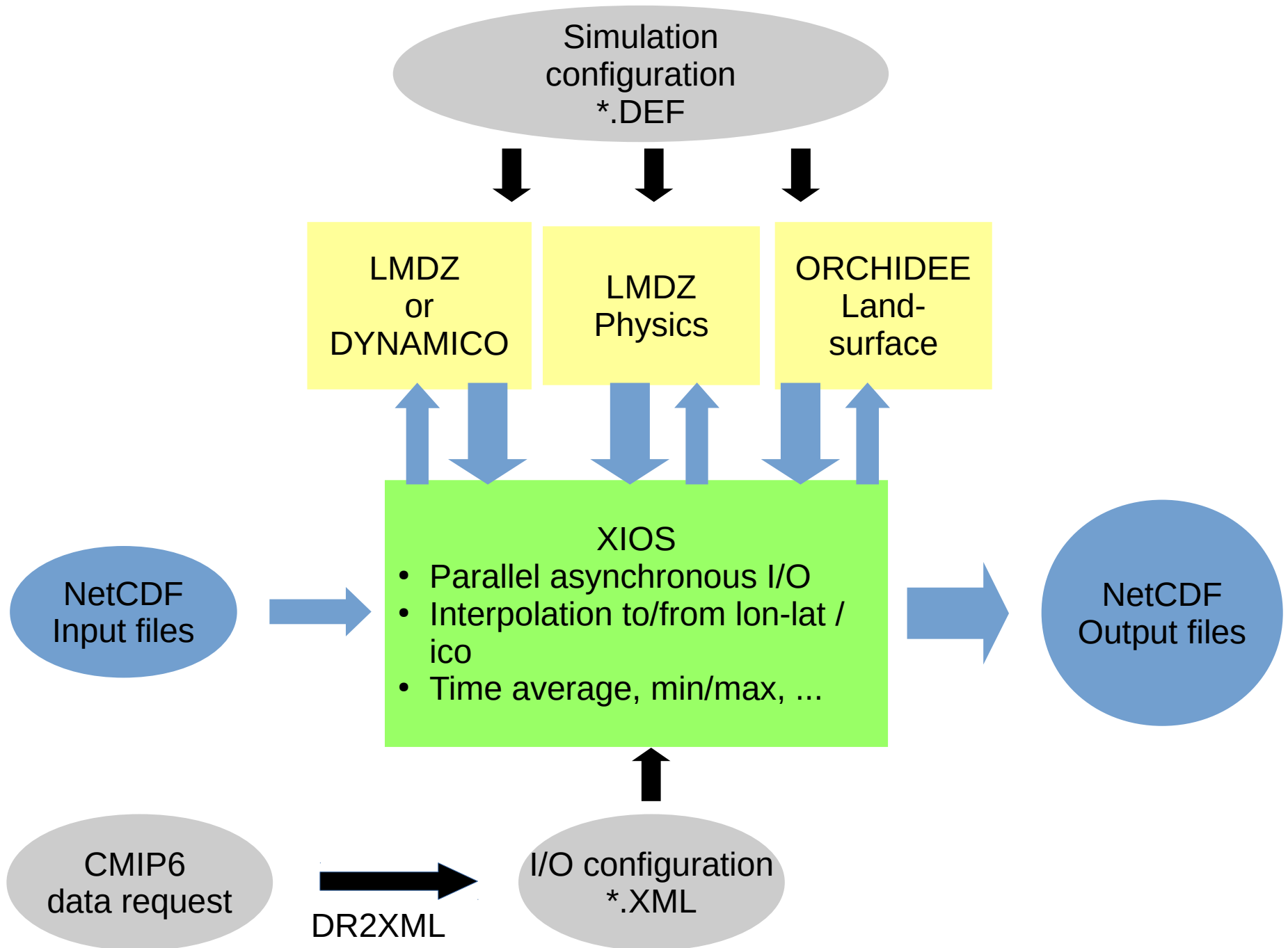
$$\alpha_{ik} = \alpha(p_{ik}, s_{ik}),$$



- Discrete integration by parts (Bonaventura & Ringler, 2005 ; Taylor, 2010)
- Energy- and vorticity- conserving Coriolis discretization (TRISK : Thuburn et al., 2009 ; Ringler et al., 2010)

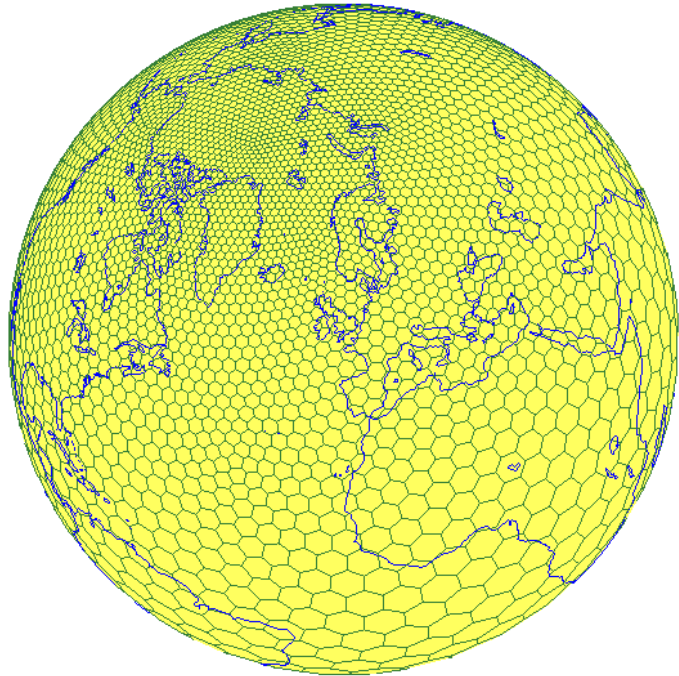
Energy-conserving
3D core

Atmosphere-only configurations

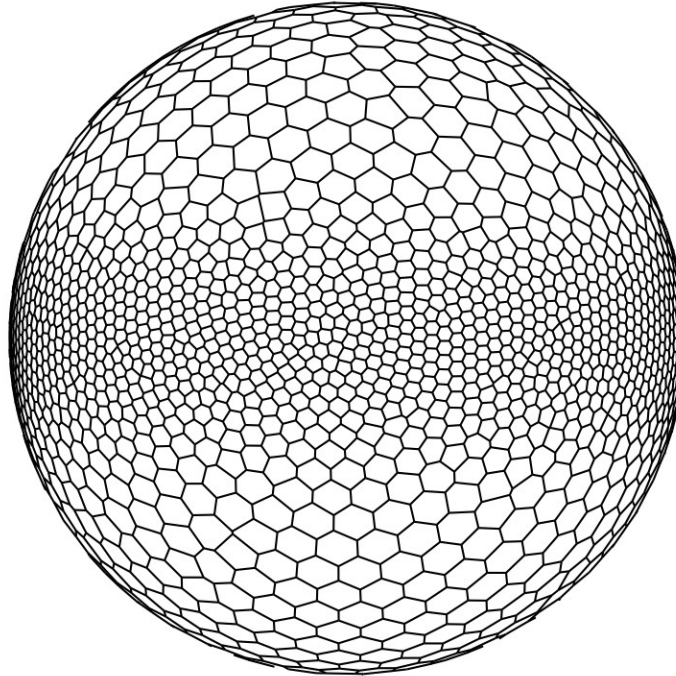


Towards regional configurations with DYNAMICO

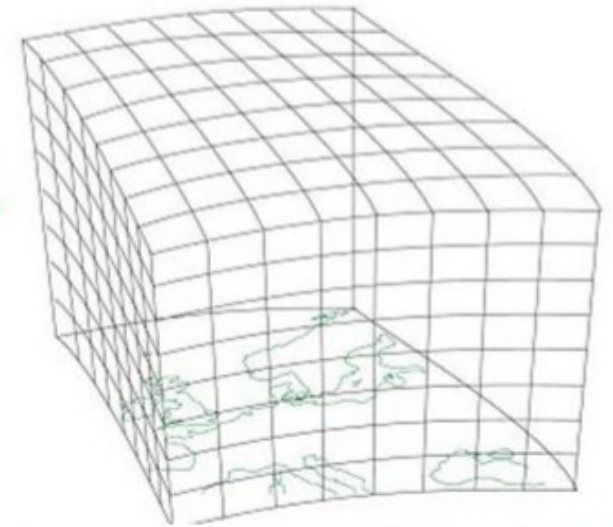
<https://lmdz.lmd.jussieu.fr/le-coin-des-developpeurs/reunions/2020-01-13>
<https://cmc.ipsl.fr/online-session-on-ipsl-cm-configurations/>



Zoom (Schmidt transform)



Zoom (unstructured mesh)



Limited area (LAM)

Status : many pieces of the puzzle there, nothing operational yet

Work could start now with the Schmidt transform (some tuning of dissipation needed)

AWACA science-driven modelling goals

Past

- Goal : go beyond simple statistical relationships (e.g. isotopic thermometer) to reconstruct past variability from isotopic records
- How : inverse modelling, e.g. Steiger (2017)
- Requires : isotope-aware model, multi-centennial simulations

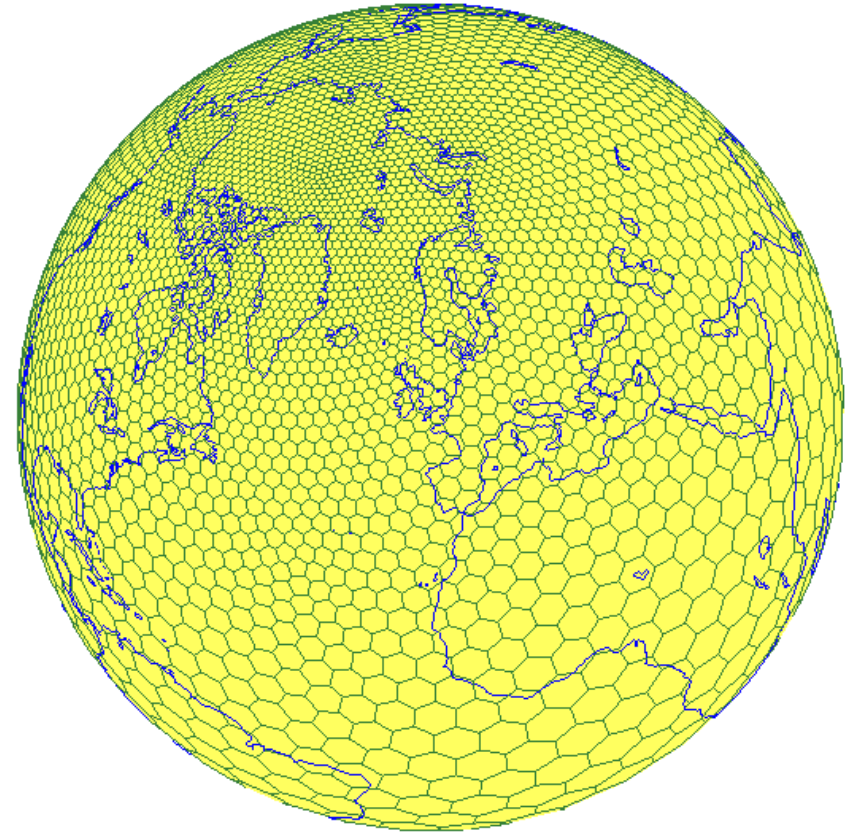
Present

- Goal : reduce biases in simulated present-day Antarctic precipitation ; characterize contribution of moist processes to uncertainty
- How : characterization, understanding and representation (parameterization) of small-scale moist processes
- Requires : process-rich, high resolution, limited-area modelling (+tailored observations ...)

Future

- Goal : disentangle contributions to future changes in Antarctic precipitation from ocean and sea-state, atmospheric circulation, small-scale processes
- Goal : deliver high-resolution projections of Antarctic precipitation, with quantified uncertainty, for use beyond AWACA, especially for ice sheet modelling
- How : ensembles of simulations 1800-2100
- Requires : large multi-centennial ensembles, integration with CMIP-class ESM

A modelling portfolio for AWACA



- LMDZ the atmospheric component of the IPSL-CM ESM
- MAR regional atmospheric model with particular focus on polar regions
- DYNAMICO Dynamical Core on Icosahedral mesh : a massively parallel atmospheric solver

	Existing			AWACA	
	MAR	DYNAMICO	LMDZ	MAR+ICO	LMDZ+ICO
Polar processes	√			√	√
Massively parallel		√		√	√
LAM	√	(√)		√	√
Isotope-aware	√	(√)	√	√	√
CMIP-capable		√	√		√

